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CROSSING the SIZE LIMIT of RADIO SOURCES with LOFAR

with Dorota Kozieł, Natalia Żywucka, Arti Goyal, Marek Jamrozy and others



AGNs ZOO - hybrid radio galaxies

THERSYSTEL

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Giant sources in the ROUGE I catalog



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ROUGE I

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(Kozieł-Wierzbowska+, 2020) a catalog of radio sources associated with optical hosts with unresolved or extended morphologies

- The SDSS Main Galaxy sample (Strauss+, 2002)
- the Red Galaxy Sample (Eisenstein+, 2001)
- · FIRST (Becker+, 1995)
- NVSS (Condon+, 1998)

NVSS (black) & FIRST (red) contour images overload on SDSS grey-scale maps. The contours start from a 3σ confidence level and are scaled of $\sqrt{2}$. Negative 3σ contours are marked with dashed lines.

Giant sources in the ROUGE I catalog



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NTWI RSYTET

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> Proctor (2016): faint elliptical galaxy, LAS 668'

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1st host) z_{phot} =0.782 \rightarrow 5.0 Mpc 2nd host) z=0.115 \rightarrow 1.5 Mpc

Giant sources in the ROUGE I catalog



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Possibly DDRG One-sided jet-like structure LAS 578'

THERSYSTEM

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1st host) z_{phot} =0.586 → 3.9 Mpc 2nd host) z_{phot} =0.657 → 4.9 Mpc 3rd host) z=0.068 → 0.9 Mpc

Galaxy cluster ZwCl 1550.5+2310



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NVSS & VLASS data

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(right) L-band NVSS contours of J155210.97+224508.0 (left) VLASS (Lacy+, 2020) contours of J155235.1+223418 Contours start from a 3σ confidence level and are scaled of $\sqrt{2}$. The negative 3σ contours are marked with dashed lines.

GMRT 325 MHz & ASKAP observations



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GMRT 325 MHz (left; the project code 32_117) & ASKAP 889 MHz (right; McConnell+, 2020) contour images of J155210.97+224508.0 and J155235.1+223418ASKAP. The contours start from a 3 σ confidence level and are scaled of $\sqrt{2}$. Negative 3 σ contours are marked with dashed lines.

LOFAR data

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0.05

0.04

0.03 peam

0.02

0.01

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LOFAR 144 MHz colorscale image of two GRGs candidates:

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J155210.97+224508.0 and **J155235.1+223418**. The fluxscale starts below the 0.06 Jy level to present very faint structures.

LOFAR - ASKAP spectral index

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α _{mean}	
N lobe	-0.90 (0.18)
S lobe	-0.77 (0.18)
E lobe	-0.91 (0.14)
W lobe	-0.84 (0.15)

α_{MAX}

J155210.97+224508.0 N lobe: -1.4.

 spectral index values divide N lobe area into 2 vertical areas

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- the bridge possesses value of α similar to the southern part of the N lobe

J155235.1+223418 W lobe: -1.5

• one point-like component located at RA=15 52 41.53, Dec=22 34 55.39 with steep $\alpha_{MEAN} \cong -1.1$

Central regions of both sources possess $\alpha_{_{\text{MEAN}}}\cong$ –0.8, typical for synchrotron emission

[S~ν^α]

LOFAR–ASKAP spectral index image of J155210.97+224508.0 and J155235.1+223418 with LOFAR convolved (25" ×25") contour plot. The contours start from a 3 σ confidence level and are scaled of $\sqrt{2}$. Negative 3 σ contours are marked with dashed line.

Pajdosz-Śmierciak+, in prep. to MNRAS

JVLA S- & C-band observations

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S-band JVLA data of J155210.97+224508.0 tapered at 5 k λ (big panel).

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C-band JVLA C-config. data of:

- A) the northern lobe of J155210.97+224508.0
- B) the southern lobe of J155210.97+224508.0
- C) the eastern lobe of J155235.1+223418
- D) the western lobe of J155235.1+223418

The contours start from a 3σ confidence level and are scaled of $\sqrt{2}$. Negative 3σ contours are marked with dashed lines.



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Spectral modeling

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Spectral (synchrotron) modeling parameters derived for different components of J155210.97+224508.0 & J155235.1+223418 SYNAGE (Murgia 1996) for JP (Jaffe & Perola 1973) & CI (Pacholczyk 1970) models

STRUCTURE	z	AGE [Myr]	MODEL
total	0,782	90-160	JP
total	0,115	700-740	JP
S lobe	0,782	25-45	JP
S lobe	0,115	200-210	JP
N lobe	0,782	130-230	JP
N lobe	0,115	1020-1070	JP
N lobe w/bridge	0,782	120-200	JP
N lobe w/bridge	0,115	910-950	JP

Conclusions, future work

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Crossing the size limit of radio sources with LOFAR

 \rightarrow the dedicated HBA LOFAR observations revealed extended lobes of both radio galaxies with high resolution (13."80 × 5."31) and sensitivity (0.2 mJy/beam) with backflow signatures and bridge-like structure

 \rightarrow these extended lobes are also detectable with the JVLA S-, and C-band observations

 \rightarrow the total extension (assuming the probable hosts) is 5 or 1.5 Mpc for J155210.97+224508.0 and (unfortunately) most likely 0.9 Mpc for J155235.1+223418

- → spectral index map revealed i.a. similarity of the bridge and the N lobe of J155210.97+224508.0
- \rightarrow SED model fits were done for different sources' areas: **assymetry in synchrotron ages** for the lobes pair of both sources (for J155210.97+224508.0 more reasonable seems z=0.782)
- → J155235.1+223418 morphology could be affected by the galaxy cluster environment restarting source?
- \rightarrow J155210.97+224508.0 assymetry or two separate sources aligned with each other?
- → SED model fits for different regions within J155210.97+224508.0 lobes to check the above hypotesis
- \rightarrow probably dynamical modeling to evaluate and compare the dynamical ages and physical parameters

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THANK YOU!

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