

Observing ionized stellar outflows with LOFAR

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POLFAR



LOFAR

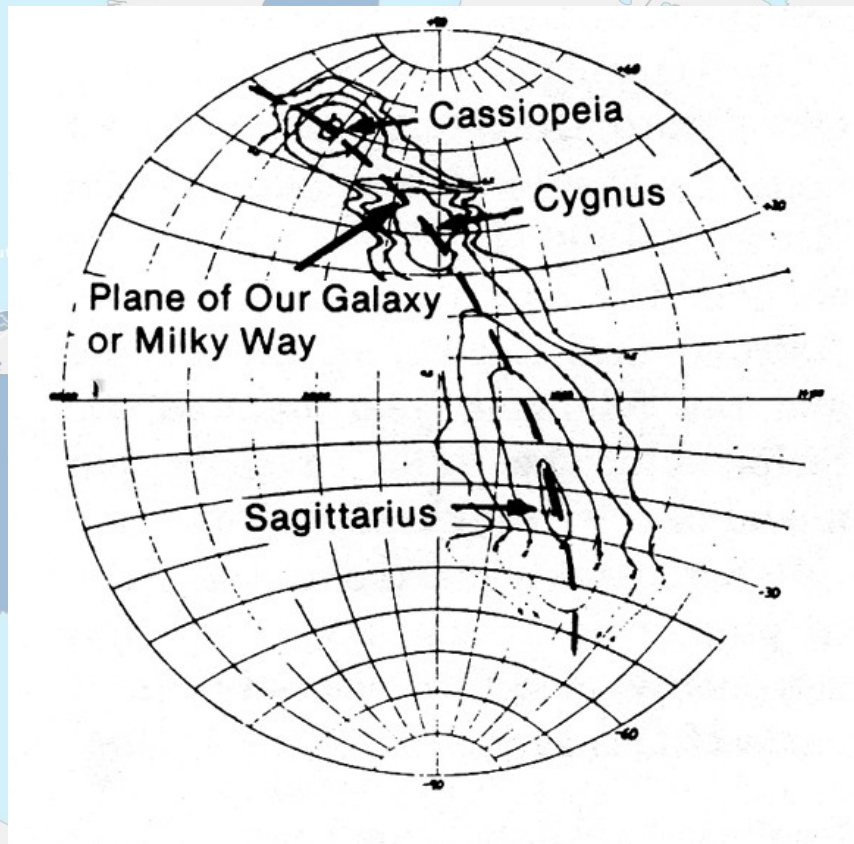


NARODOWA AGENCJA
WYMIANY AKADEMICKIEJ

Stellar outflows observed in radio

Two-metre survey by Grote
Rober

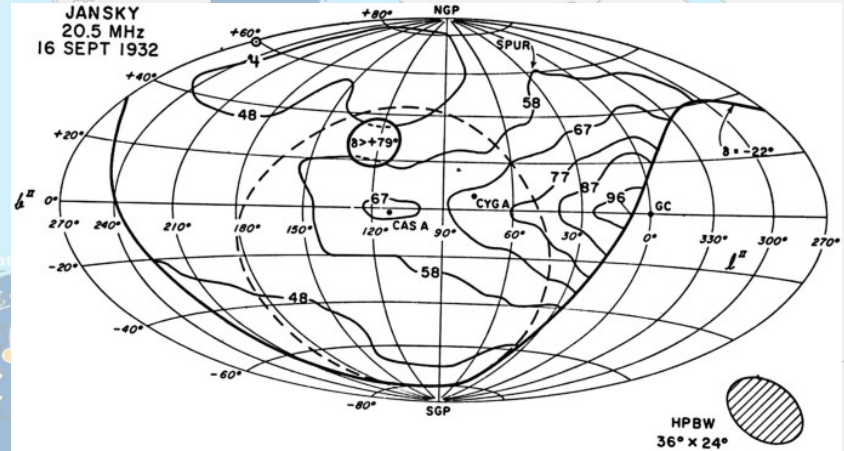
Rober G., 1944, **ApJ**, 100, 279



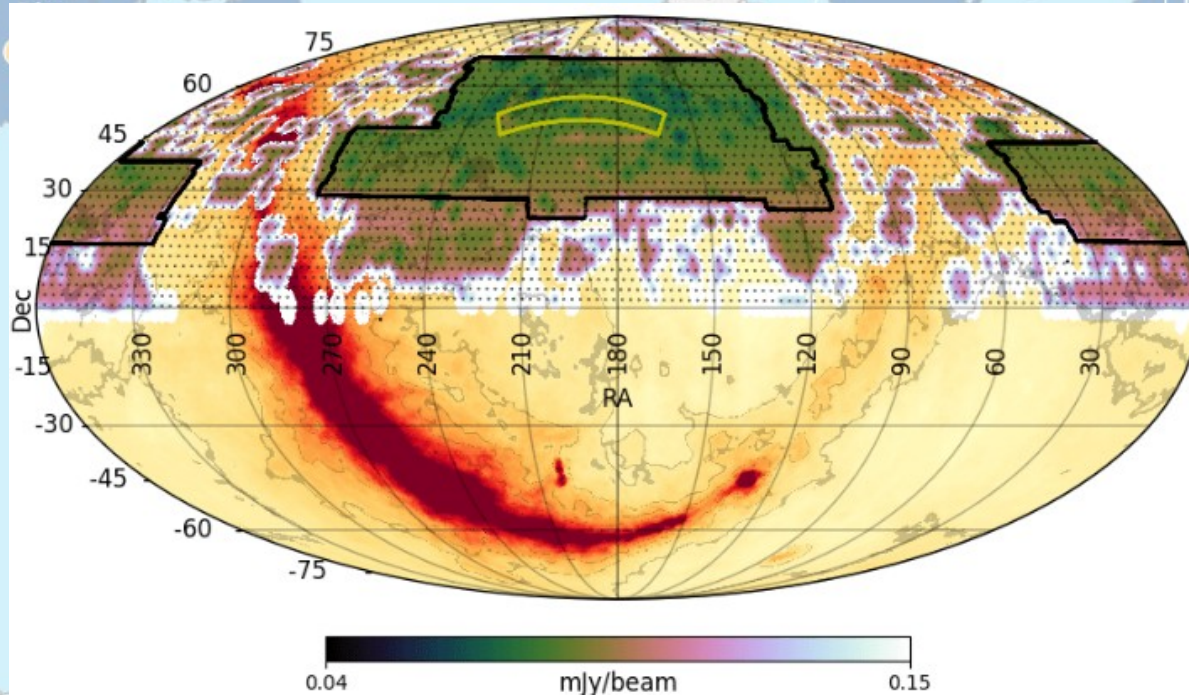
Stellar outflows observed in radio

20.5 MHz survey by Karl Jansky in 1932

Sullivan, W., 1978, **S&T**, 56, 101

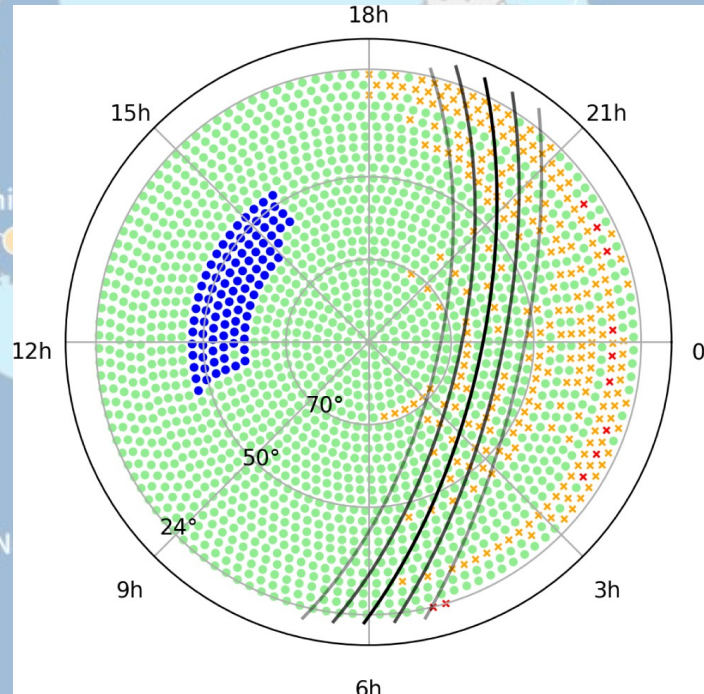


LoTSS 120-168 MHz



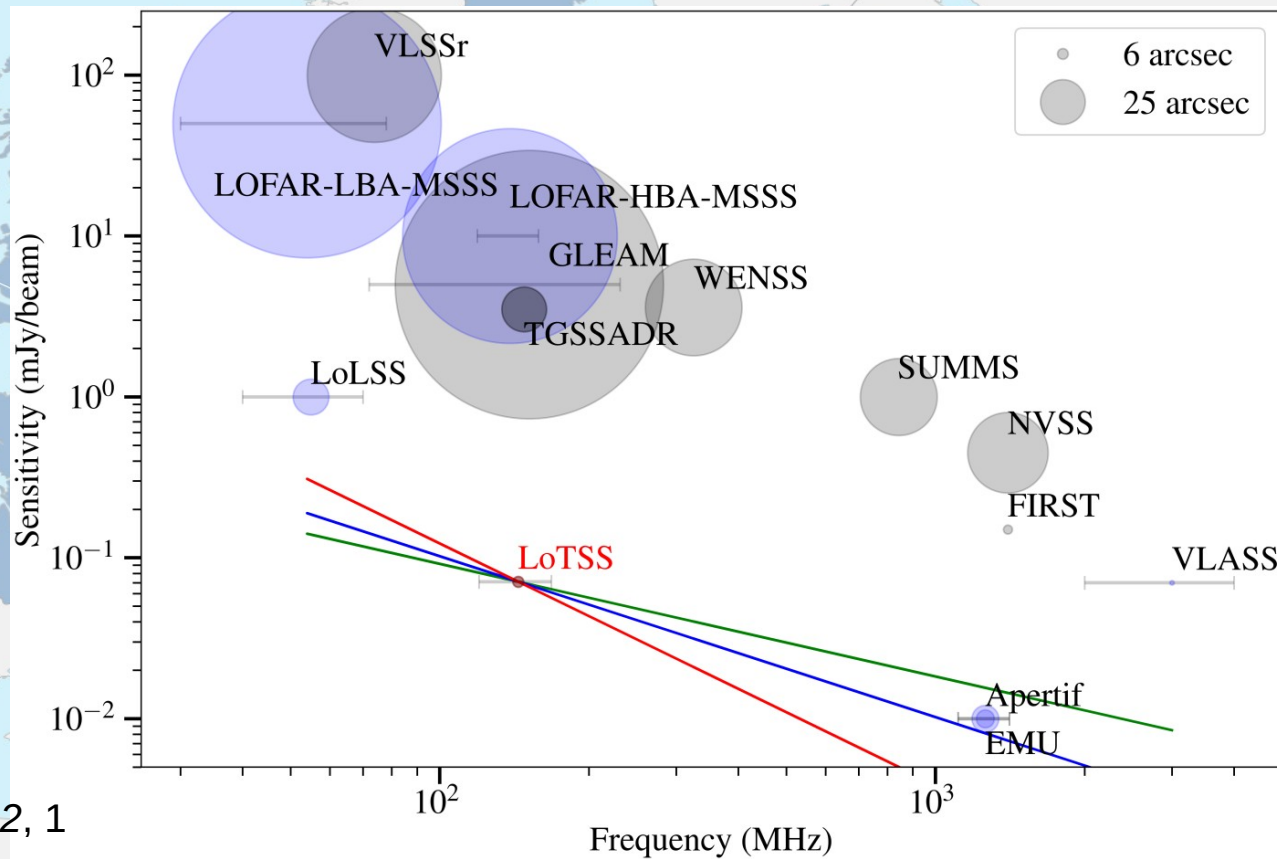
Shimwell, T., 2022, **A&A**, 659A, 1

LoLLS 41-66 MHz



de Gasperin, F., et al., 2023, **A&A**, 673A, 165

LOFAR wide-area surveys



Stellar sources in LOFAR

Kavanagh, R., et al., 2022, **MNRAS**, 514, 675

Vedantham, H. K., et al., 2022, **ApJ**, 926L, 30

Hajduk, M., et al., 2022, **A&A**, 665A, 152

Pope, B. J. S., et al., 2021, **ApJ**, 919L, 10

Vedantham, H. K.; et al., 2020, **ApJ**, 903L, 33

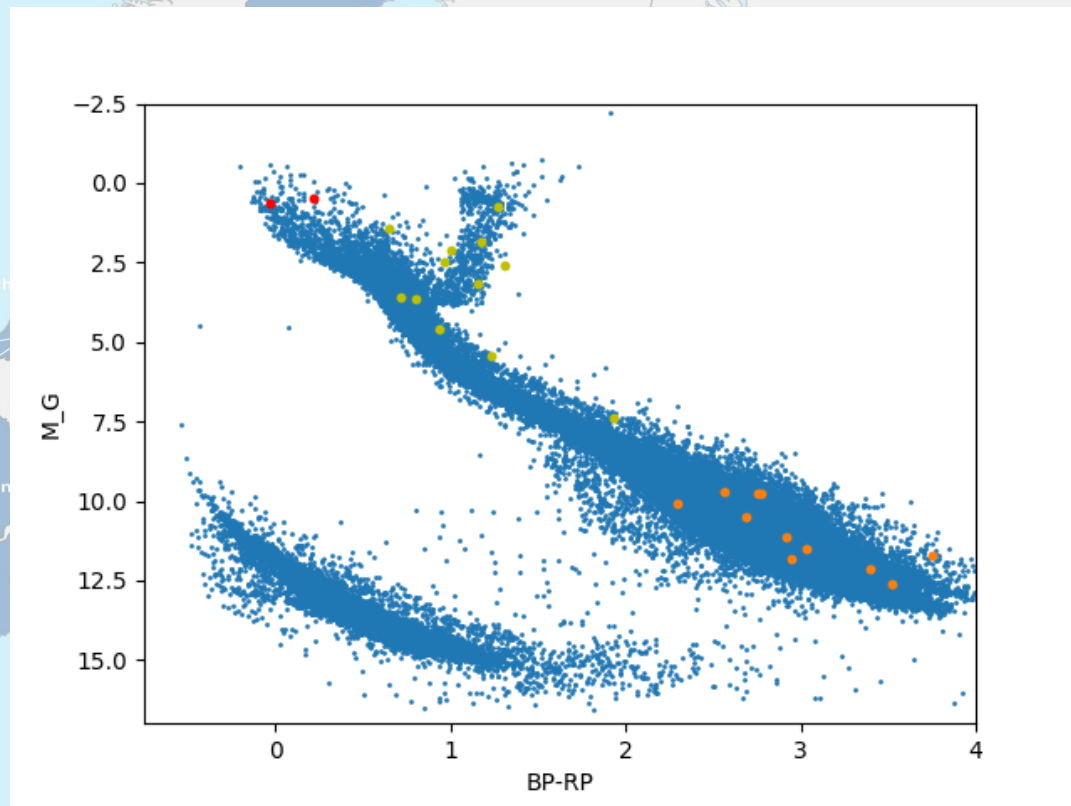
Callingham, J. R., et al., 2021, **Nature Aston.**, 5.1233

Vedantham, H. K., et al., 2020. **Nature Astron.**, 4, 577

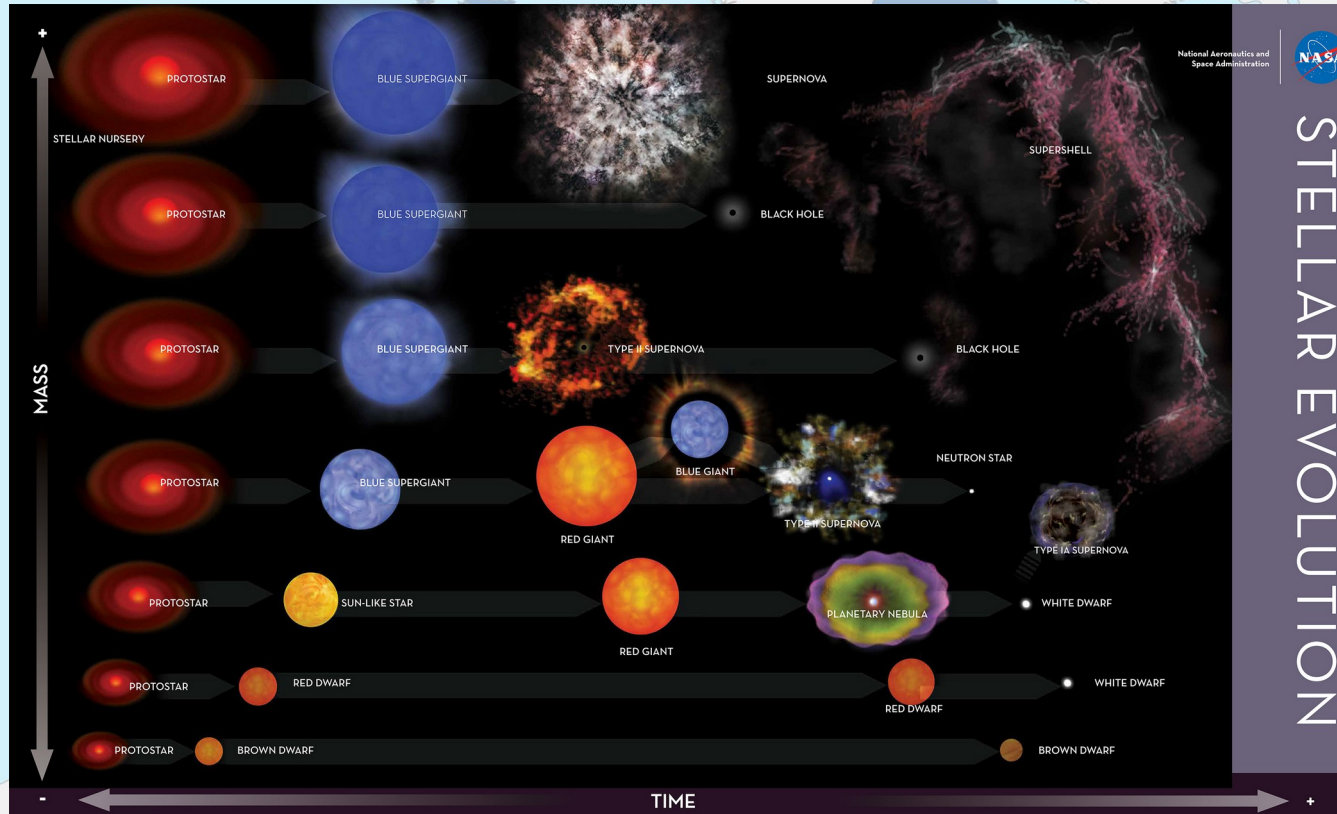
Vedantham, H. K., et al., 2020, **ApJ**, 903L, 33

Callingham, J., et al., 2021, **A&A**, 648A, 13

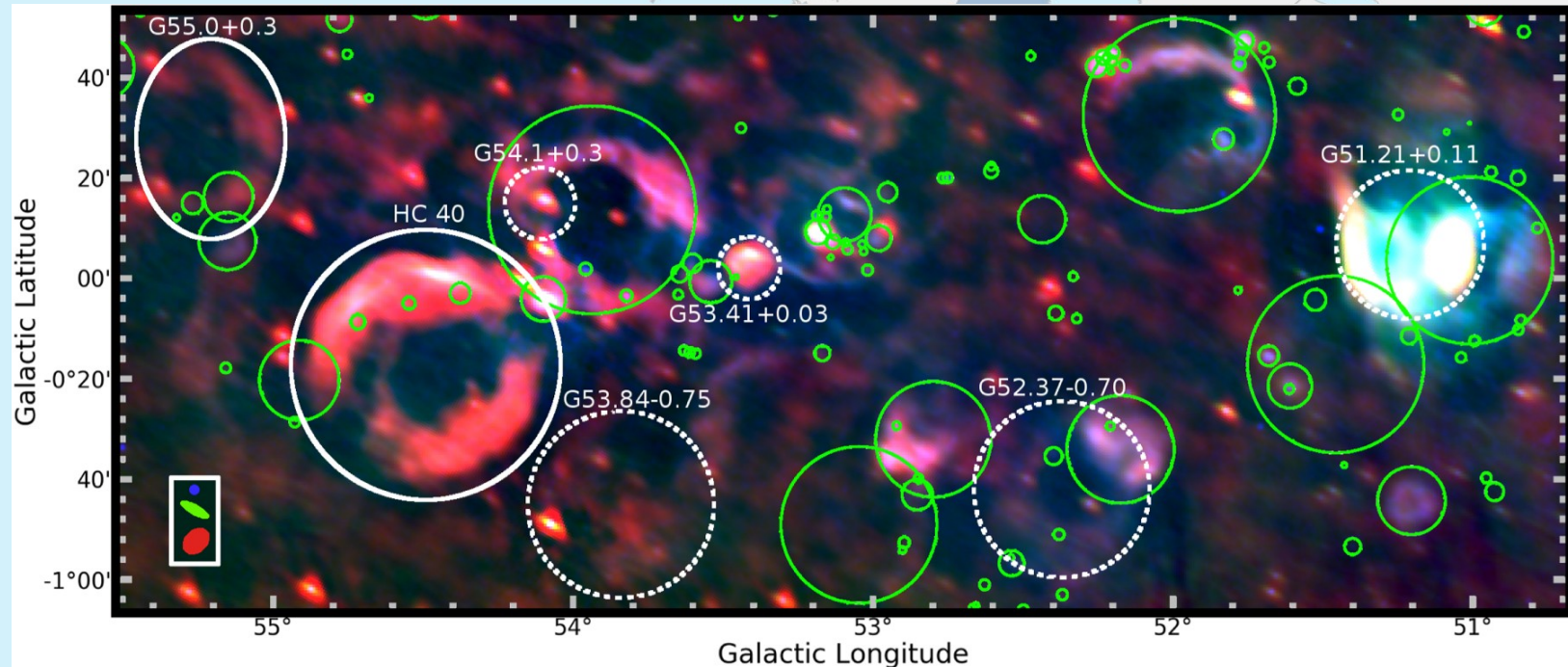
Toet, S. E. B., et al., 2021, **A&A**, 654A, 21



Ionized nebulae throughout stellar evolution

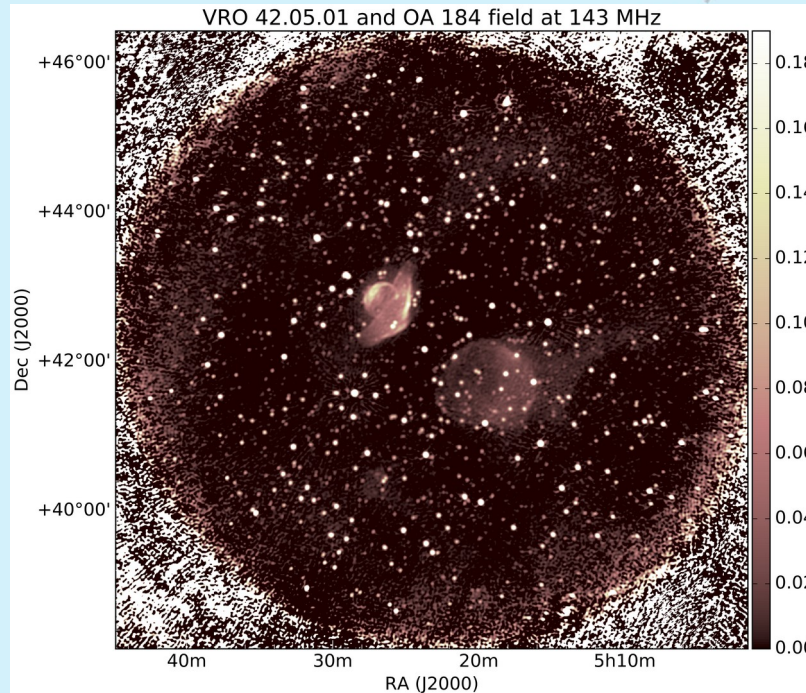


Supernova remnants

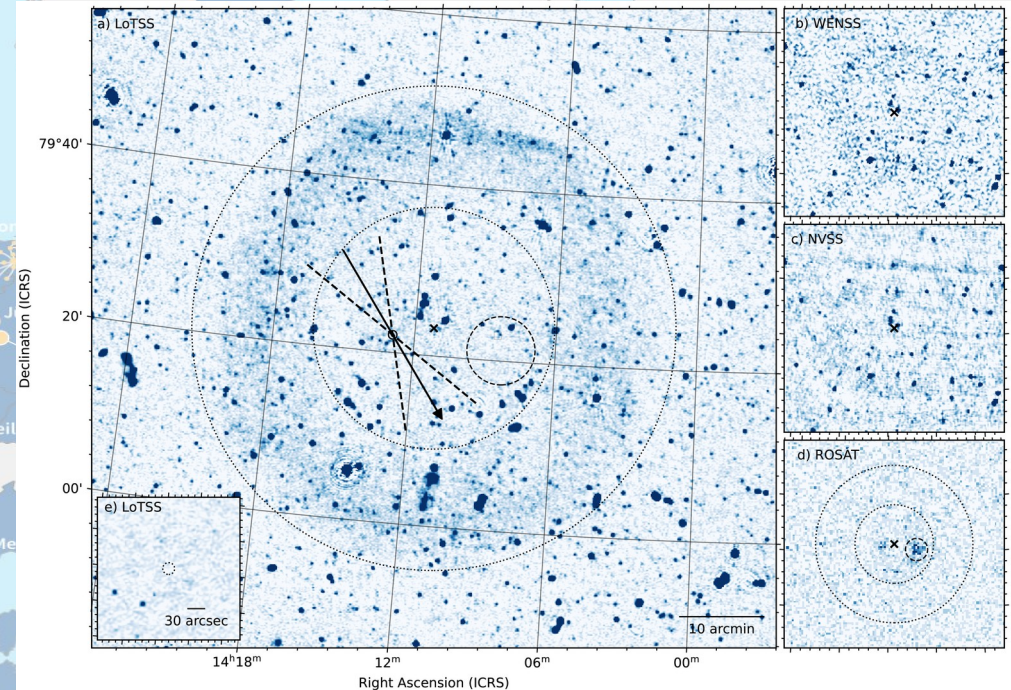


Driessen, L., et al., 2018, **ApJ**, 860, 133

Supernova remnants

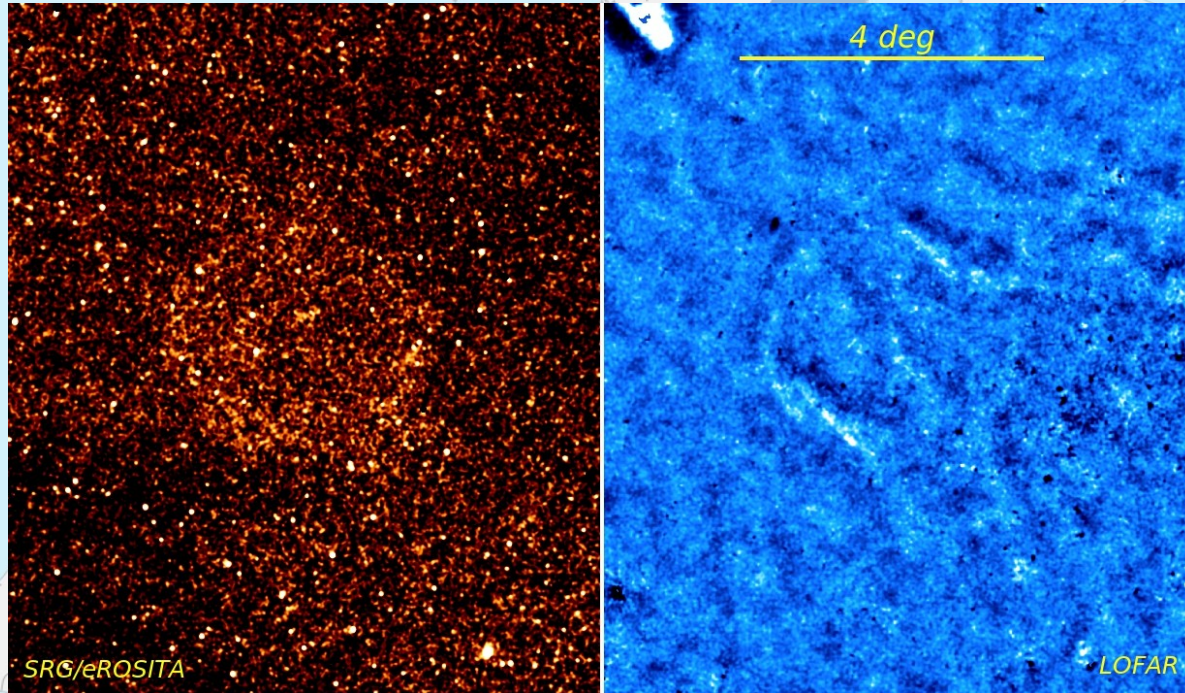


Arias, M., et al., 2019, **A&A**, 622A, 6



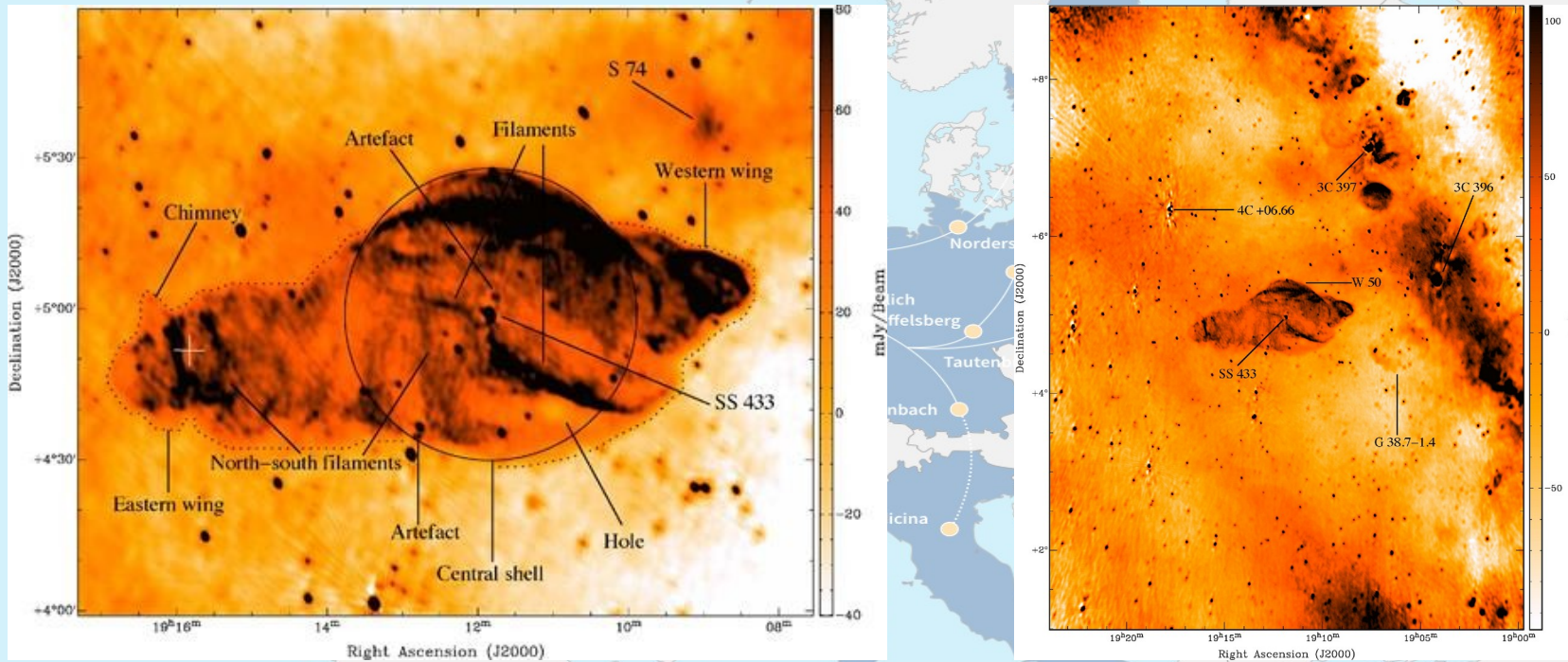
Arias, M., et al., 2022, **A&A**, 667A, 71

Supernova remnants



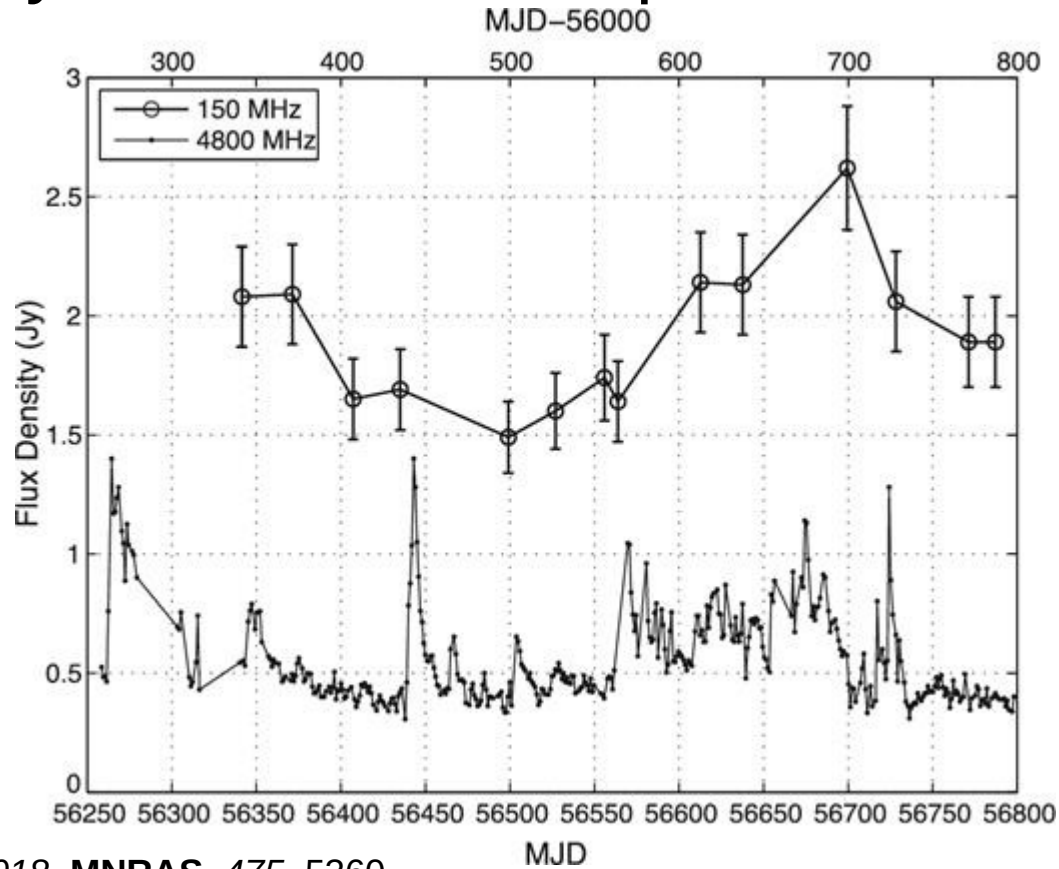
Churazov, E.M., et al., 2022, **MNRAS**, 513L, 83

X-ray binary SS 433 and a supernova remnant W 50



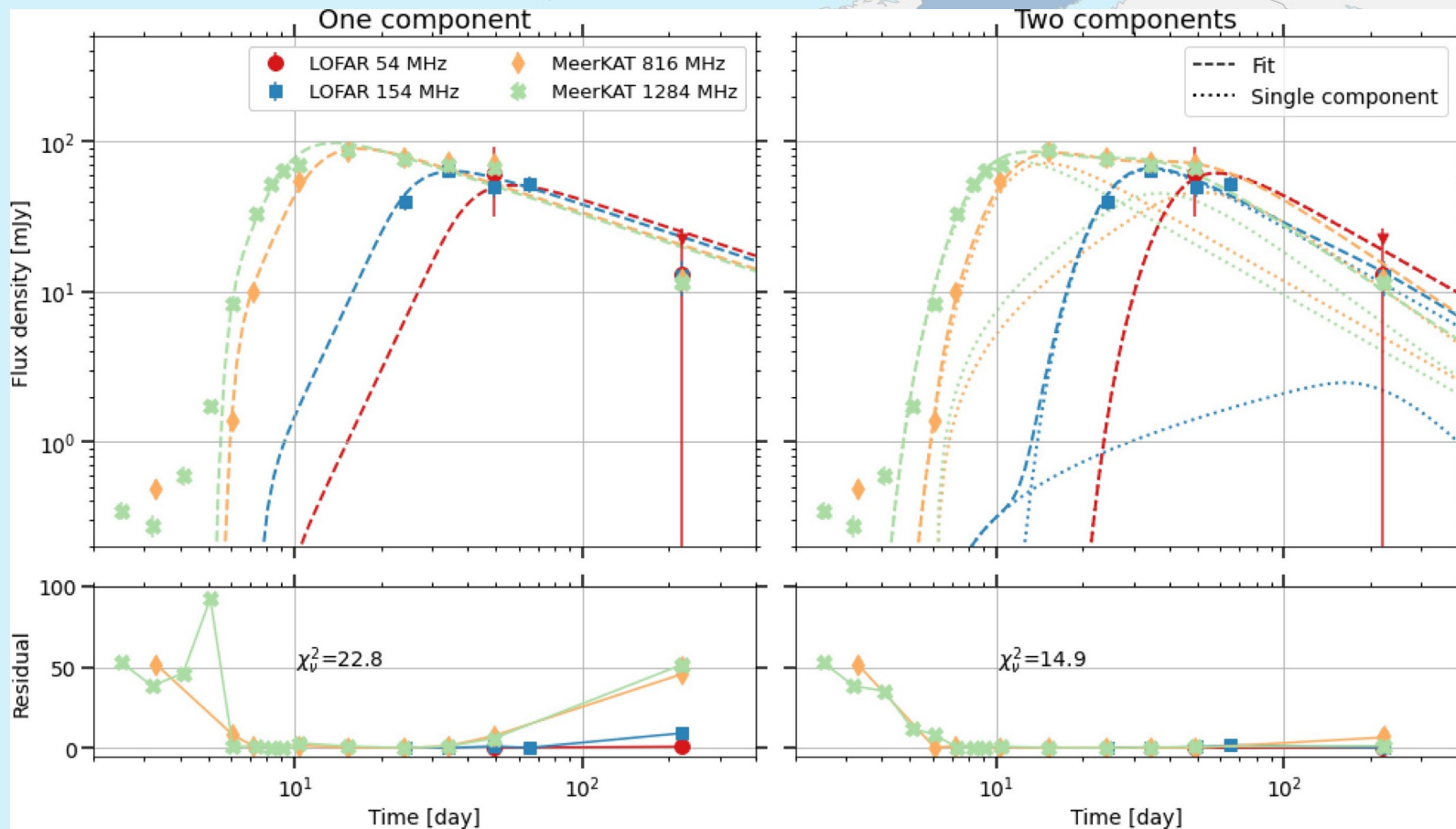
Broderick, J. W., et al., 2018, **MNRAS**, 475, 5360

X-ray binary SS 433 and a supernova remnant W 50

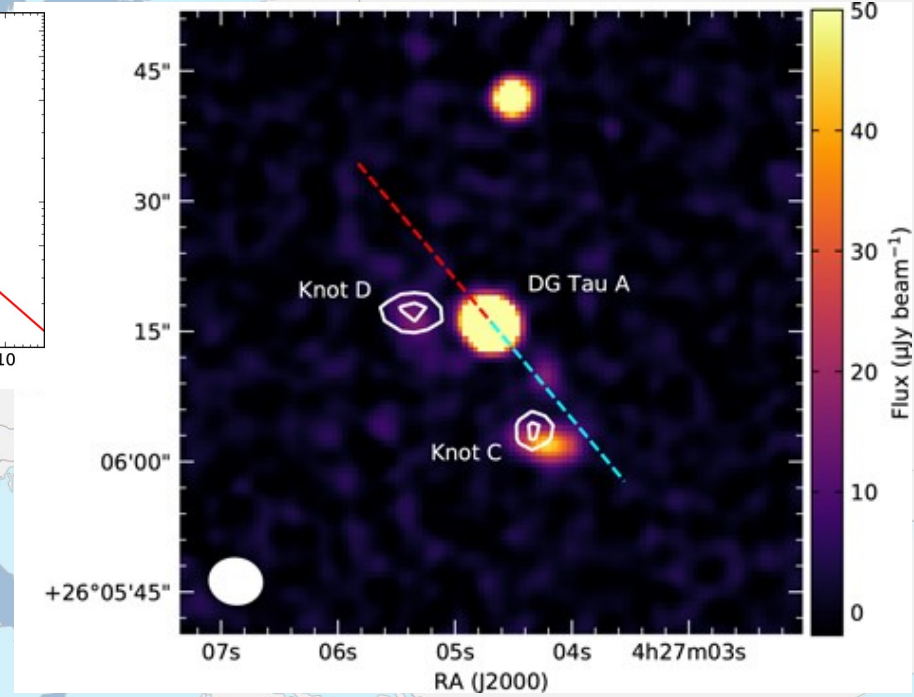
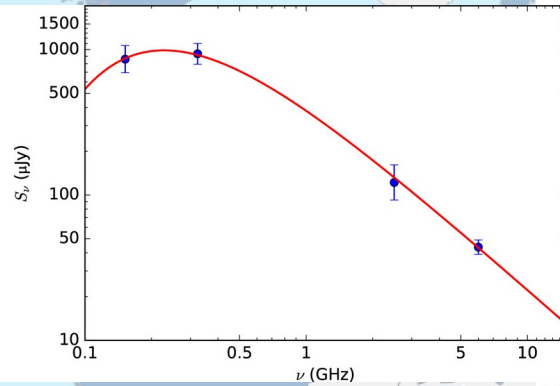
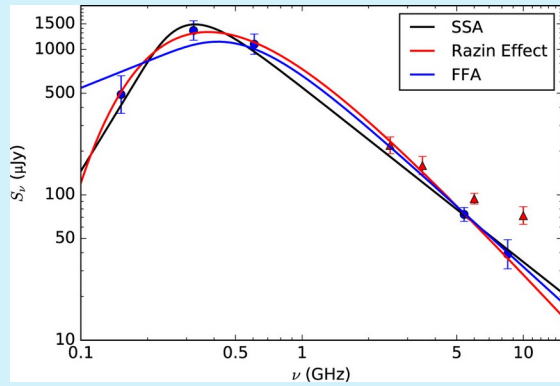


Broderick, J. W., et al., 2018, **MNRAS**, 475, 5360

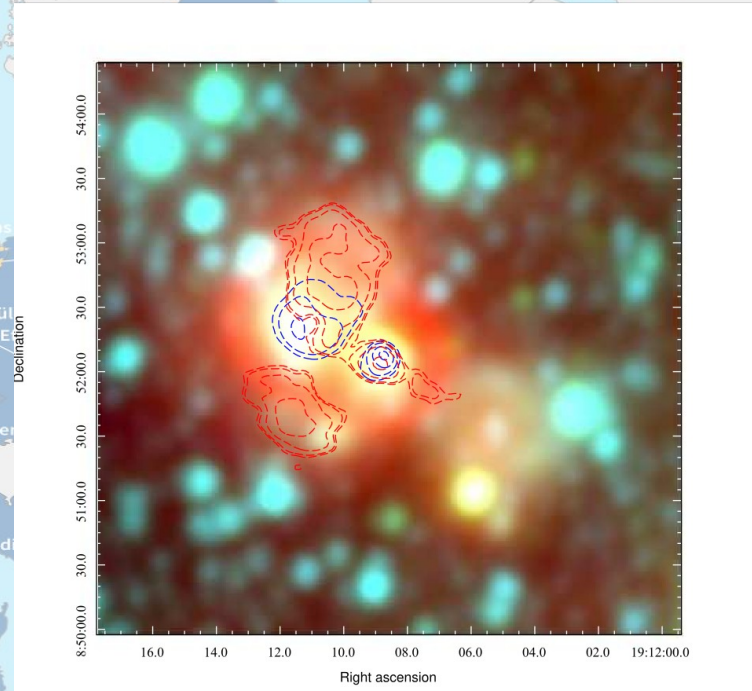
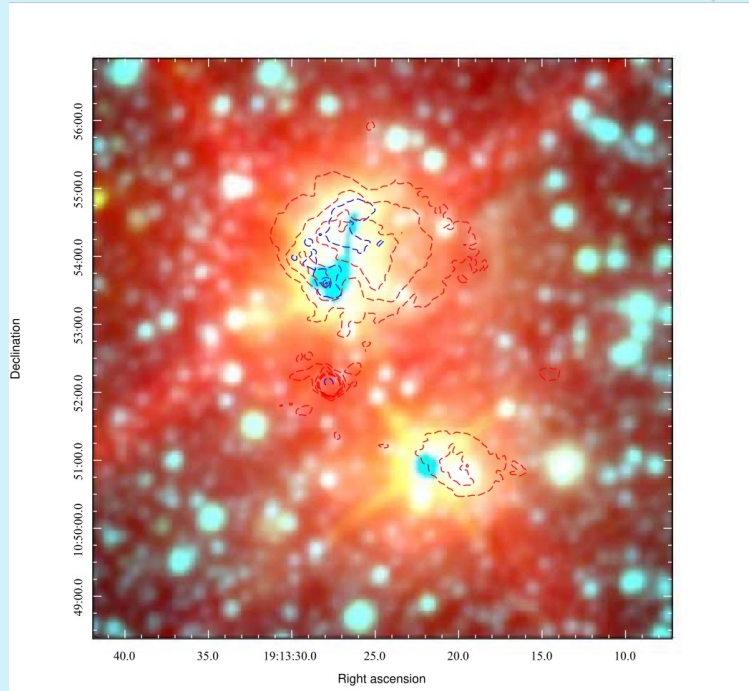
Recurrent nova RS Oph



LOFAR observations of DG Tau

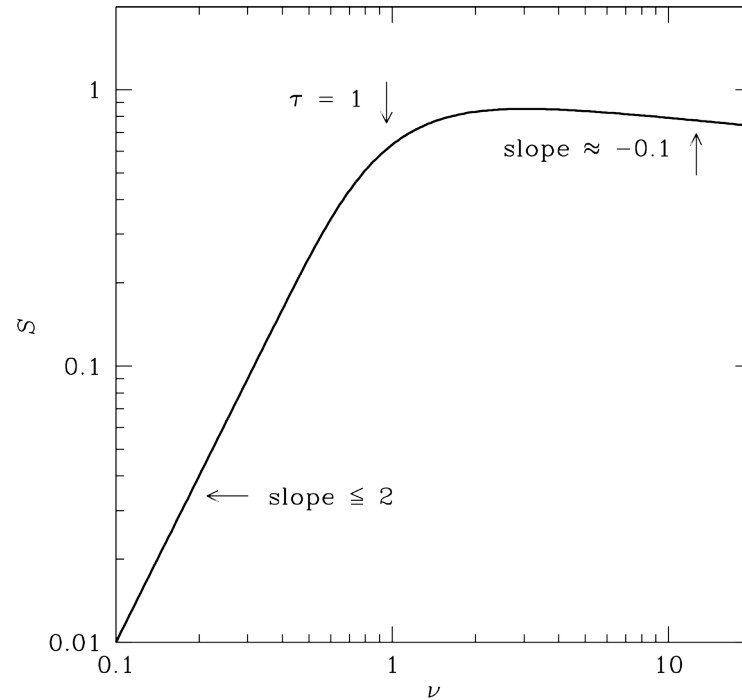


Young stellar objects

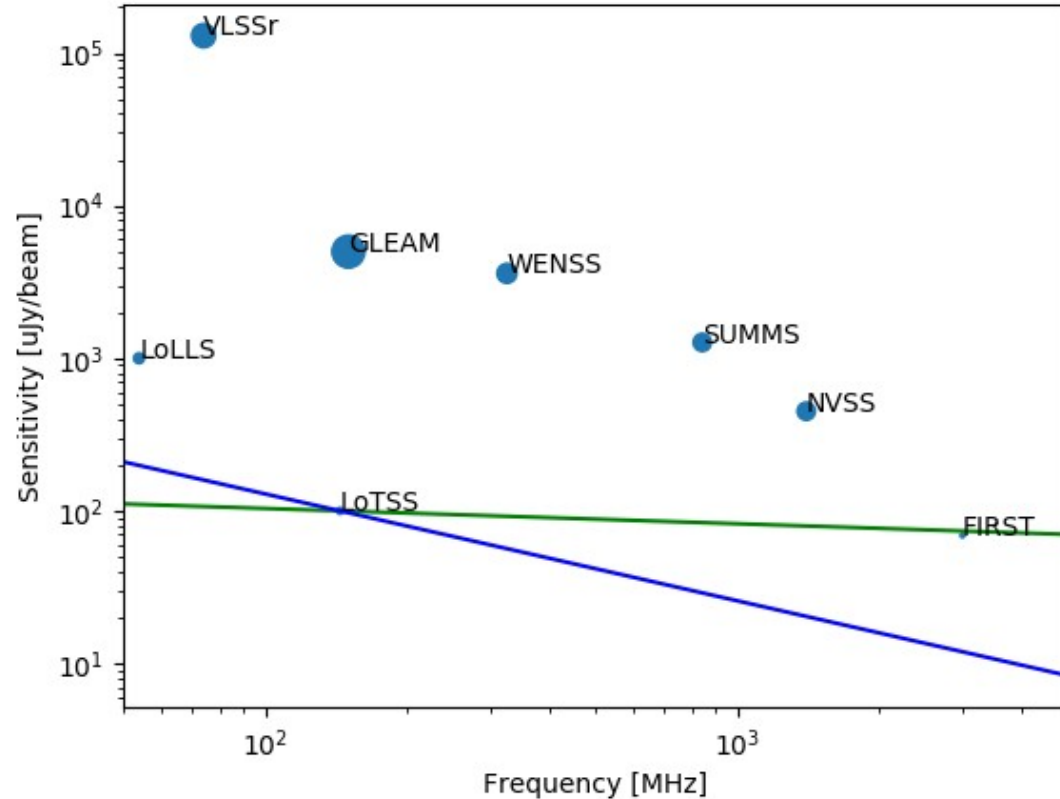


Olech et al., in prep.

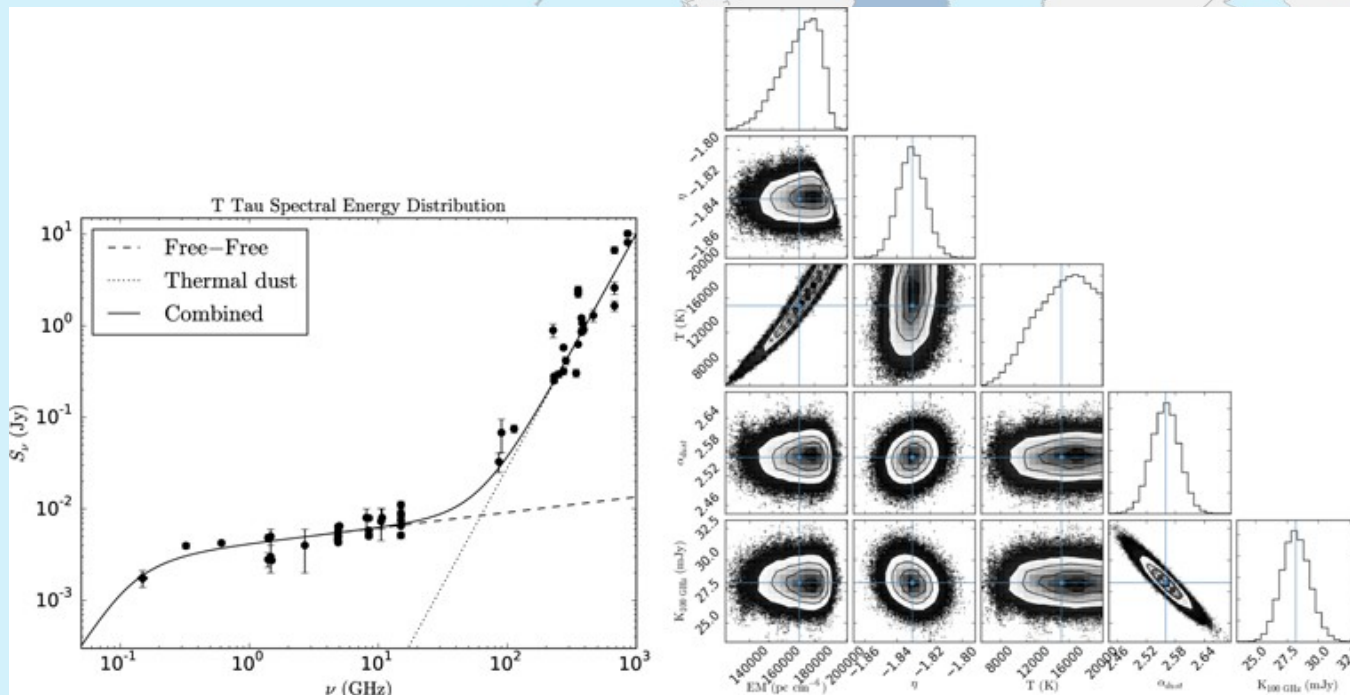
Thermal radio spectrum of an ionized shell



LOFAR Sensitivity for thermal emission

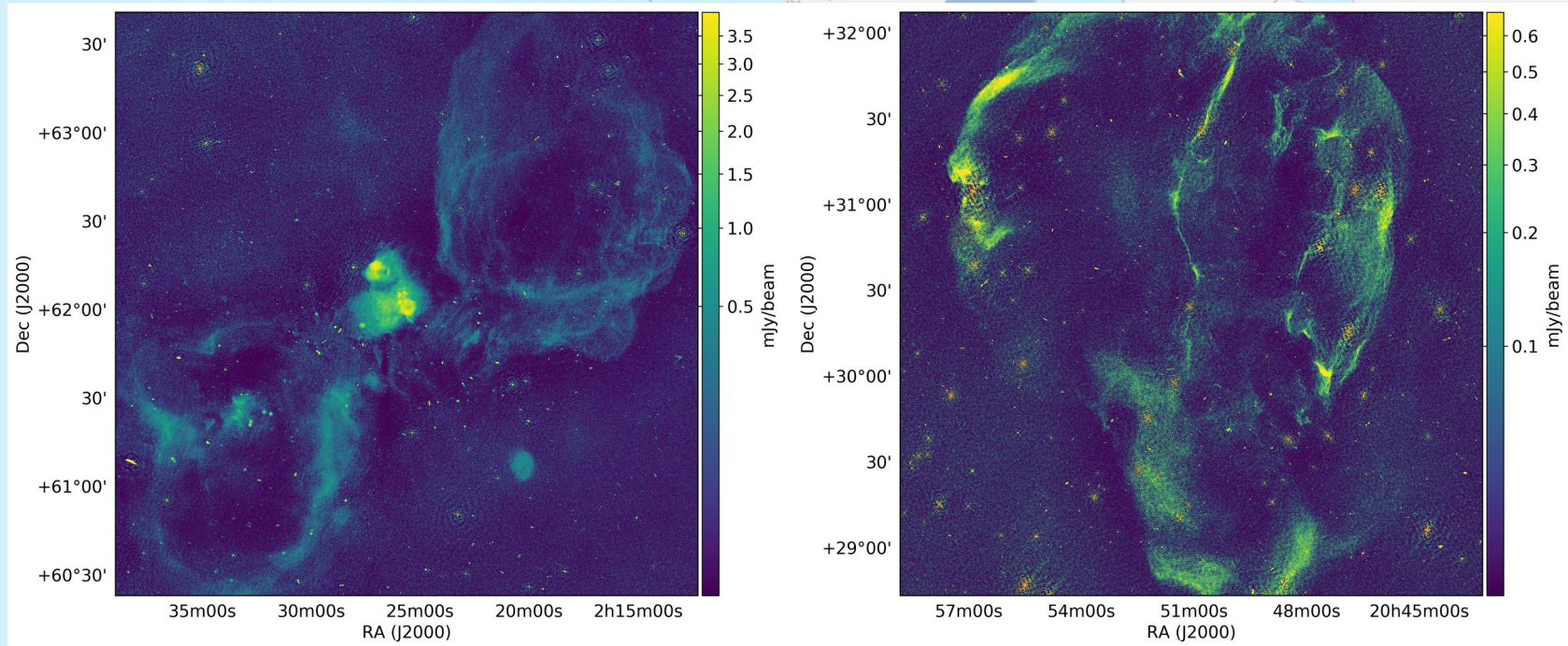


Turnover in T Tau



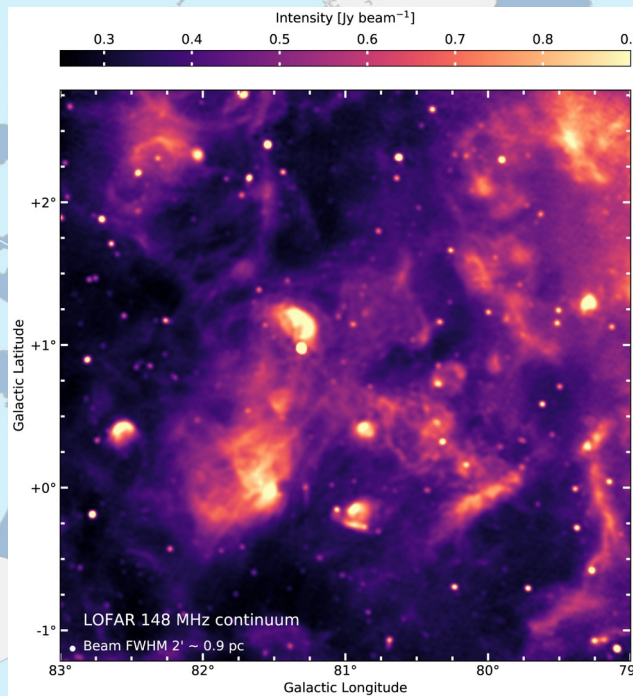
Coughlan C. P., et al., 2017, ApJ, 834, 206

W3/W4/HB3 star forming region (left) and the Cygnus loop supernova remnant



Shimwell, T., et al., 2022, **A&A**, 659A, 1

Cygnus X



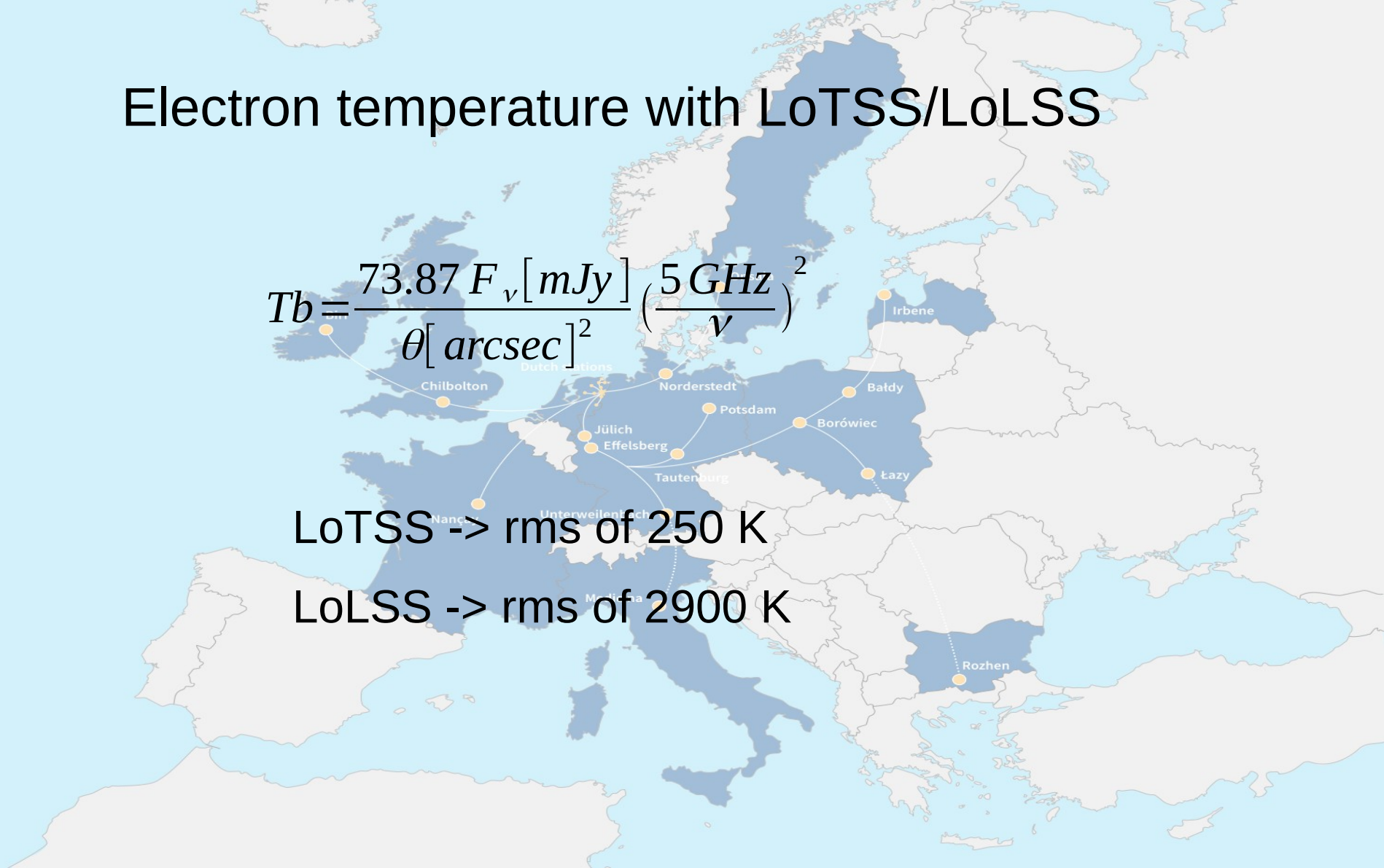
Emig, K. L., et al., 2022, **A&A**, 664A, 88

Electron temperature with LoTSS/LoLSS

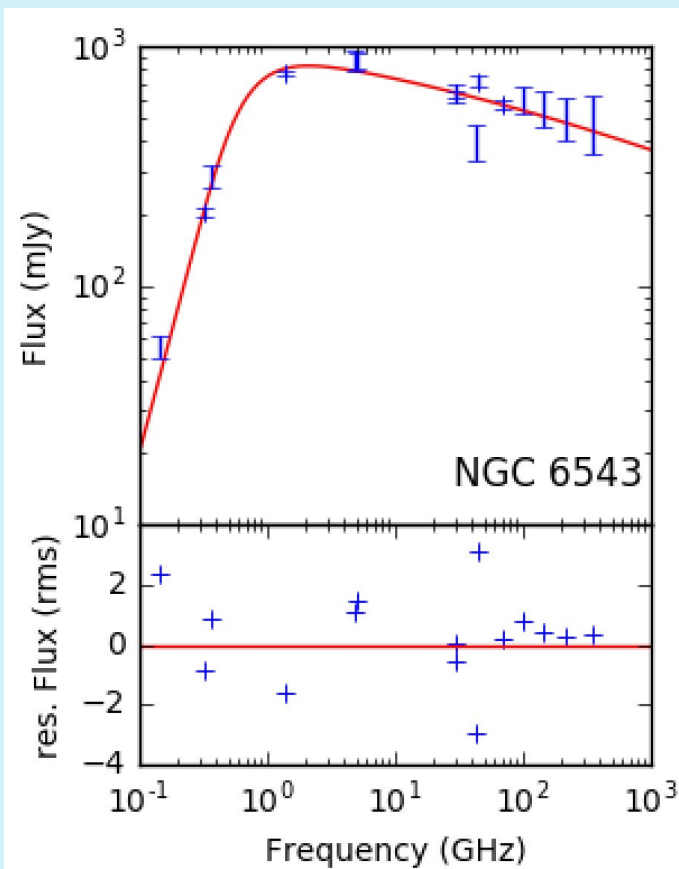
$$T_b = \frac{73.87 F_\nu [mJy]}{\theta [arcsec]^2} \left(\frac{5 GHz}{\nu} \right)^2$$

LoTSS -> rms of 250 K

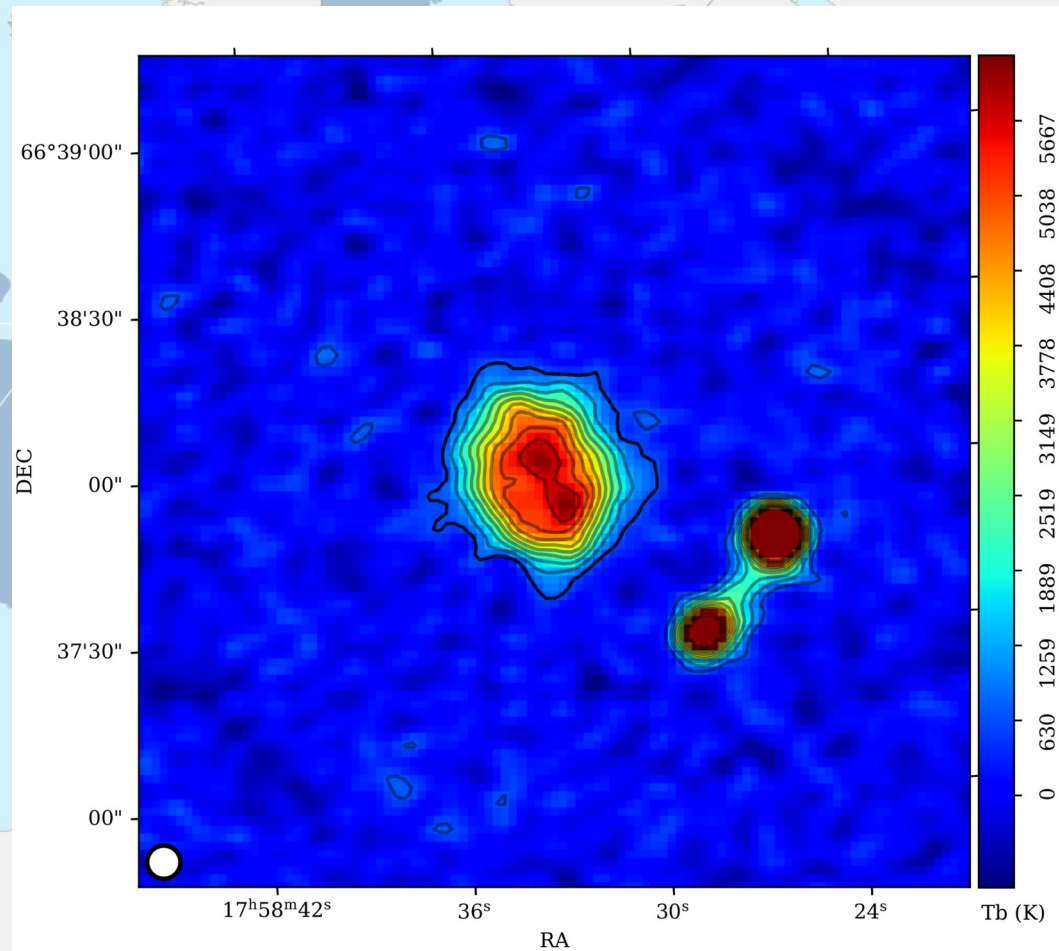
LoLSS -> rms of 2900 K



Cat's Eye planetary nebula (NGC 6543)

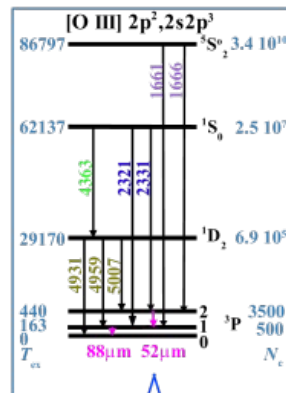


Hajduk, M., et al., 2021, **ApJ**, 919, 121



Electron temperature determination from atomic lines

[O III] collisionally excited lines



CELs

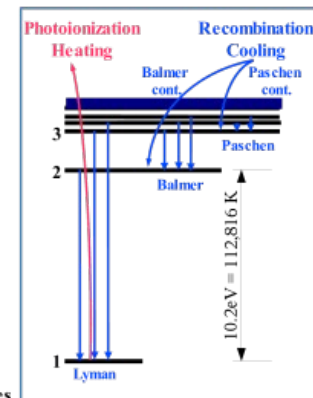
$j_v \propto T_e^{-1/2} \exp(-T_{ex}/T_e)$

$j_v \propto N(X^{+i})N_e$ for $N_e \ll N_e$

$\propto N(X^{+i})$ for $N_e \gg N_e$

j_v increases as T increases

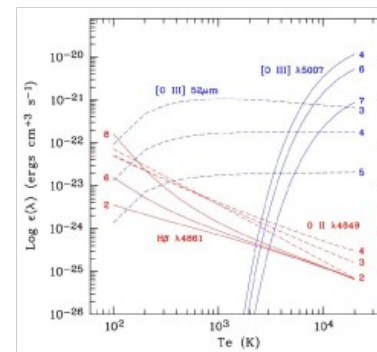
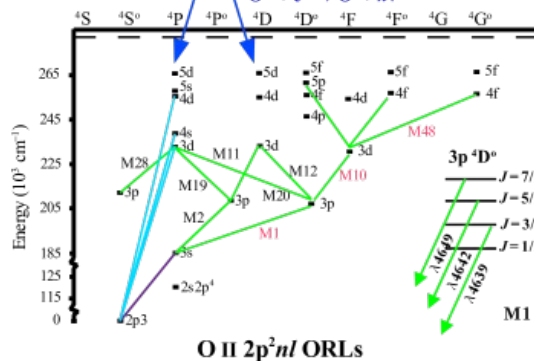
H I recombination lines/continua



ORLs/Cont.

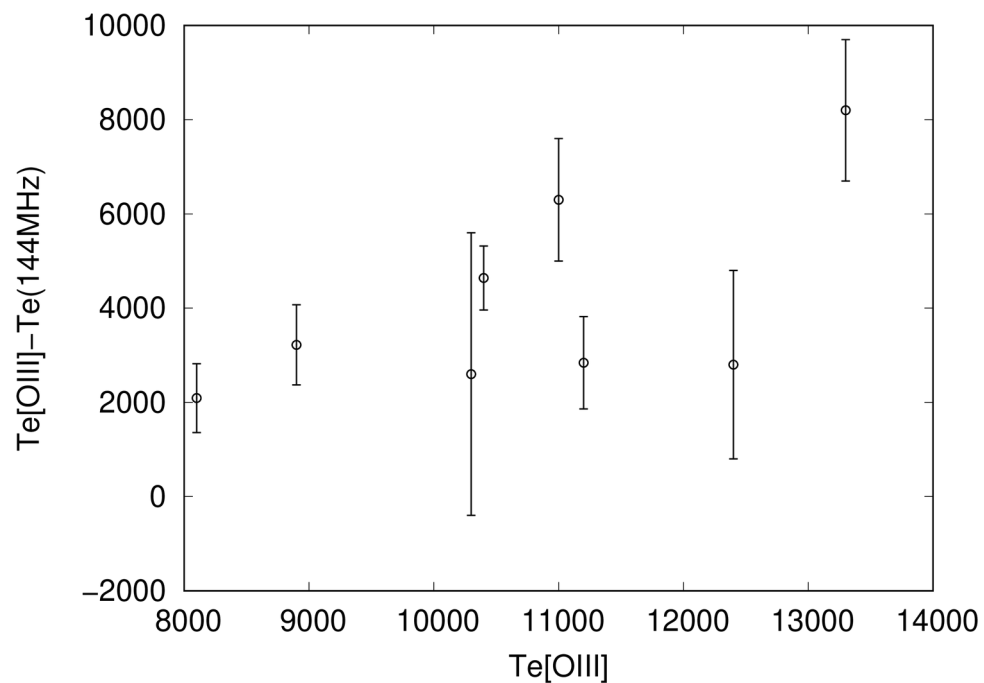
- $j_v \propto T_e^{-\alpha}$ where $\alpha \sim 1$
- $j_v \propto N(X^{Z+1})N_e$
- j_v **decreases** as T_e increases

Recombination

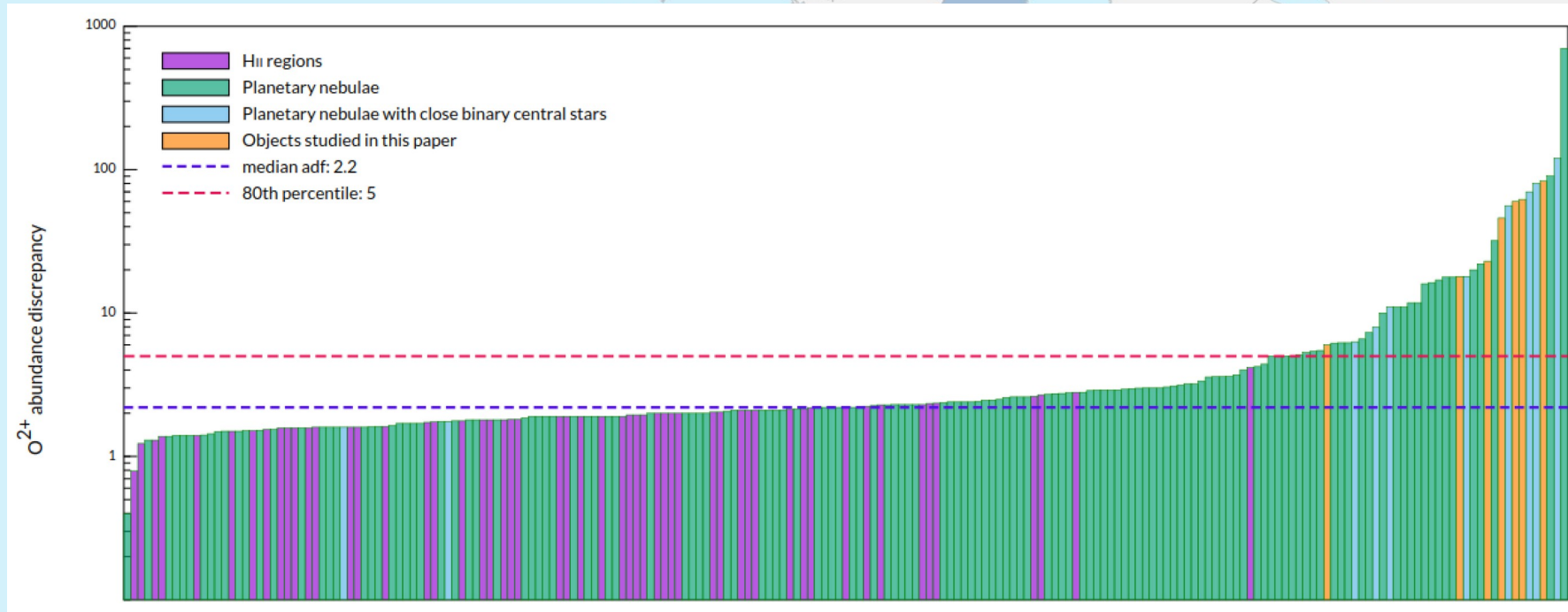
$$\text{O}^{2+} + \text{e}^{-} \rightarrow \text{O}^{+} + h\nu$$


Emissivities per ion per electron

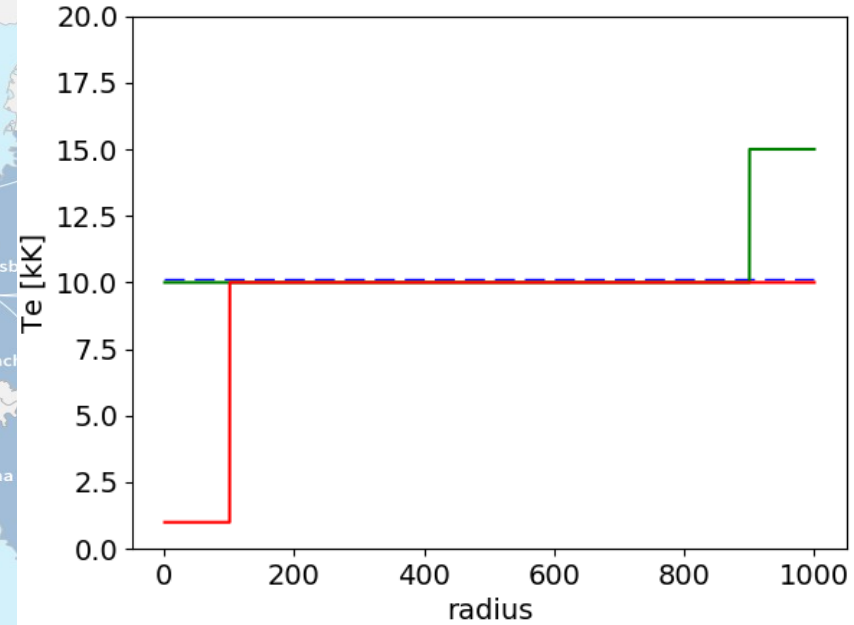
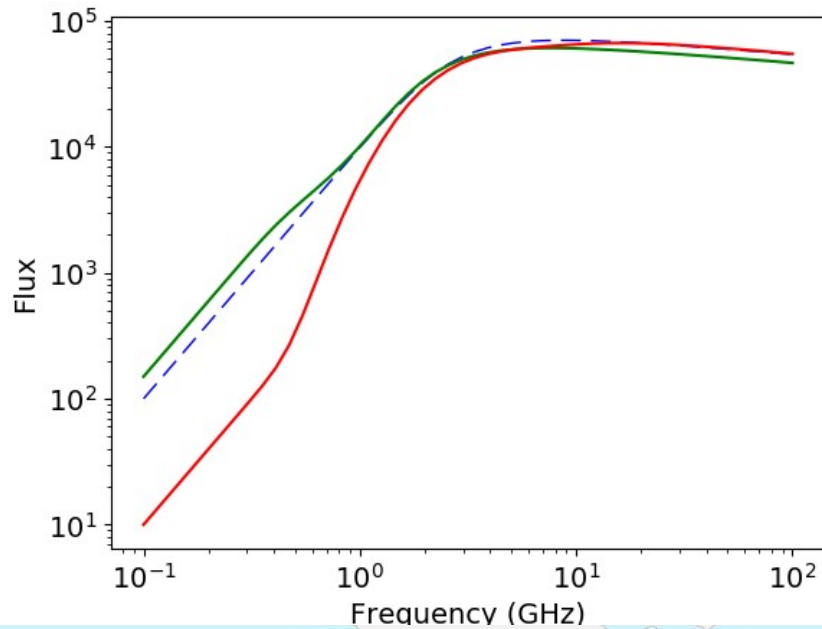
Temperature difference in $T_e(144\text{MHz})$ and $T_e([\text{OIII}])$



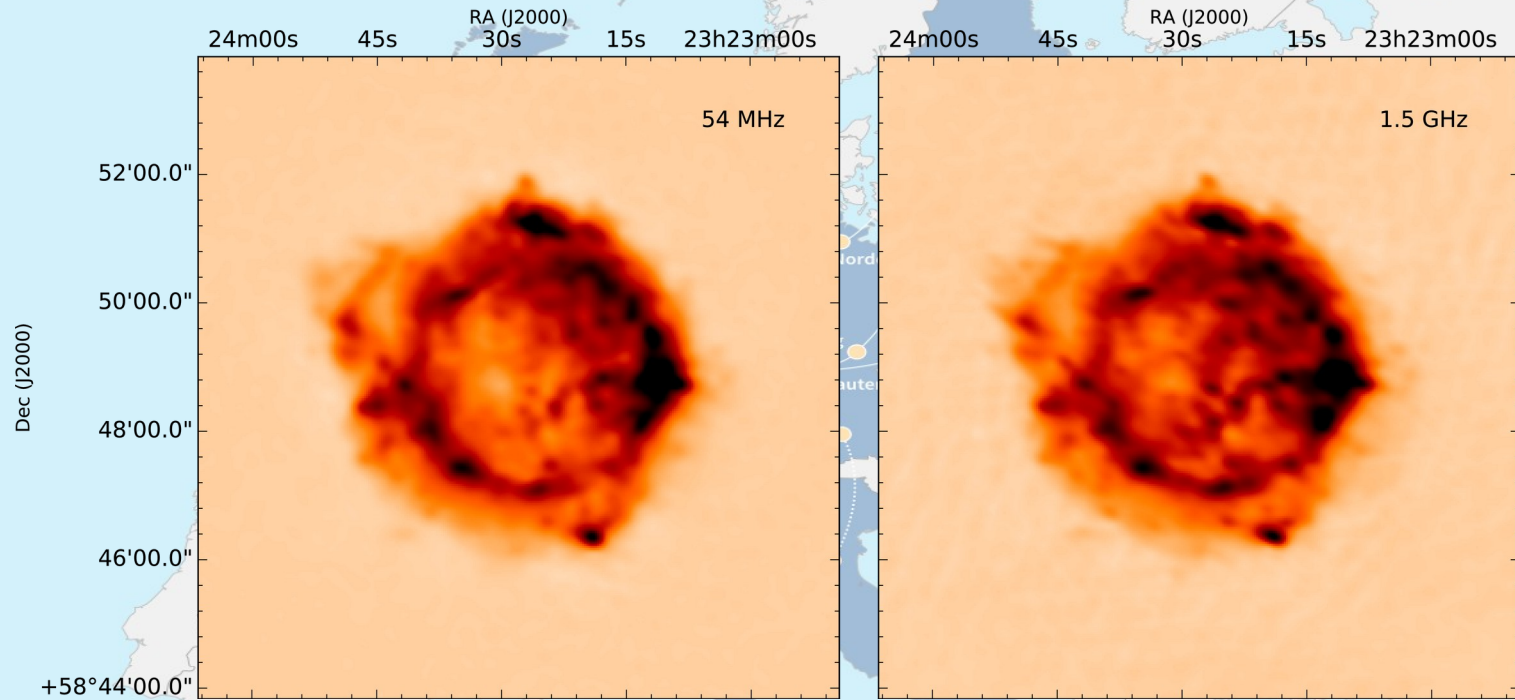
Abundance discrepancies in PNe and HII regions



Temperature gradients

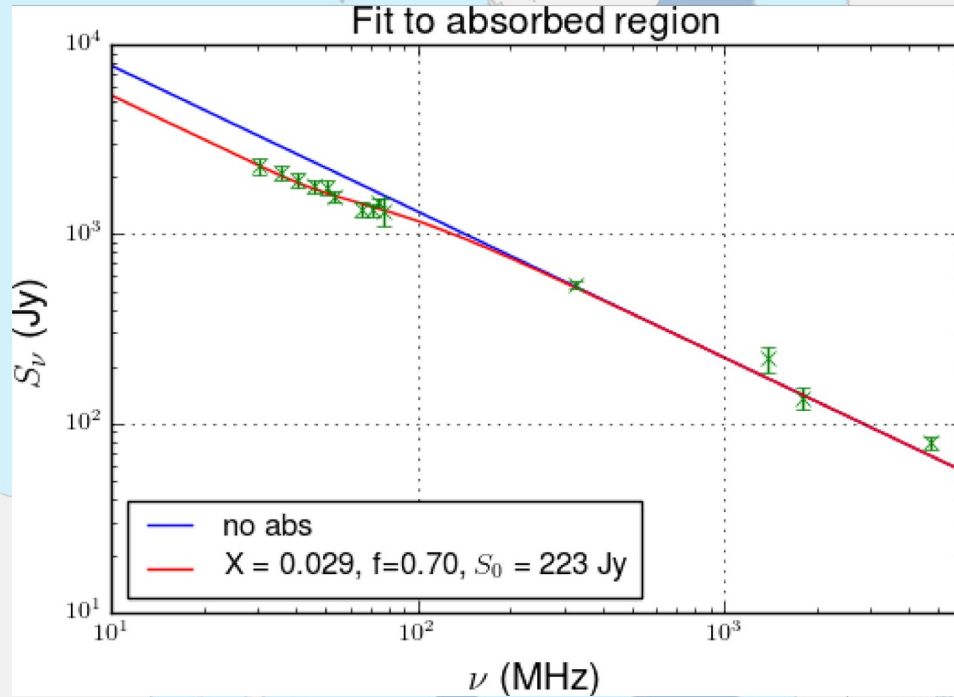


Free-free absorption in Cas A

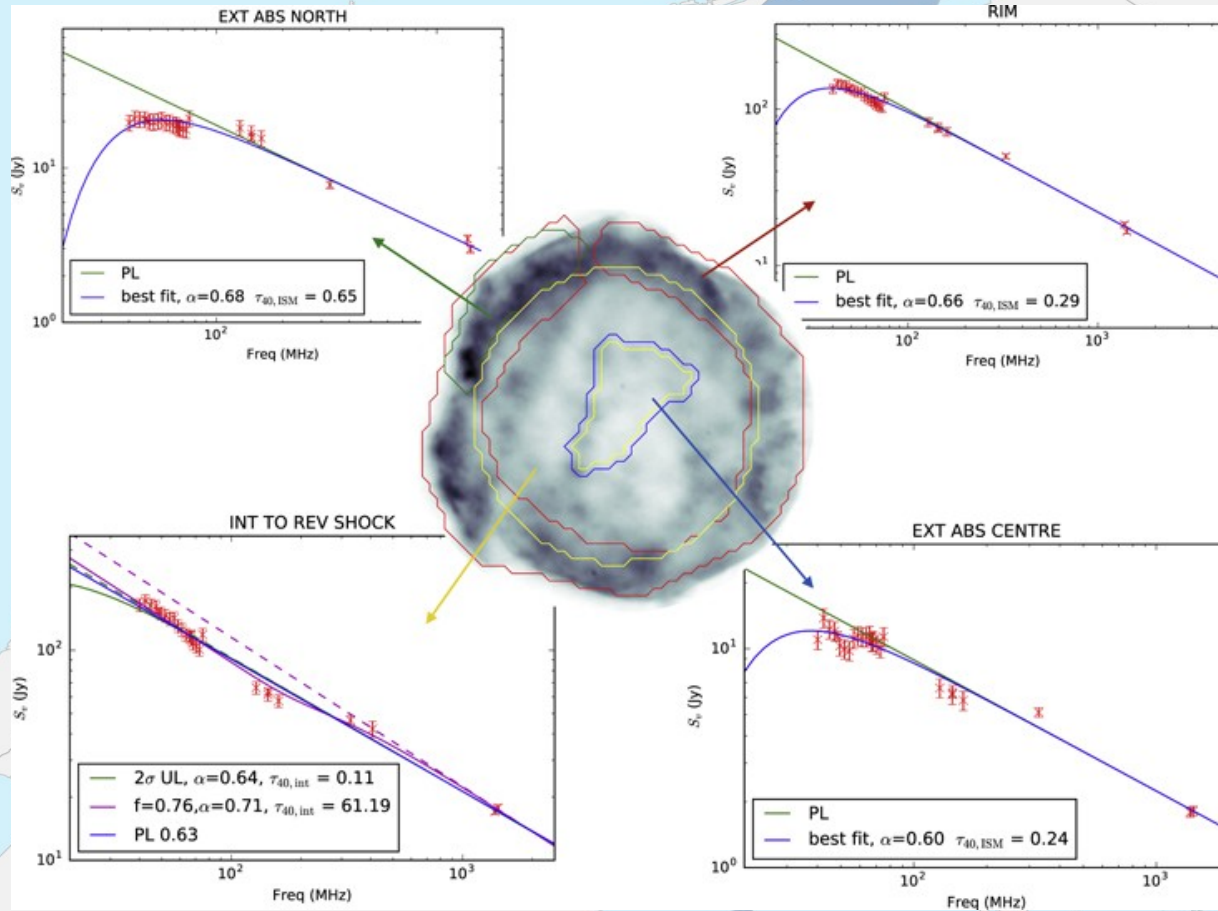


Arias, M., et al., 2018, **A&A**, 612A, 110

Free-free absorption in Cas A



Tycho's Supernova Remnant



LBA and HBA observations of HII regions

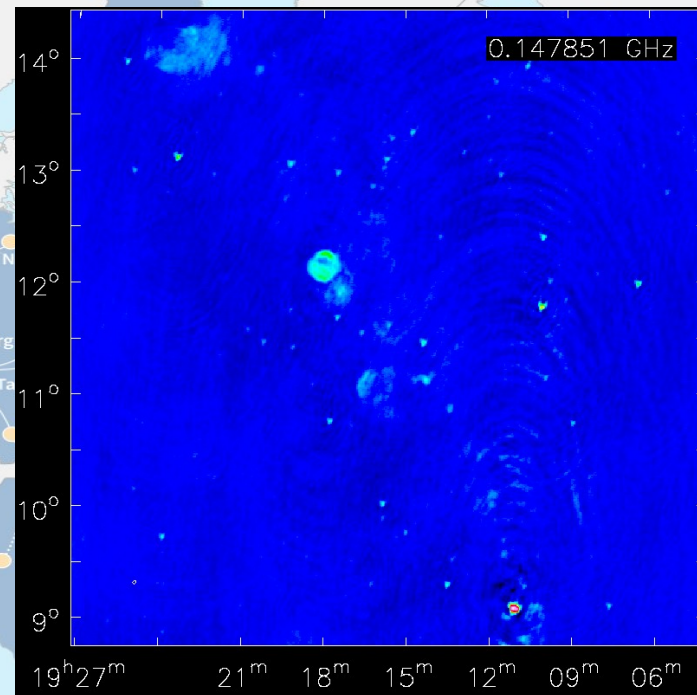
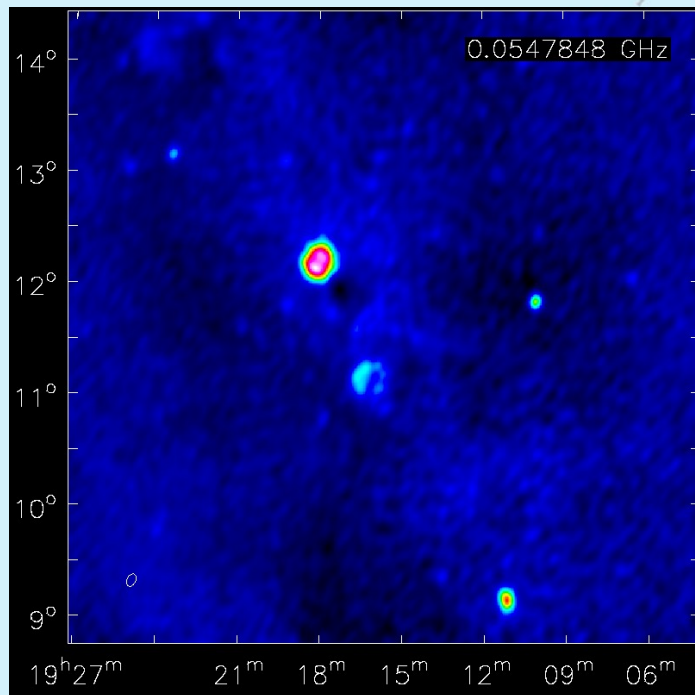


Image courtesy of Irene Polderman

LBA and HBA observations of HII regions

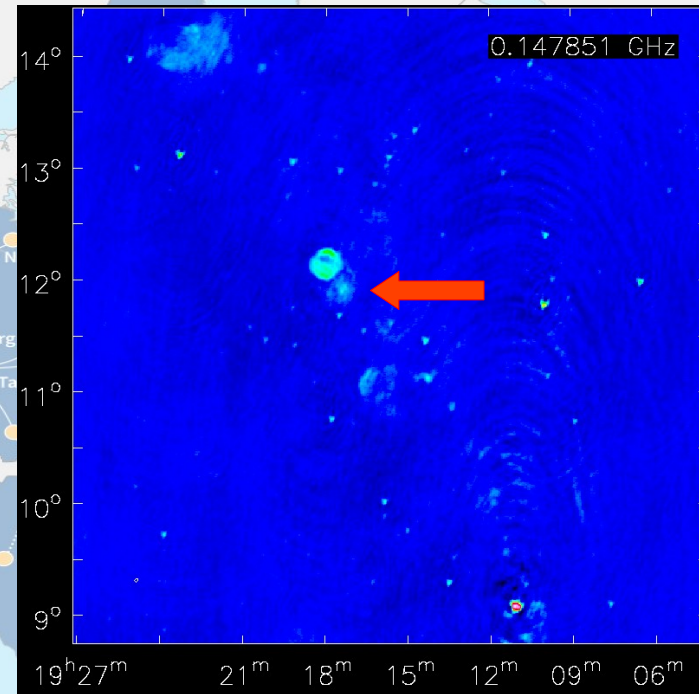
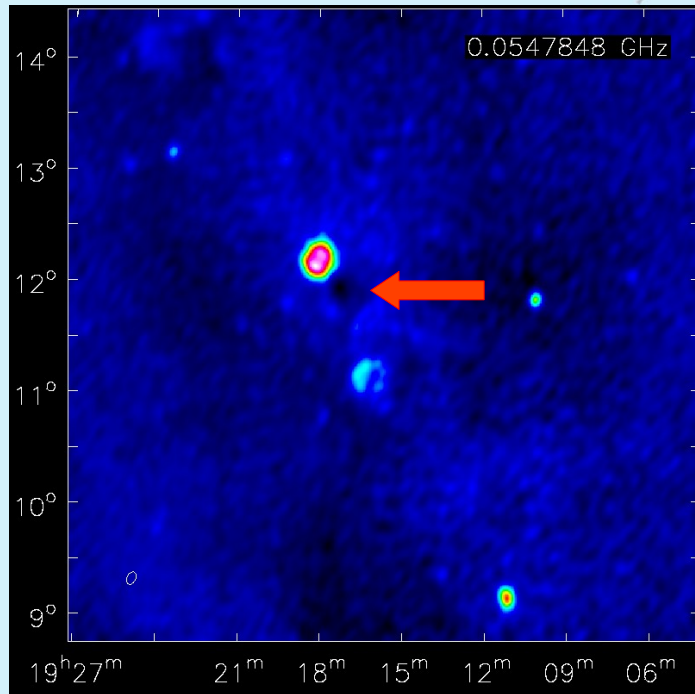
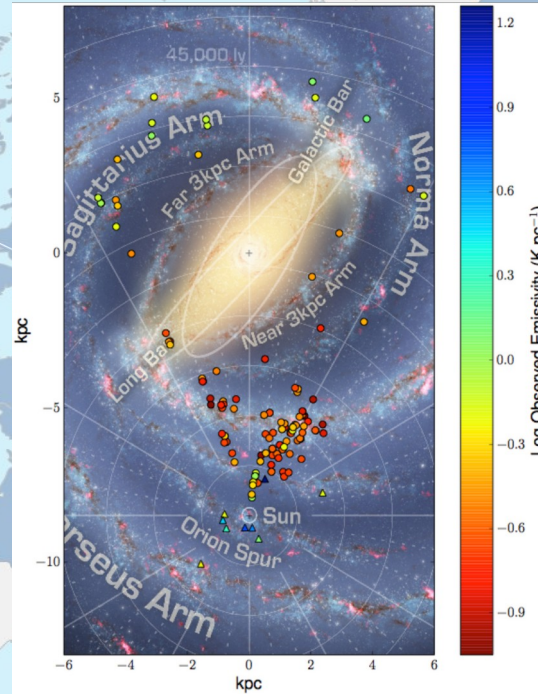


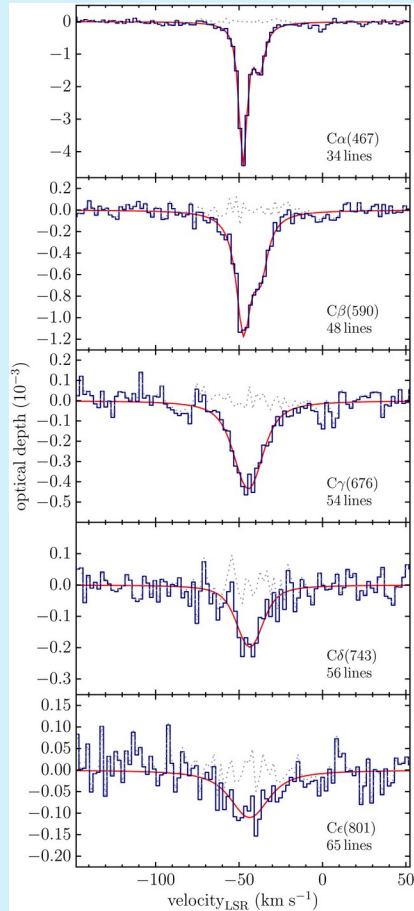
Image courtesy of Irene Polderman

Studying Galaxy using HII regions

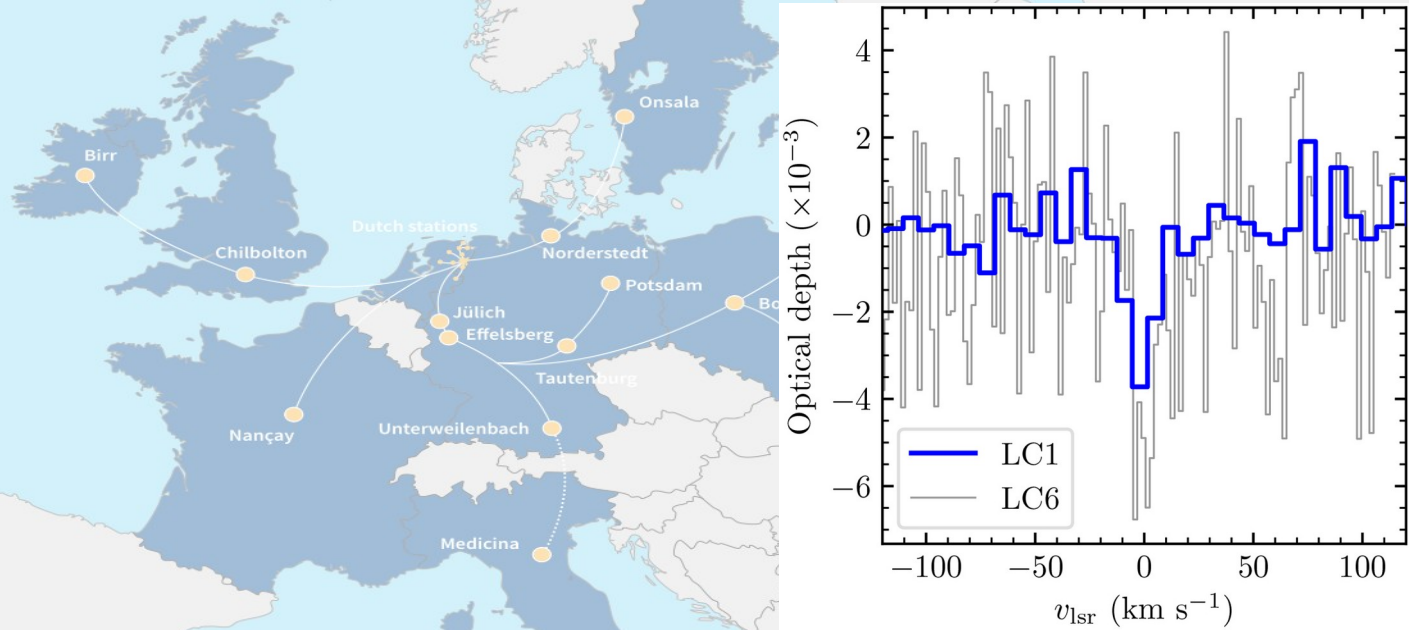


Poldrman, I., et al., 2019, **A&A**, 621A, 127

Radio recombination lines

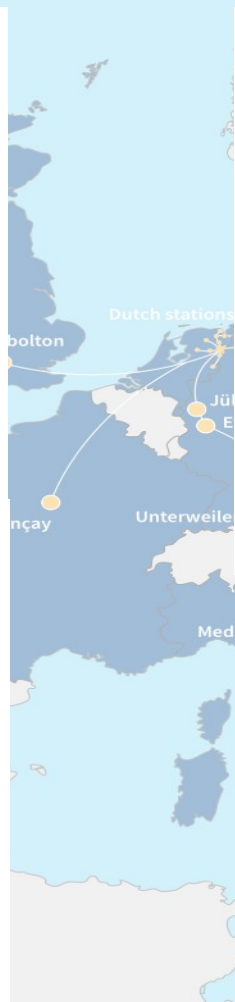
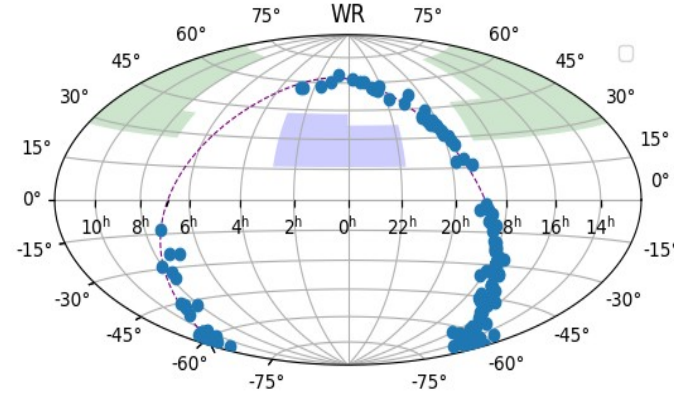
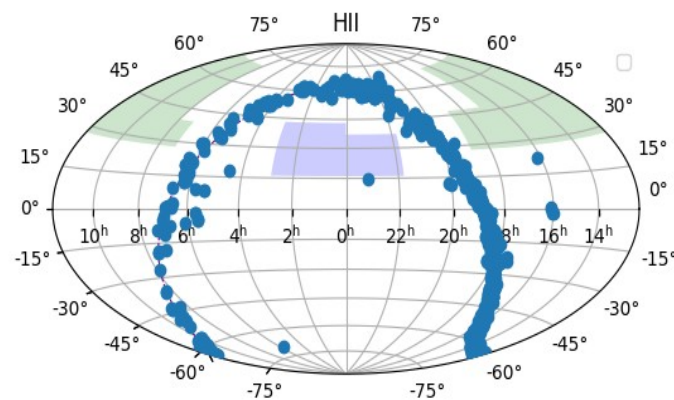
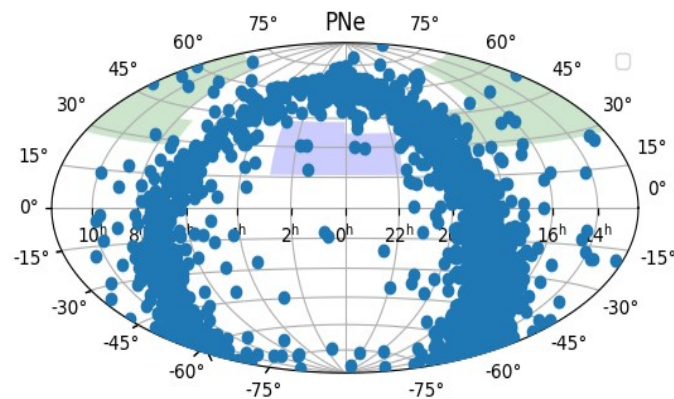
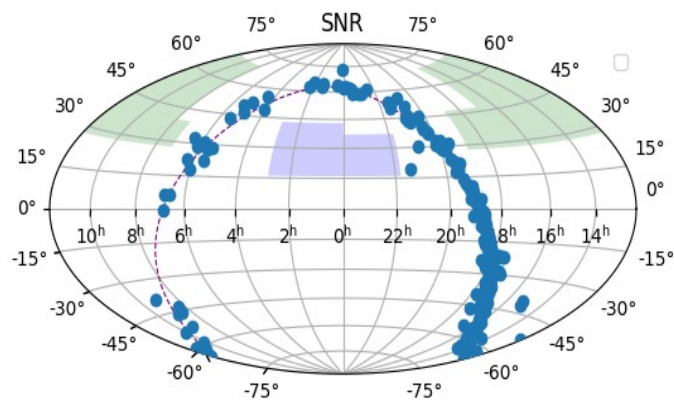


Emig, K. L., et al., 2020, **A&A**, 634A, 138

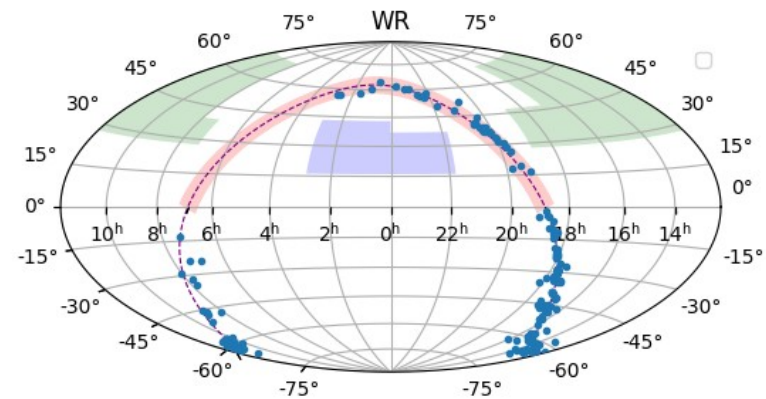
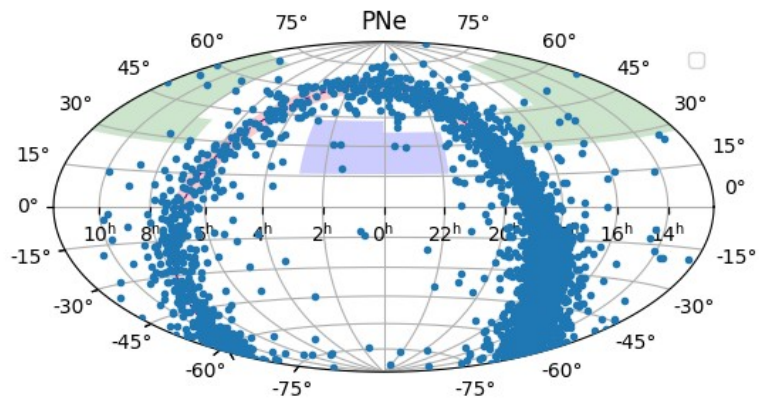
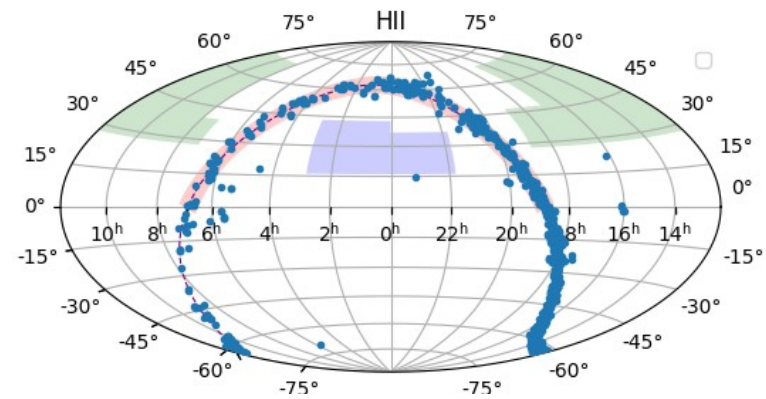
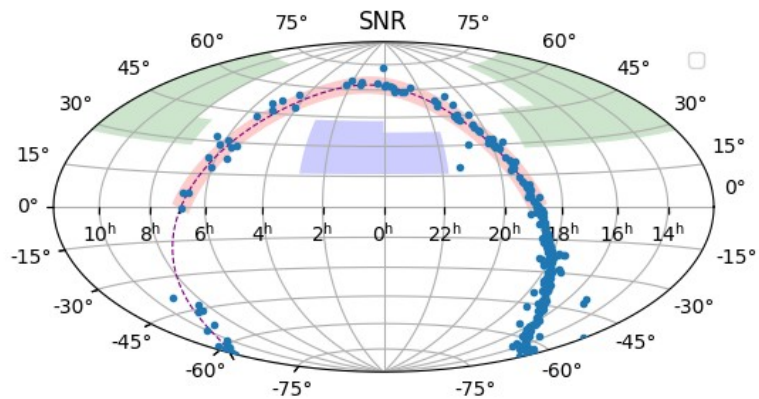


Salas, P., et al., 2019, **A&A**, 626A, 70

Sky distribution of Galactic nebulae

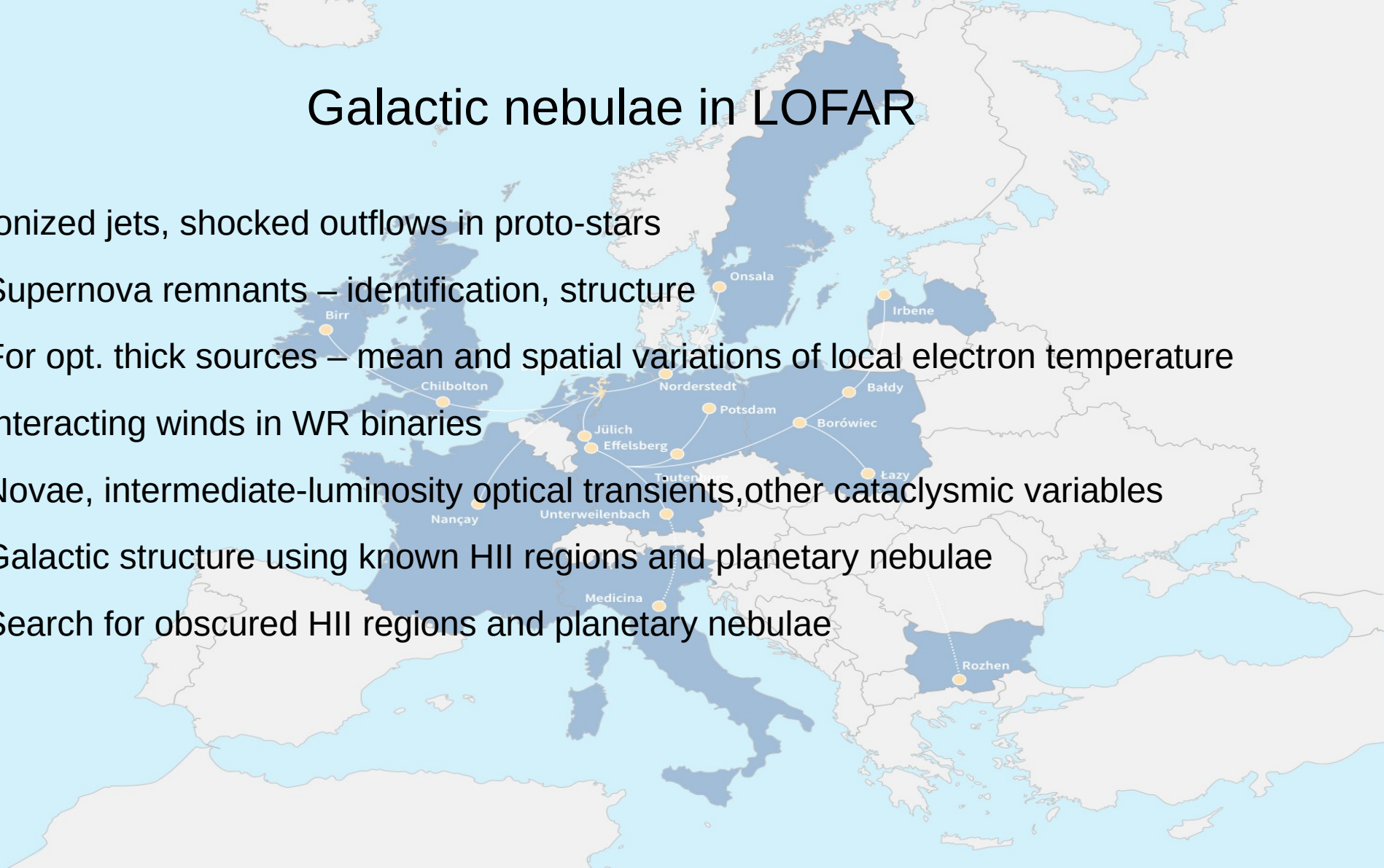


L2GPS – LOFAR 2.0 Galactic Plane Survey

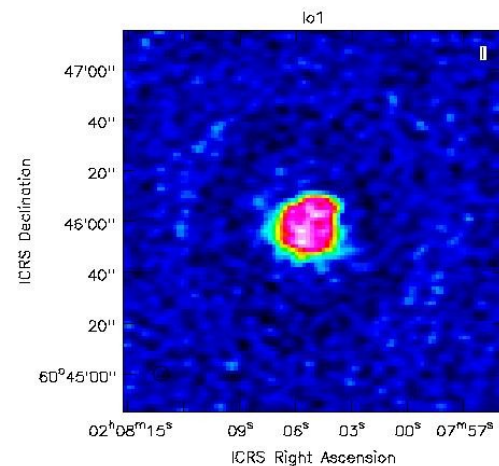
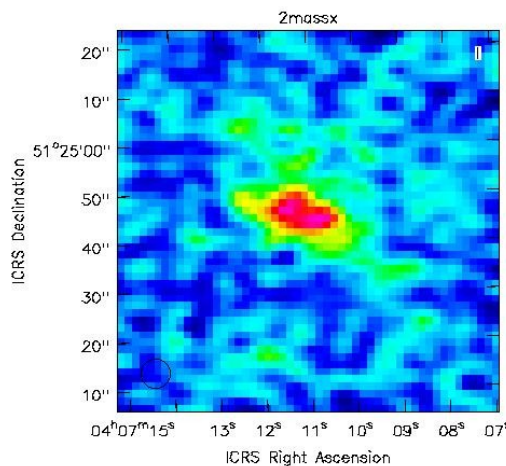
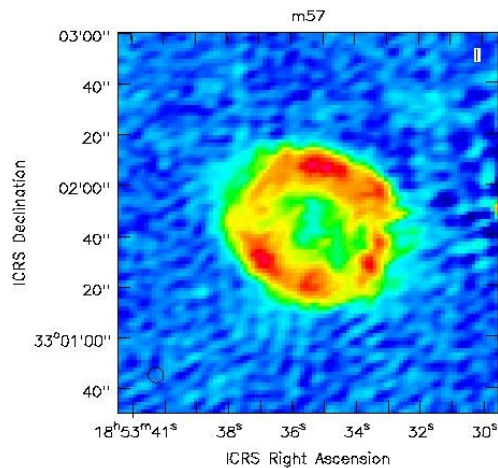
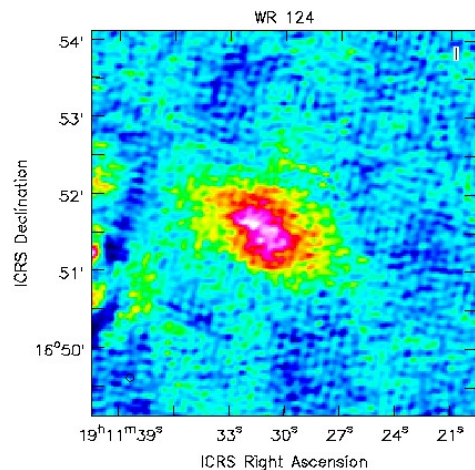


Galactic nebulae in LOFAR

- Ionized jets, shocked outflows in proto-stars
- Supernova remnants – identification, structure
- For opt. thick sources – mean and spatial variations of local electron temperature
- Interacting winds in WR binaries
- Novae, intermediate-luminosity optical transients, other cataclysmic variables
- Galactic structure using known HII regions and planetary nebulae
- Search for obscured HII regions and planetary nebulae



Galactic nebulae in LOFAR



For extragalactic objects see talks in the Nearby Galaxies session, e.g. by Etienne Bonnassieux, Lovorka Gajović, Deepika Venkattu