# Studying Single Pulse Emission from Pulsars using POLFAR

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#### Pulsed Radio Emission from Pulsars



[Image Credit : Manchester & Taylor (1977, top left); Mitra et al. (2016) - MSPES survey (bottom left and right).]

#### Prominent Single pulse Phenomena in Pulsars

• **Subpulse Drifting** - Systematic shift of the emission components (subpulses) within the pulsed window.

Nulling

Mode changing

Periodic Modulation

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### Subpulse drifting



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# Longitude Resolved Fluctuation Spectra (LRFS)



PSR B0818-13 (Basu et al. 2016).

PSR B2319+60 (Rahaman et al. 2021).

LRFS - Drifting periodicity ( $P_3 = 1/f_p$ ) and Phase variations across profile.

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#### Prominent Single pulse Phenomena in Pulsars

Subpulse Drifting

• **Nulling** - The single pulse emission is below detection threshold. Repeats at regular but mostly random intervals.

Mode changing

Periodic Modulation

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# Nulling



Nulling Fraction; Distribution of Null lengths and Burst lengths; Degree of intensity reduction; etc.

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### Prominent Single pulse Phenomena in Pulsars

Subpulse Drifting

Nulling

• **Mode changing** - Change in the emission properties like Profile shape, drifting periodicity, etc. Rapid transition between modes with large variation of mode lengths from few 100 pulses to hours at a time.

Periodic Modulation

### Emission Mode Changing



Basu, Mitra & Melikidze (2021).

- Rapid transition between two distinct emission states.
- Change in Profile shape; Subpulse Drifting behaviour; etc.

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### Prominent Single pulse Phenomena in Pulsars

Subpulse Drifting

Nulling

Mode changing

• **Periodic Modulation** - Emission within the entire window periodically changes in intensity.

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#### Periodic Modulation



- Two types : Periodic nulling; Periodic amplitude modulation.
- Longer periodicities and constant phase variations.

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#### Observations with GMRT

#### • Meterwavelength Single-pulse Polarimetric Emission Survey (MSPES)

(Mitra et al. 2016, 2023; Basu et al. 2016, 2017, 2021)

- 123 pulsars, with P > 0.1 s, observed with GMRT.
- $\blacktriangleright$  ~ 200 hours of observing time.
- Observations at both 325 MHz and 610 MHz frequency bands.
- Around 2100 polarized single pulses observed from each pulsar at both frequencies.

#### Assorted Studies

(Basu & Mitra 2018a, 2018b, 2019; Basu et al. 2019a, 2019b, 2020; Mitra et al. 2020; Rahaman et al. 2021)

- A separate list of >100 pulsars from multiple observing proposals as well as archival data (Mitra & Rankin, 2011; and other previously unpublished observations).
- More than 500 hours of observing time in total.
- Observations primarily at 325 MHz frequency band.
- Longer observations (> 5000 single pulses) on specific pulsars as well as shorter observations (~ 2000 pulses) on larger samples.

#### Observations with GMRT : Results

#### Subpulse drifting :

- Detailed classification of the drifting behaviour in  $\sim$ 80 pulsars.
- Phase variations characterised for the first time in majority of them.
- 20% increase in drifting population.
- Dependence on spin-down energy loss  $(\dot{E})$ .

#### • Periodic Modulation :

- First identified as a separate phenomenon in pulsar radio emission.
- ▶ Found ~70 pulsars with periodic modulations.
- Classification into two groups : Periodic Nulling and Periodic Amplitude Modulation.
- $\blacktriangleright~\sim\!15$  pulsars with both subpulse drifting and periodic modulation seen as separate peaks in the Fluctuation spectra.

#### • Nulling and Mode Changing :

- Studied nulling behaviour in around 60 pulsars. Several new detections.
- Characterized the emission modes in around 10 pulsars, which represent around 30% of the population.

#### Subpulse drifting and Periodic Modulation



Basu, Mitra & Melikidze (2020).

- Subpulse drifting seen in low energetic pulsars with  $\dot{E} < 5 imes 10^{32} \mbox{ erg s}^{-1}$
- Bunching of drifting periodicity,  $P_3$ , around 2P and  $P_3 \propto \dot{E}^{-0.5}$

• Periodic modulation seen across the entire  $\dot{E}$  range, mostly longer periodicities > 10P.

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### Observations with Individual POLFAR stations

• The steep spectral nature of Pulsars makes it possible to detect single pulse emission from several pulsars at lower frequencies using Individual POLFAR stations.

• Possibility of longer observations repeated at regular intervals. Ideally suited for emission mode changing behaviour with timescales of hours.

• Characterise the low frequency single pulse behaviour in pulsars to complement higher frequency results.

#### POLFAR observation of PSR B0809+74 : Mode Changing



• Prominent Subpulse drifting seen for 95% of observing time.

• Switches to a short lived Slow-Drift mode for around 50 pulses.

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#### POLFAR observation of PSR B0809+74 : Subpulse Drifting



• Normal Mode :  $P_3 = 11P$ ; Slow-Drift Mode :  $P_3 = 16P$ 

• Linear Phase variations remaining unchanged in the two modes.

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#### POLFAR observation of PSR B0809+74 : Nulling



• Two different types of nulls, short duration 1-2P and longer duration lasting 5-10P.

• The short nulls repeat with periodicity of 44P.

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#### POLFAR observation of PSR B1237+25 : Flaring



- Emission changes with frequency, showing flaring bursts at low frequency.
- The typical intensity of these flares are 20 times the average intensity.

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# Frequency evolution of Drifting



Basu, Melikidze & Mitra (2023)

- Radio emission in outflowing plasma generated due to sparking discharge in an Inner Acceleration Region (IAR) above polar cap.
- The sparks undergo **E**×**B** drift in the IAR resulting in subpulse drifting.
- The drifting phase behaviour is expected to change due to the frequency evolution of the line of sight traverse.
- Program to compare drifting phase behaviour at higher frequencies from GMRT and low frequency using individual POLFAR stations.

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## Summary

- The single pulse emission from pulsars shows a number of interesting phenomena like subpulse drifting, nulling, emission mode changing, periodic modulation, etc., that are closely associated with the physical processes within the pulsar magnetosphere.
- Understanding the nature of each of these phenomena requires dedicated observations of a sizeable sample of pulsars to characterize their behaviour in individual sources as well as finding general trends in the larger population.
- The high frequency single pulse behaviour in a significant sample of pulsars have been conducted in the past few years using the GMRT.
- We are currently in the process of extending these studies to the low frequency regime using the individual POLFAR stations, with several pulsars like PSR B0809+74 and PSR B1237+25 providing new insights which are distinct from the high frequency behaviour.

# Thank you for your attention!

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