

Universität Hamburg
DER FORSCHUNG | DER LEHRE | DER BILDUNG



LOFAR

High-resolution **low-frequency** probes of X-ray emitting **knots in blazar jets**

Clues on **High-Energy Emission** from **Low-Frequency Radio** Observations

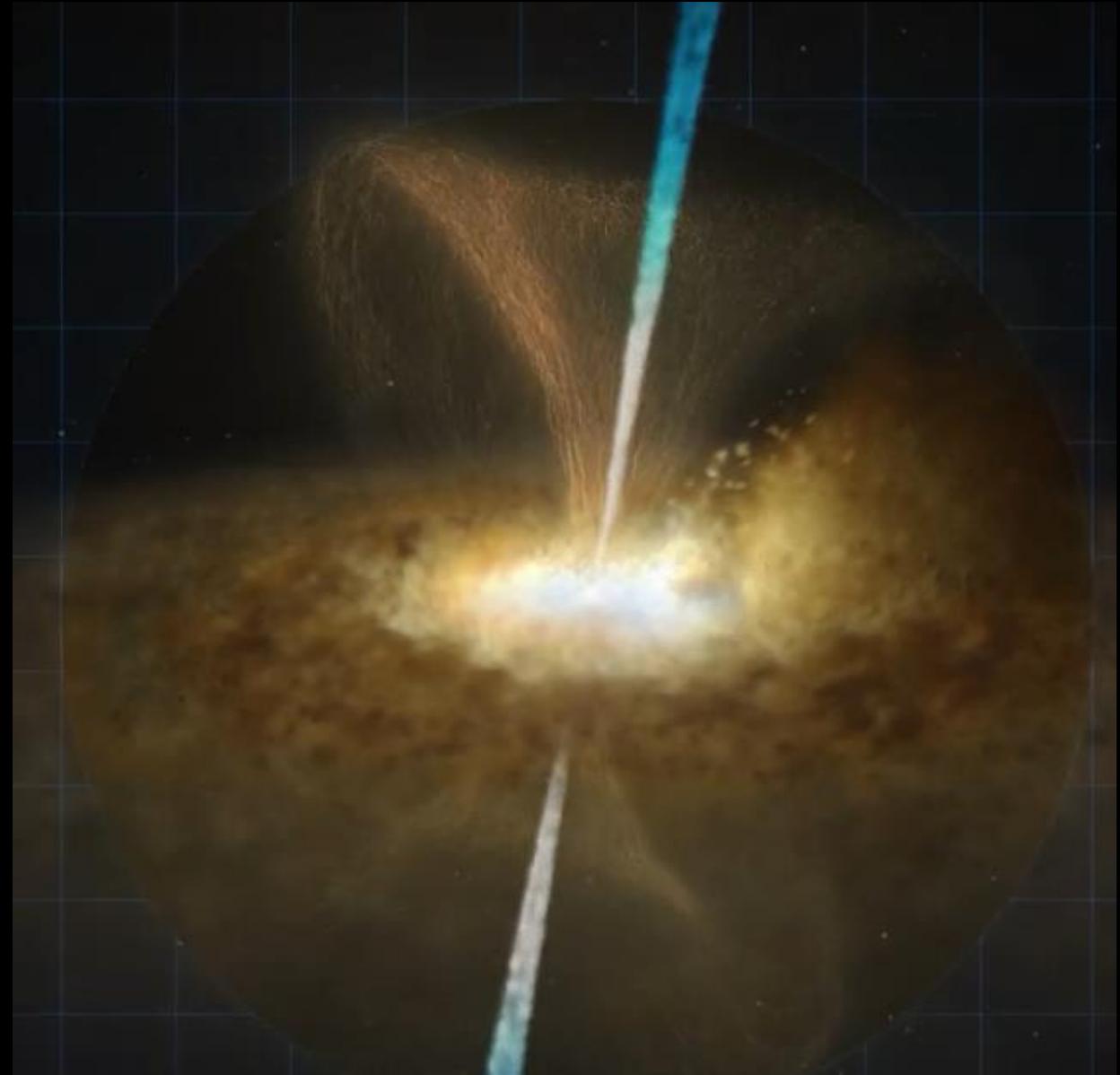
Hrishikesh Shetgaonkar

Etienne Bonnassieux , Matthias Kadler, Marcus Brüggem

LOFAR Family
Olsztyn, Poland
14th June 2023

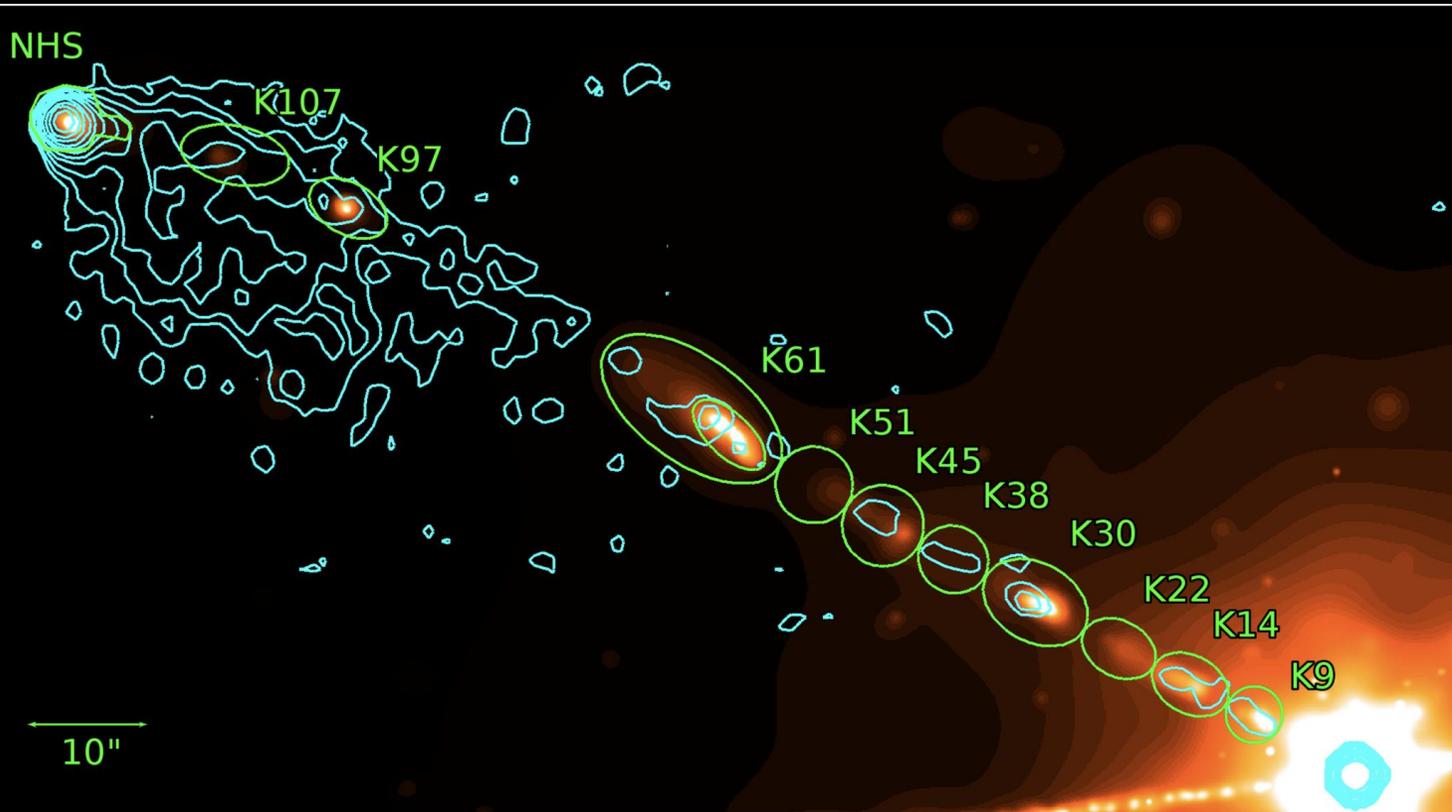
Blazars

- AGNs where jet is directed towards us
- Emit across the electromagnetic spectrum
- Relativistic beaming
- Particle accelerators
- Perfect laboratories for high-energy astrophysics!



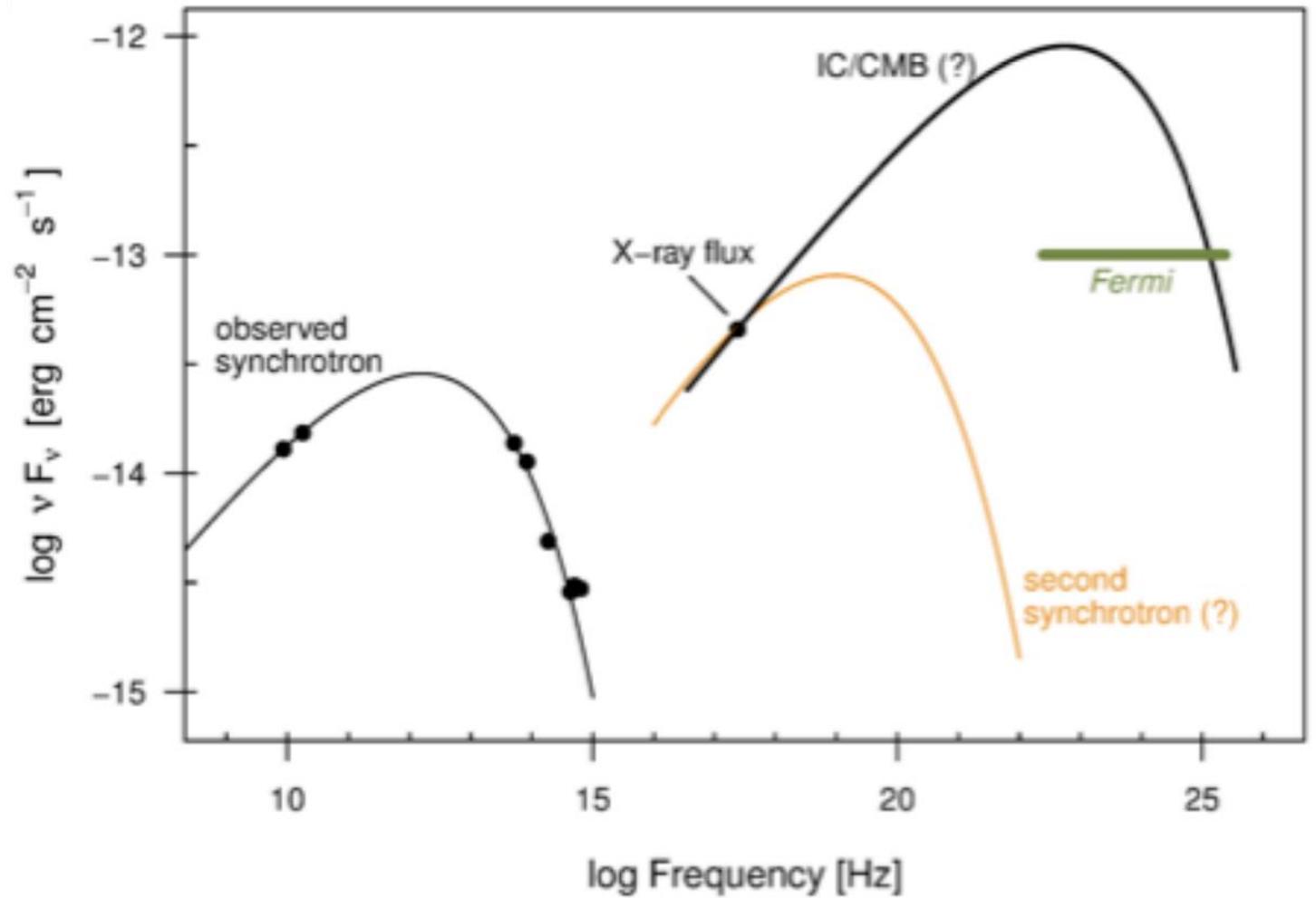
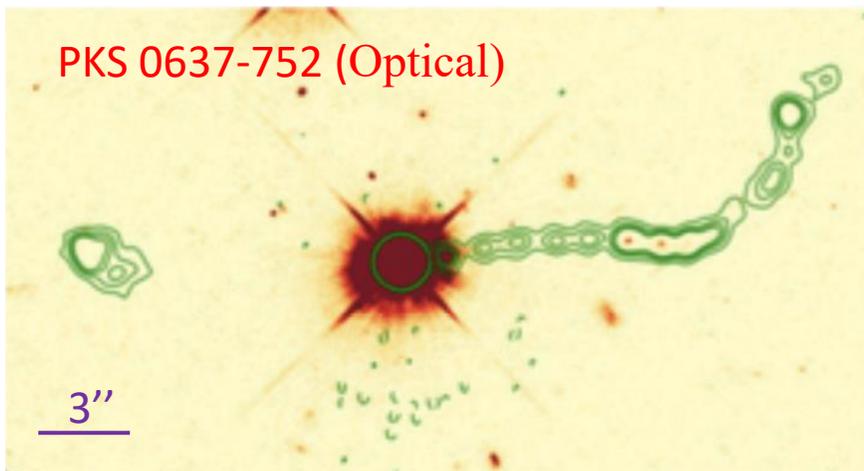
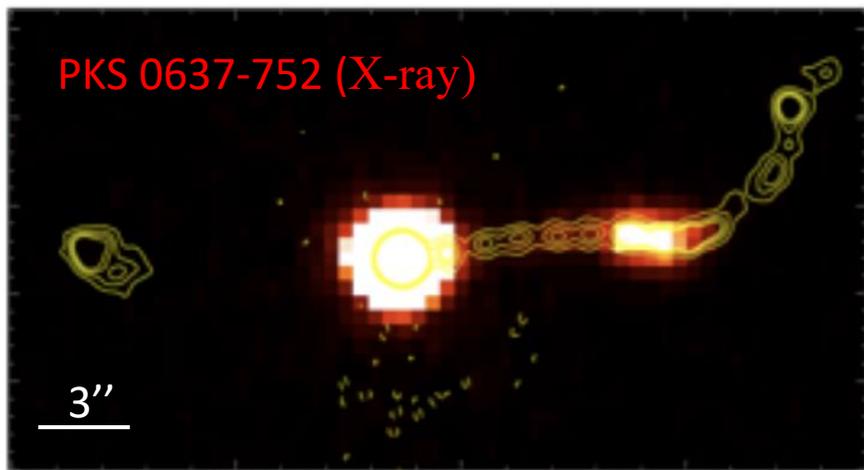
European Southern Observatory (ESO)

Knots in Jets



3C 111 in Chandra X-ray (heat scale), 8 GHz VLA (cyan contours)
Clautice et al., (2016)

- Compact bright regions in the jet
- Detection in X-rays -> emission processes?
- Re-acceleration mechanisms?
- Inverse scattering -> Photon field?
- High angular resolution necessary to separate the knots from core

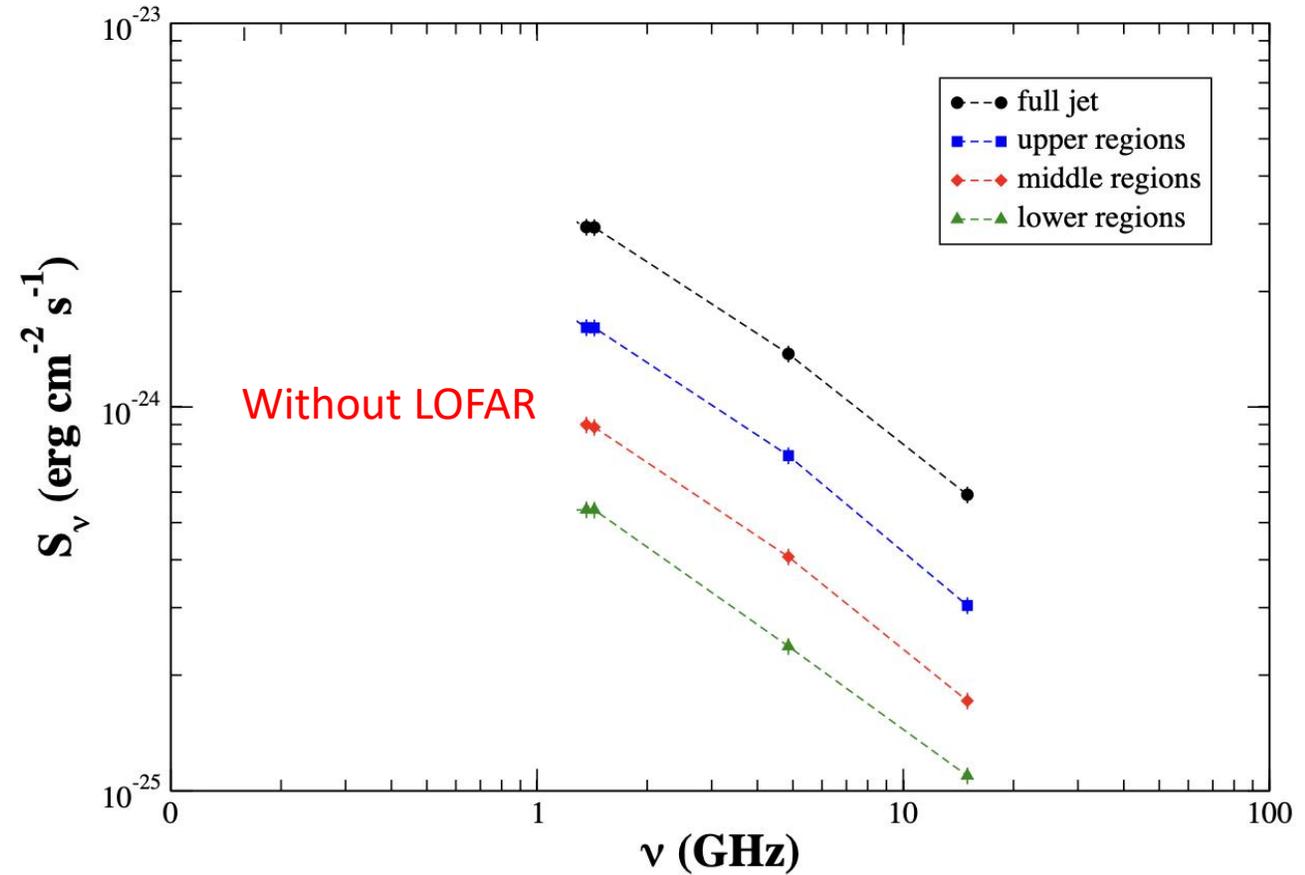
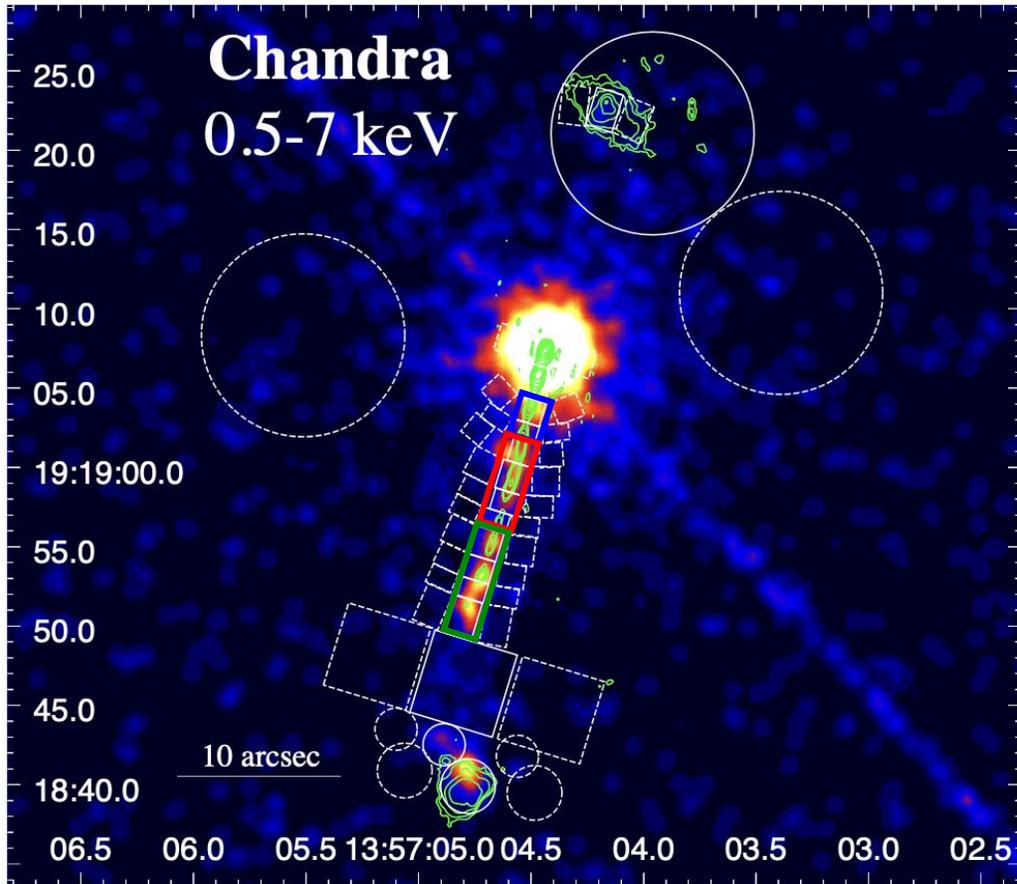


Left: PKS 0637-752 in X-ray (top), optical (bottom) with 17 GHz radio contours

Right: Radio to X-ray SED

Perlman et al. (2019)

Why LOFAR-VLBI?

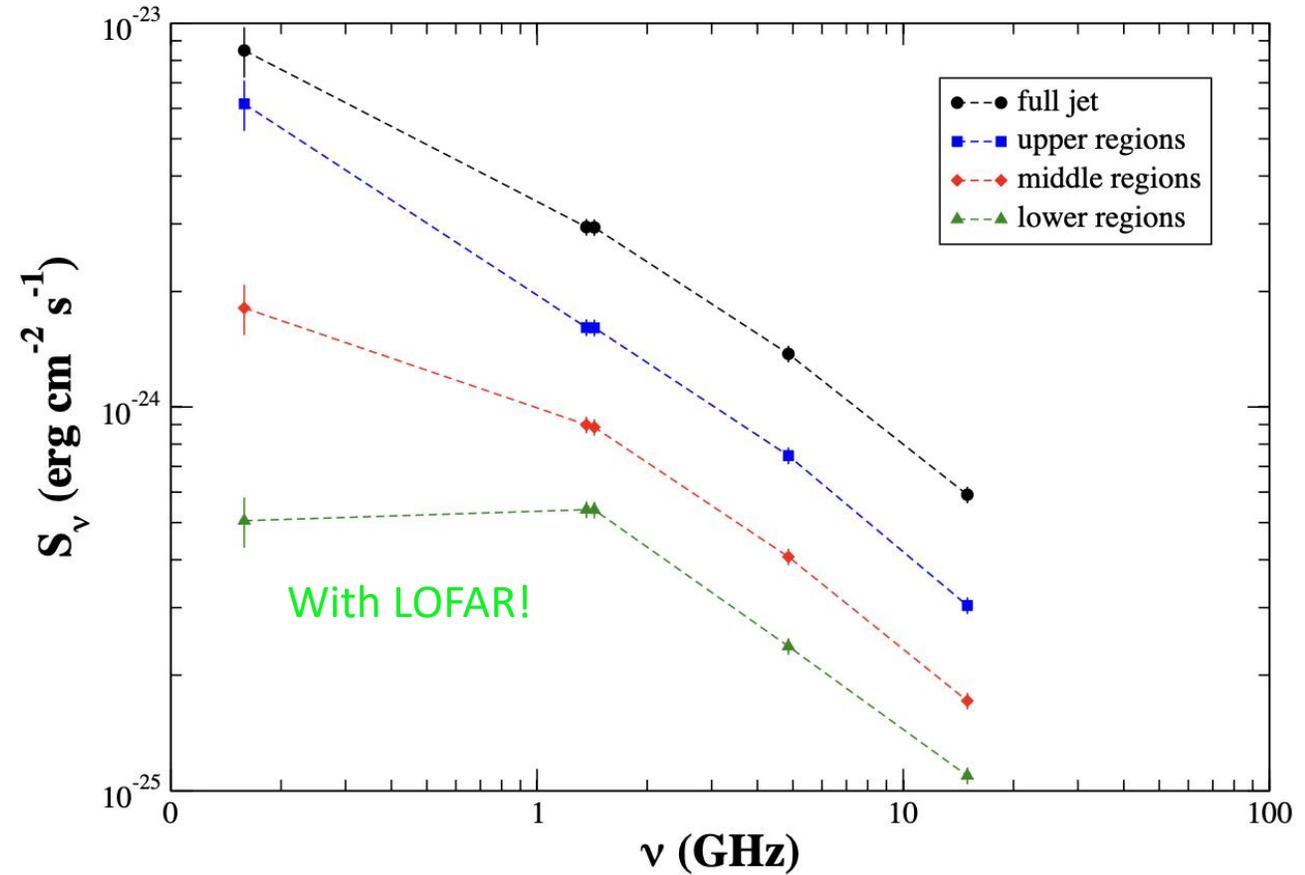
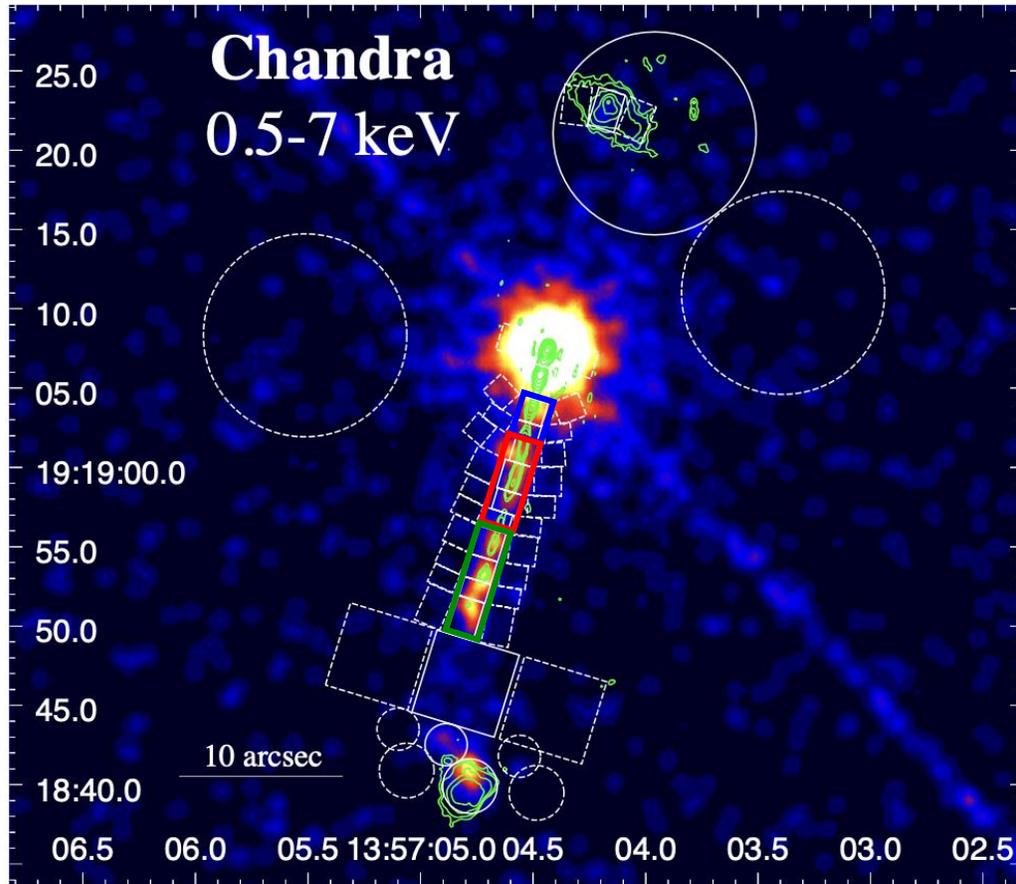


Left : 4C +19.44 in Chandra X-Ray with LOFAR radio contours overlaid

Right: MHz to GHz radio spectra of the jet.

Adapted from Harris et al., (2019)

Why LOFAR-VLBI?



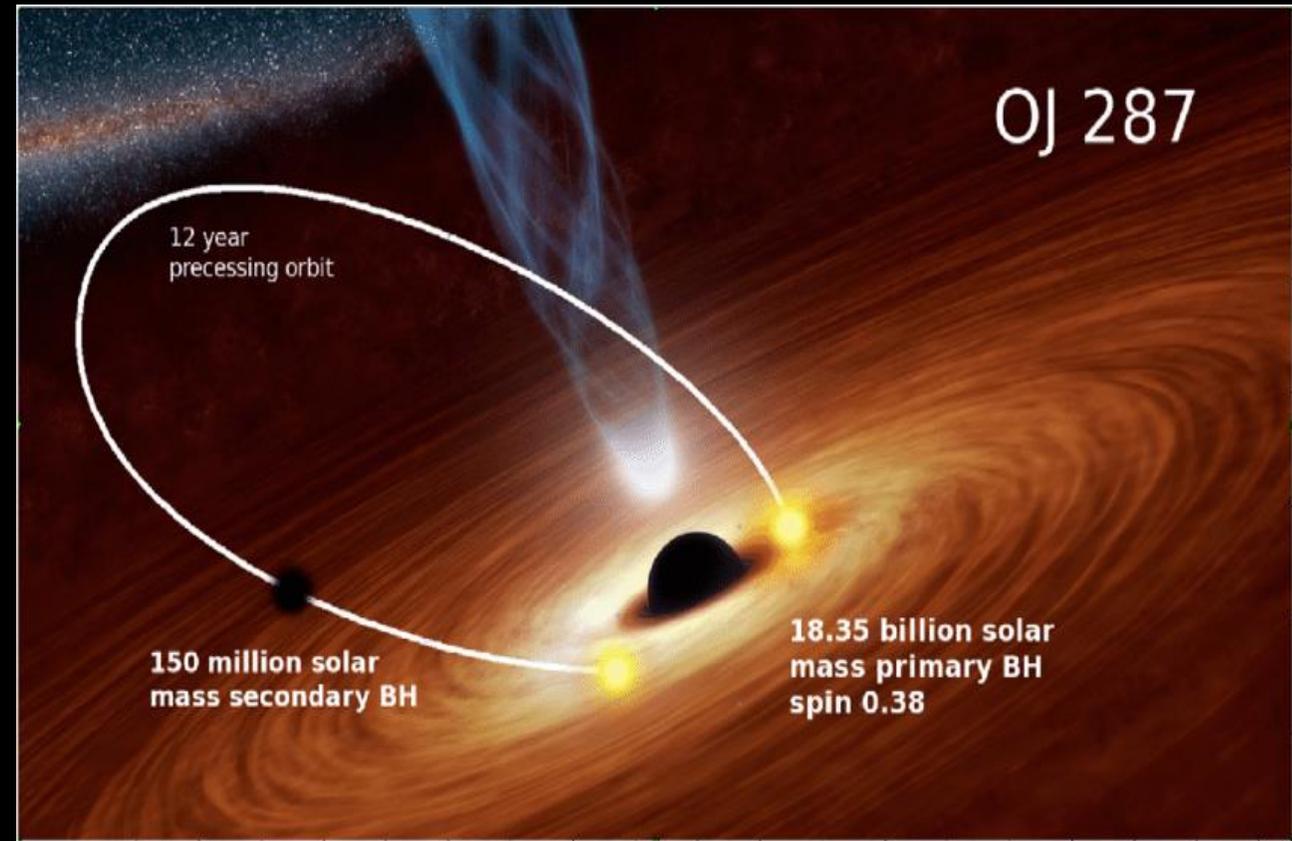
Left : 4C +19.44 in Chandra X-Ray with LOFAR radio contours overlaid

Right: MHz to GHz radio spectra of the jet.

Adapted from Harris et al., (2019)

Pilot Dataset

- Target: **OJ287**
- **8 hour** observation
 - 13 June 2019
 - 13 international stations
- **5 TB** of data!
- PI: Sean Mooney

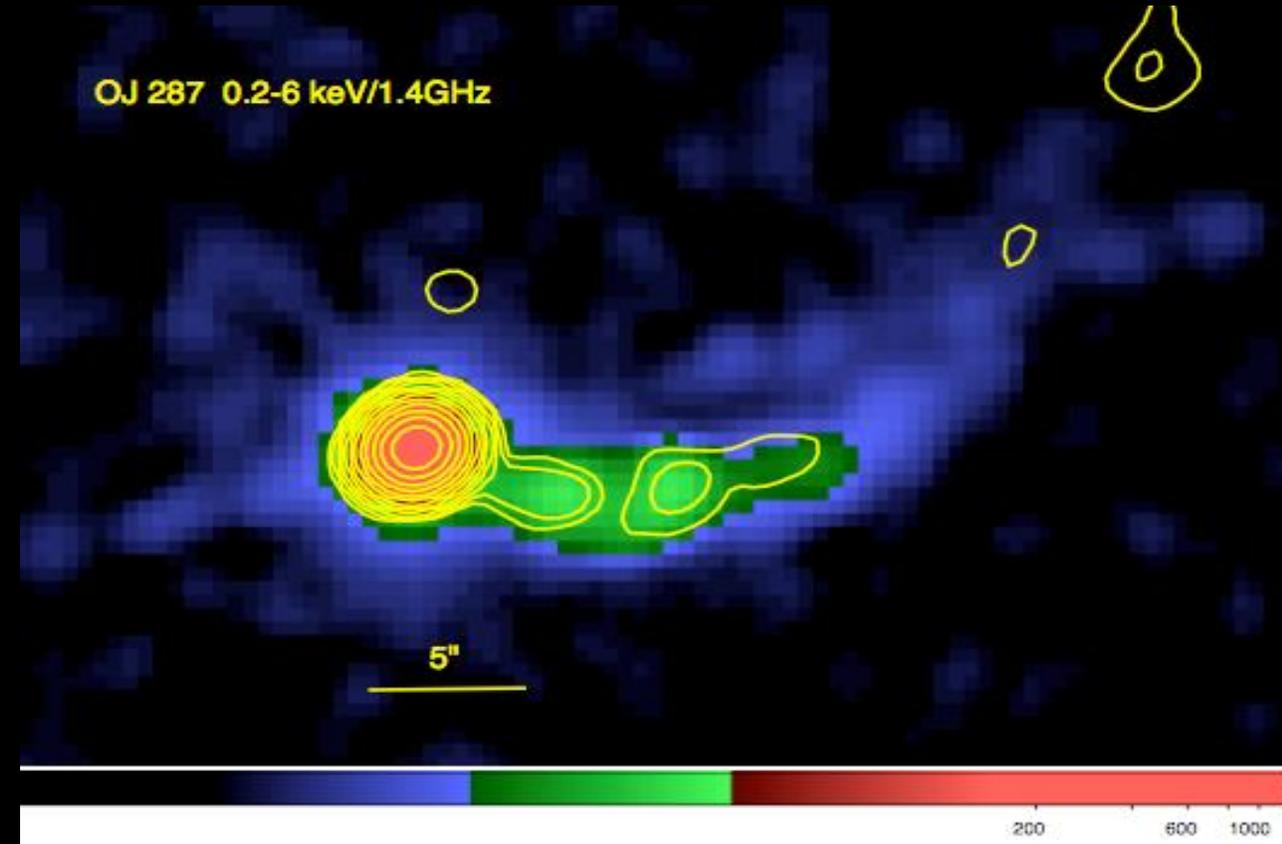


Artist impression

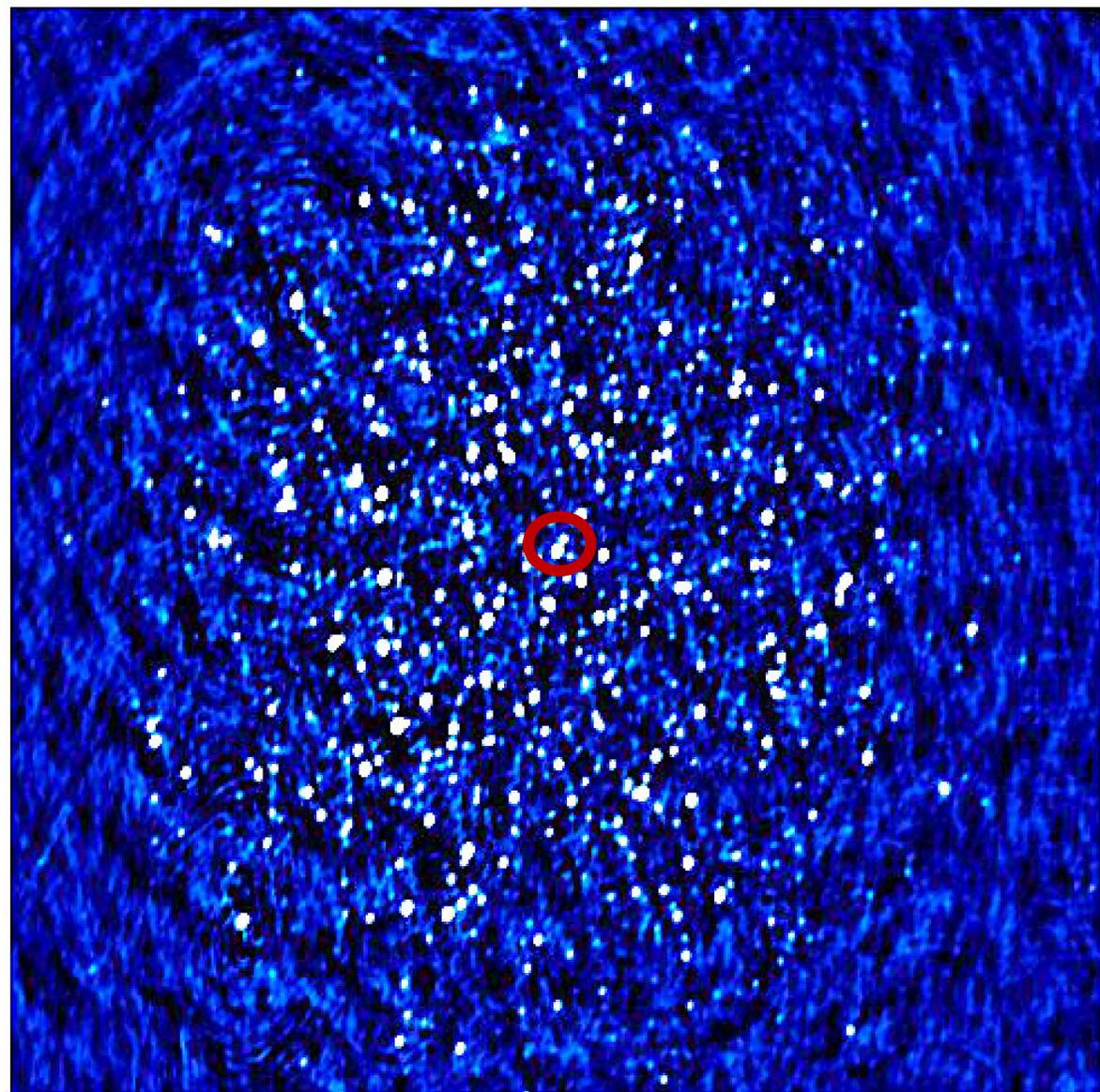
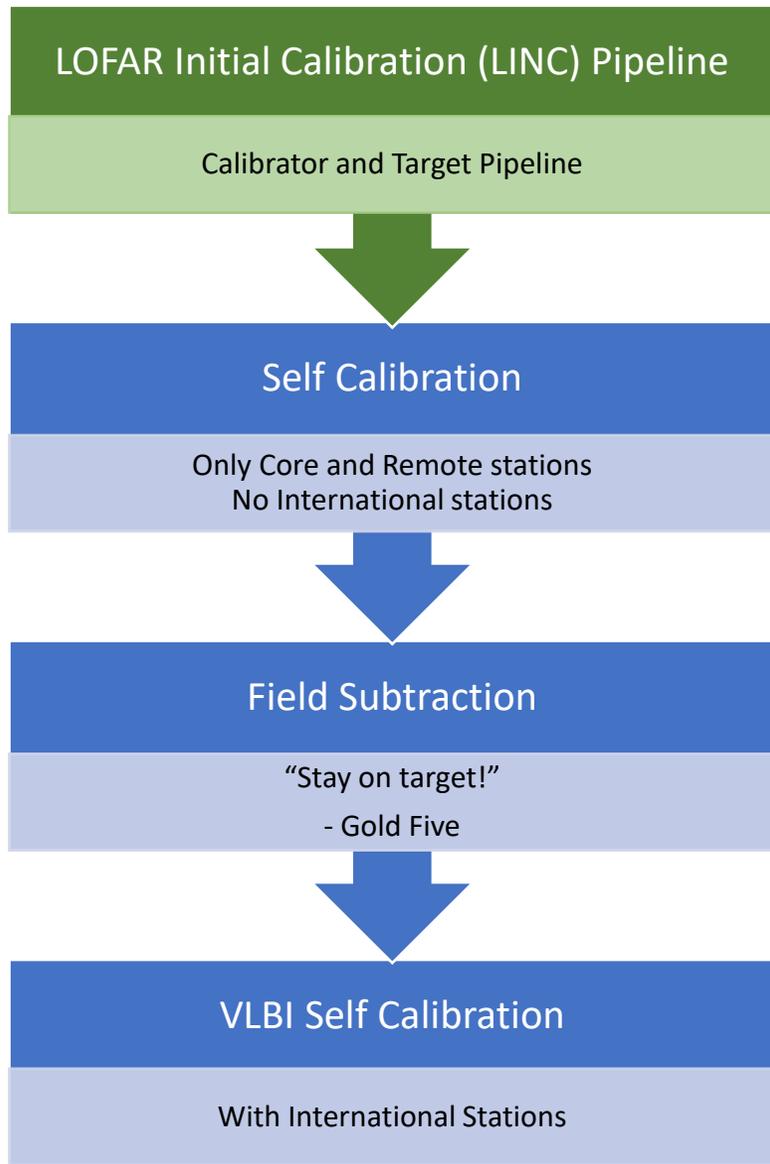
<https://www.scientificeuropean.co.uk/sciences/space/flares-from-the-supermassive-binary-black-hole-oj-287-put-constraint-on-the-no-hair-theorem/>

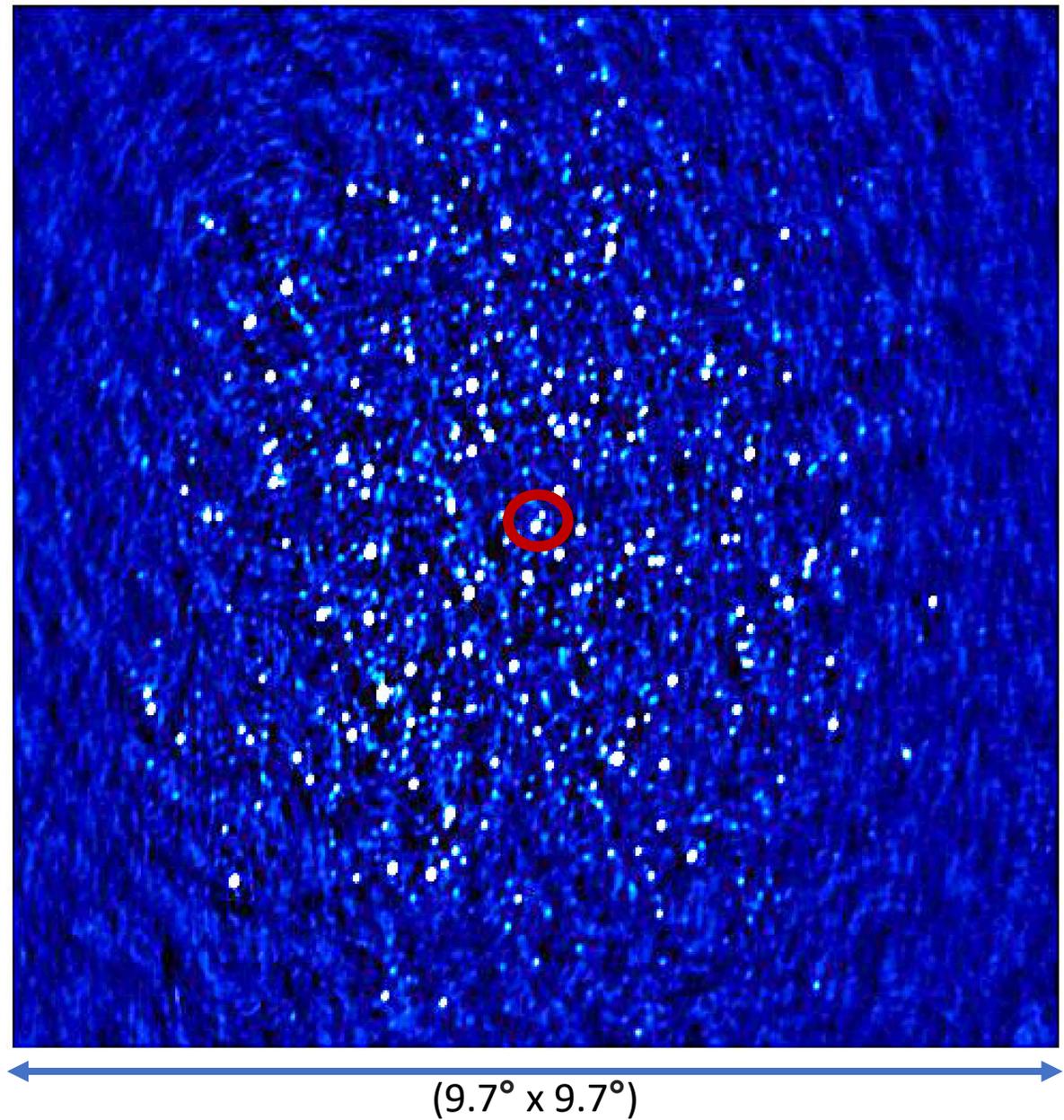
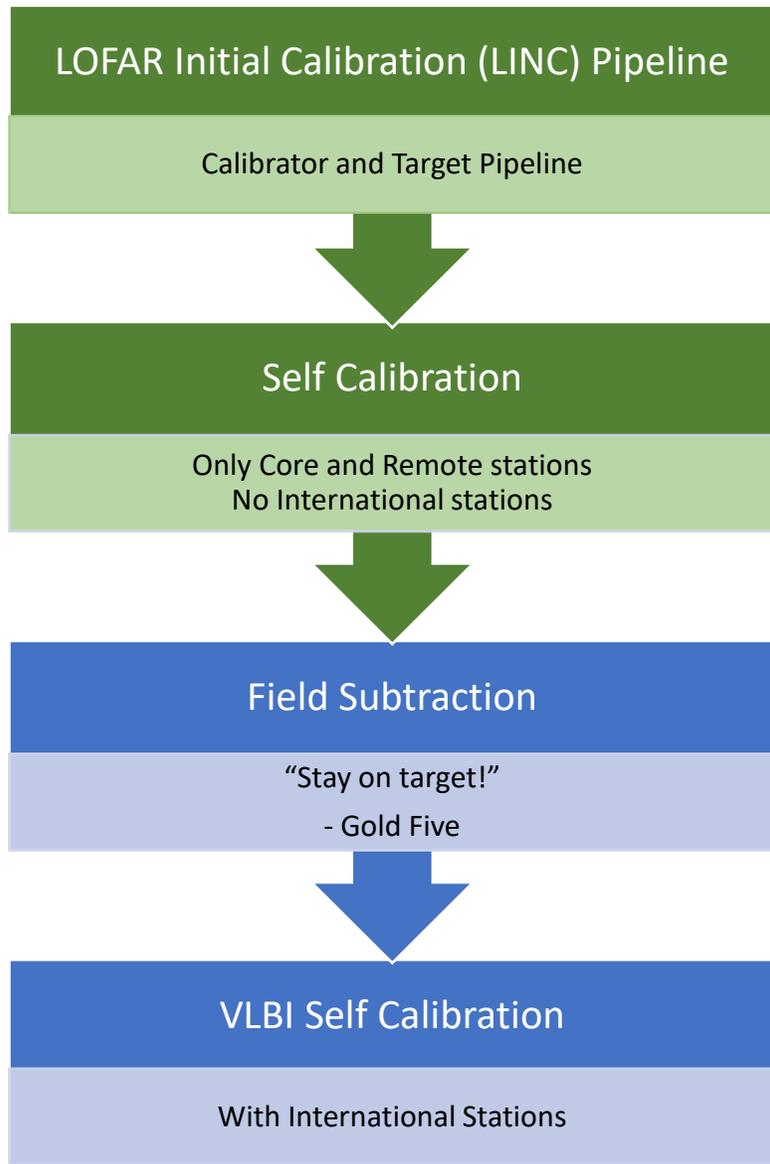
Pilot Dataset

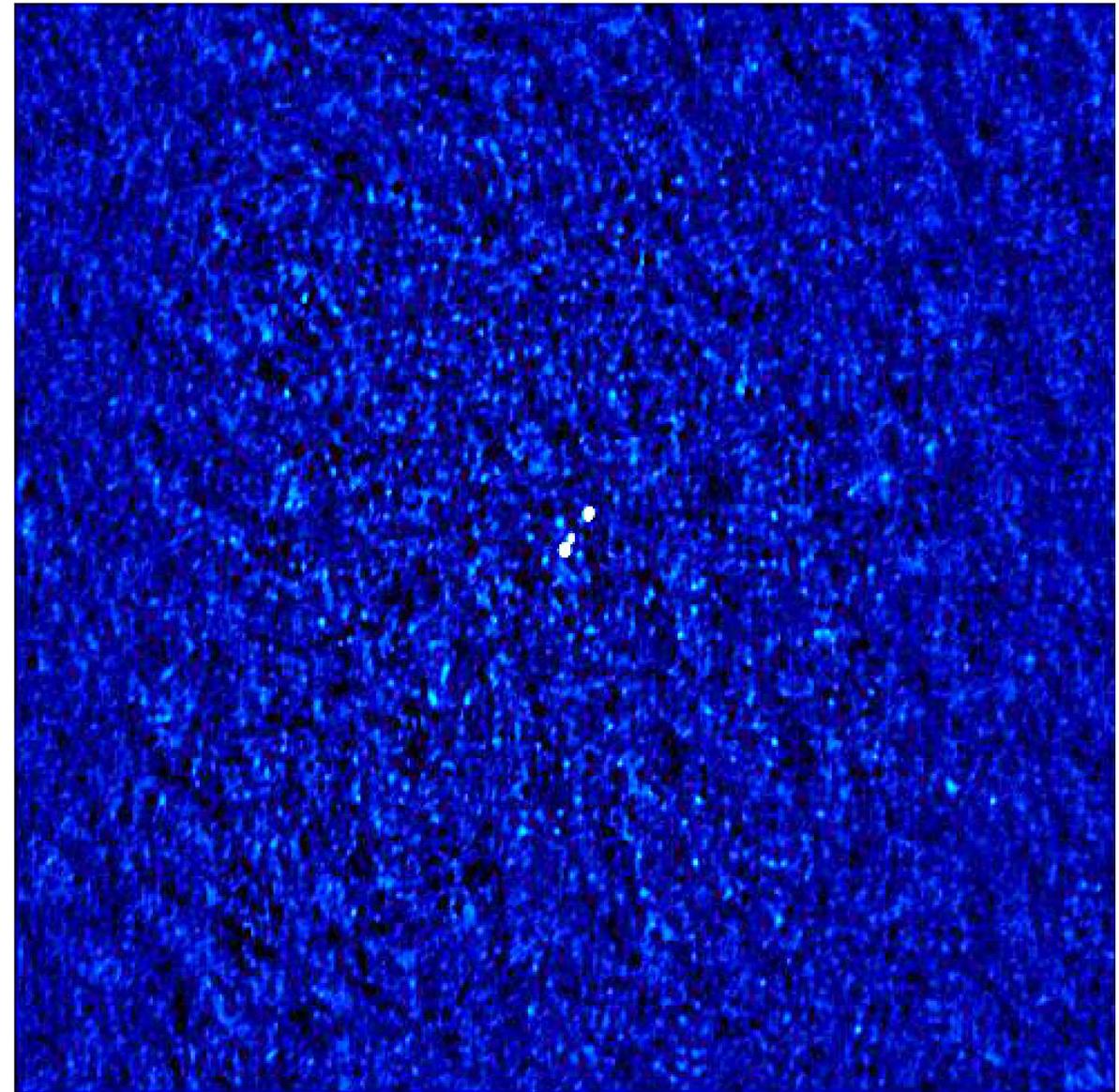
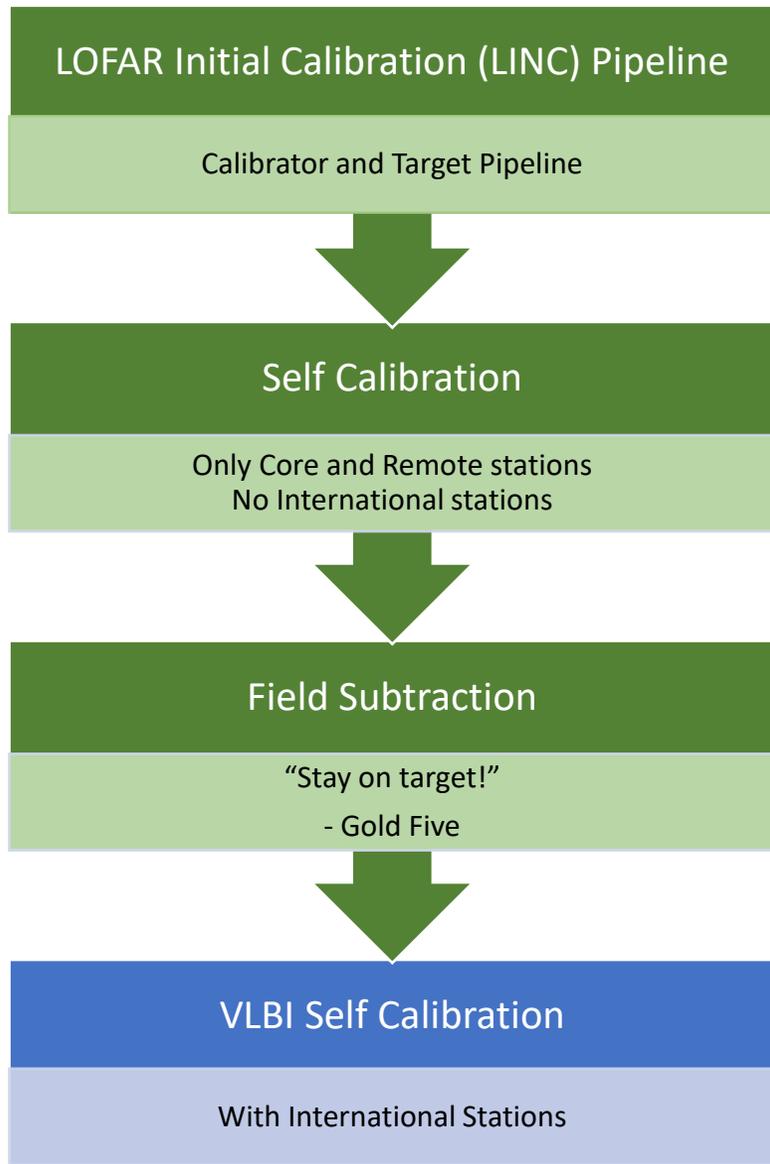
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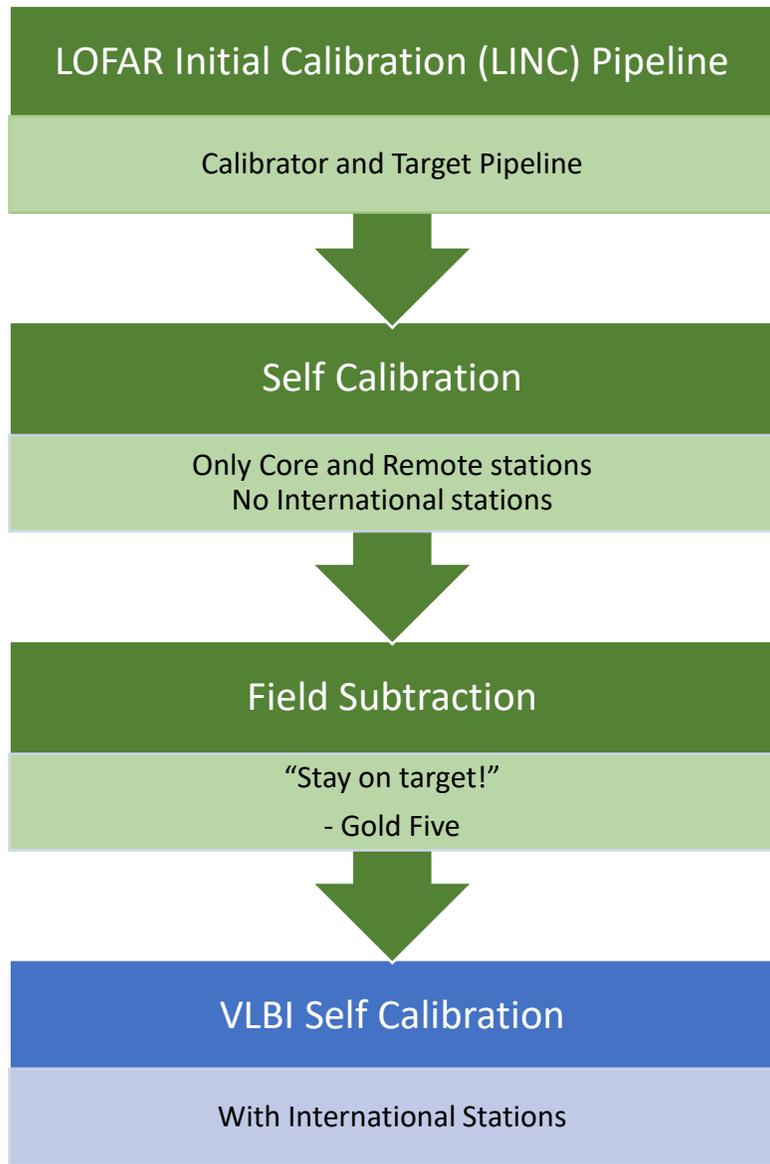
OJ287 in Chandra X-ray with VLA contours
Marscher, A.P., Jorstad, S.G. (2011)

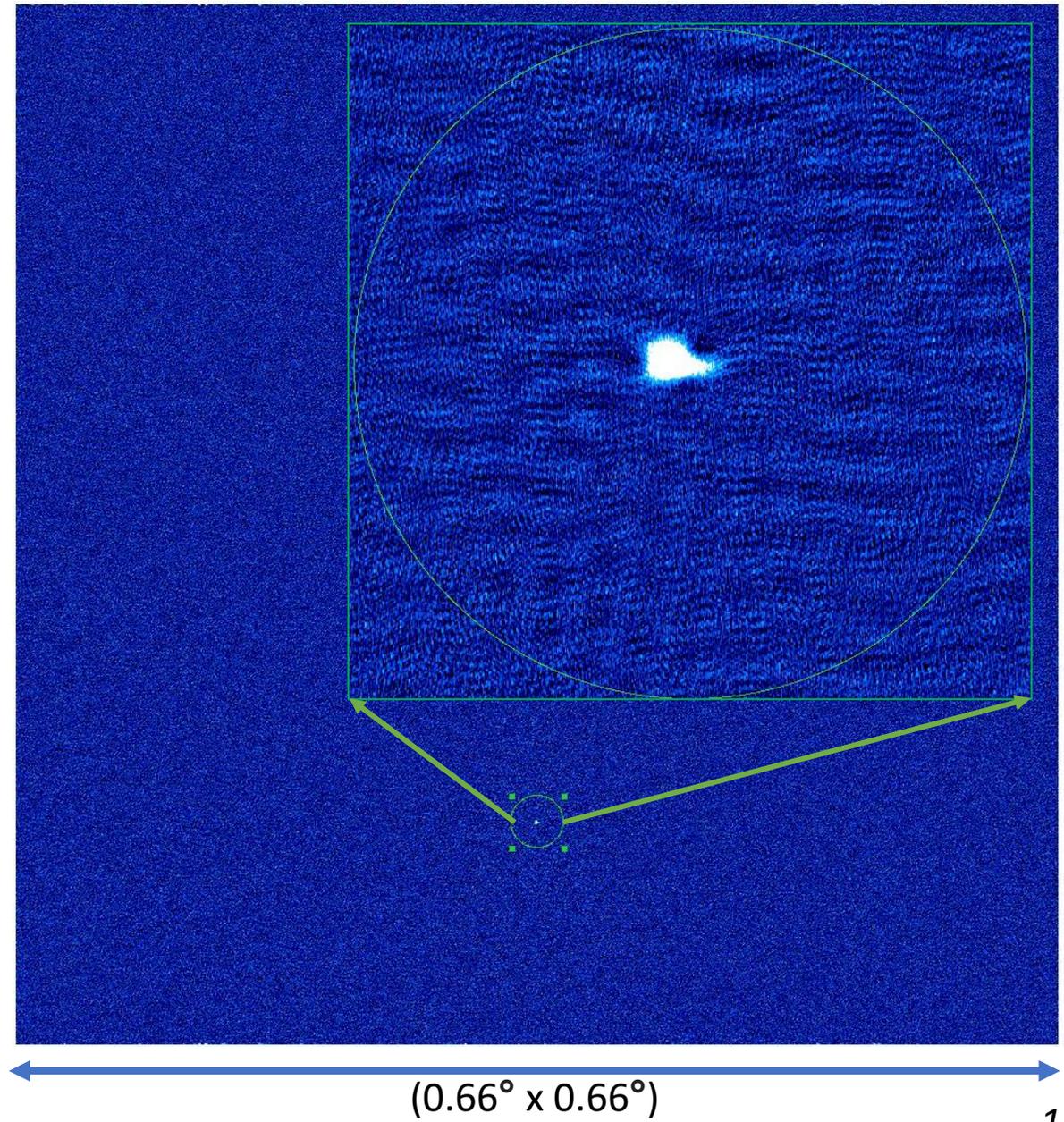
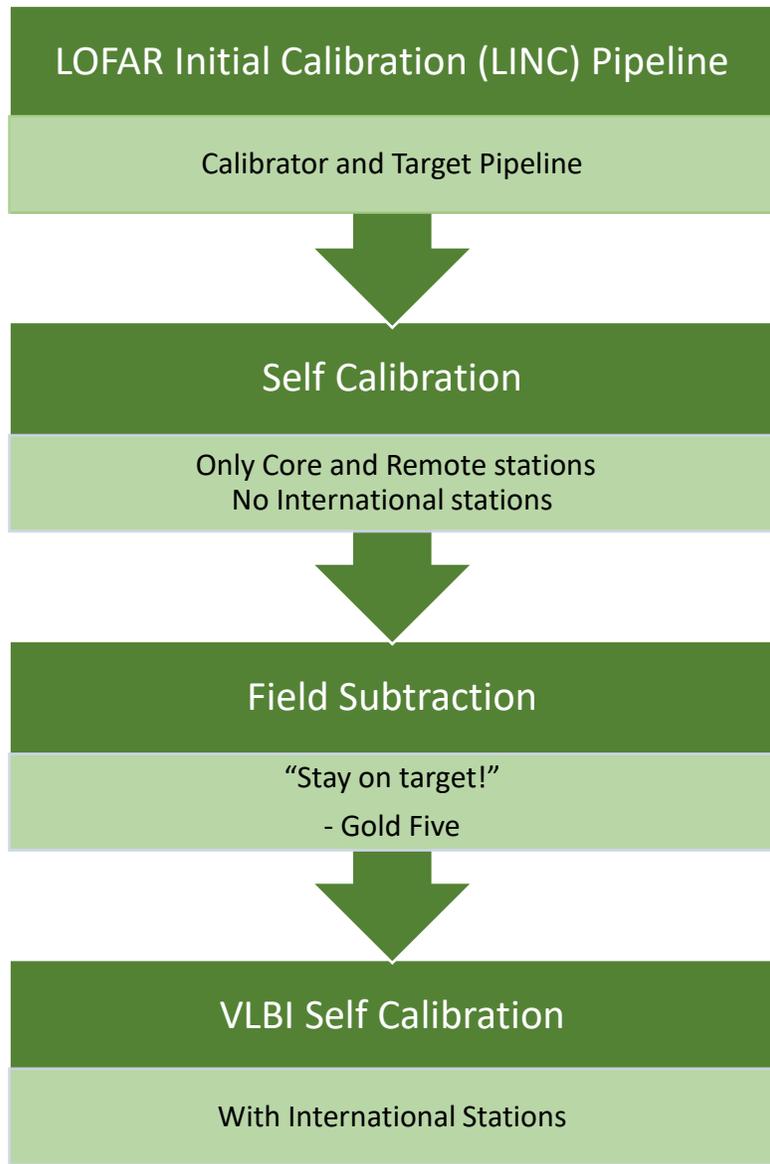




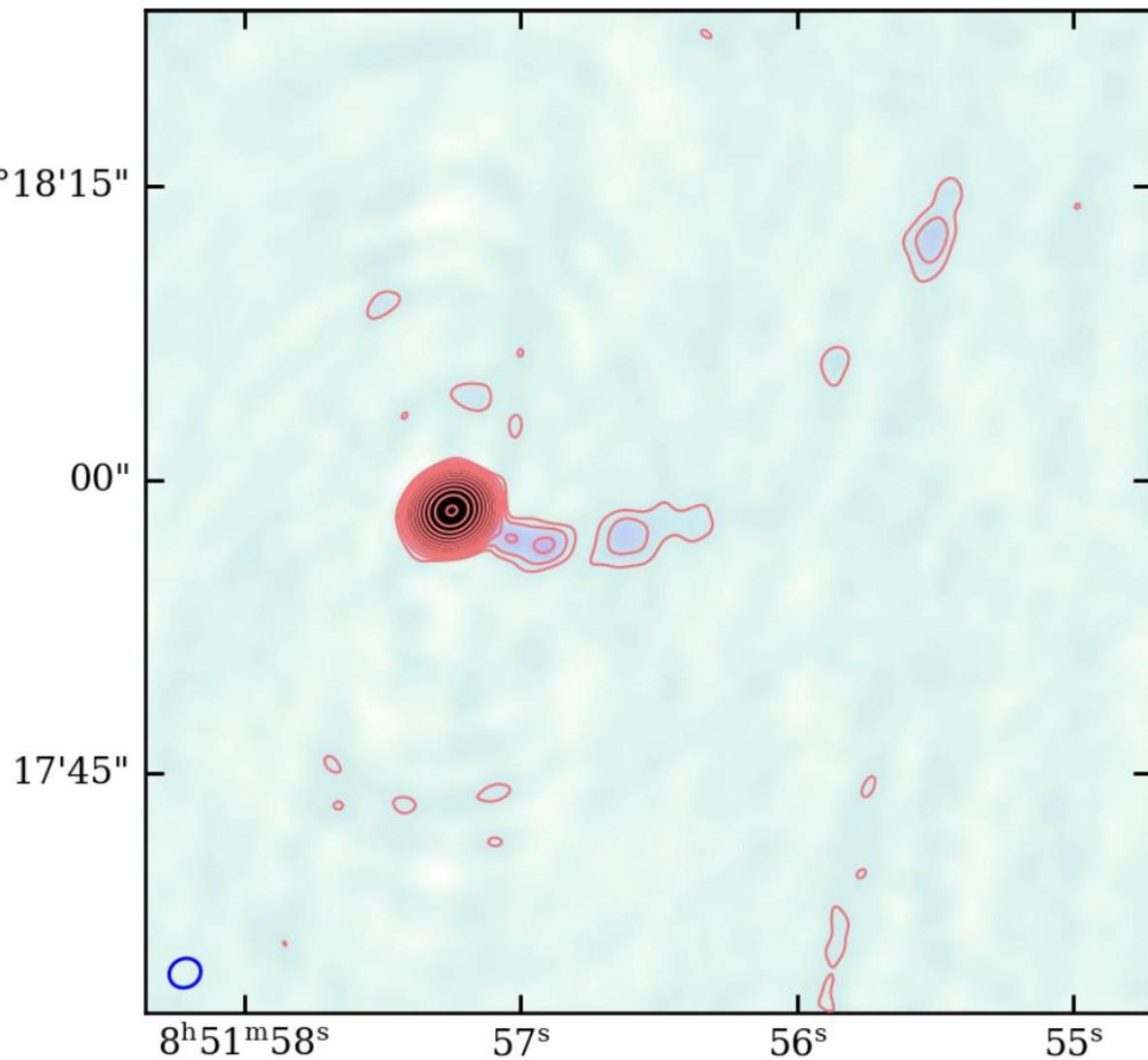


(9.7° x 9.7°)



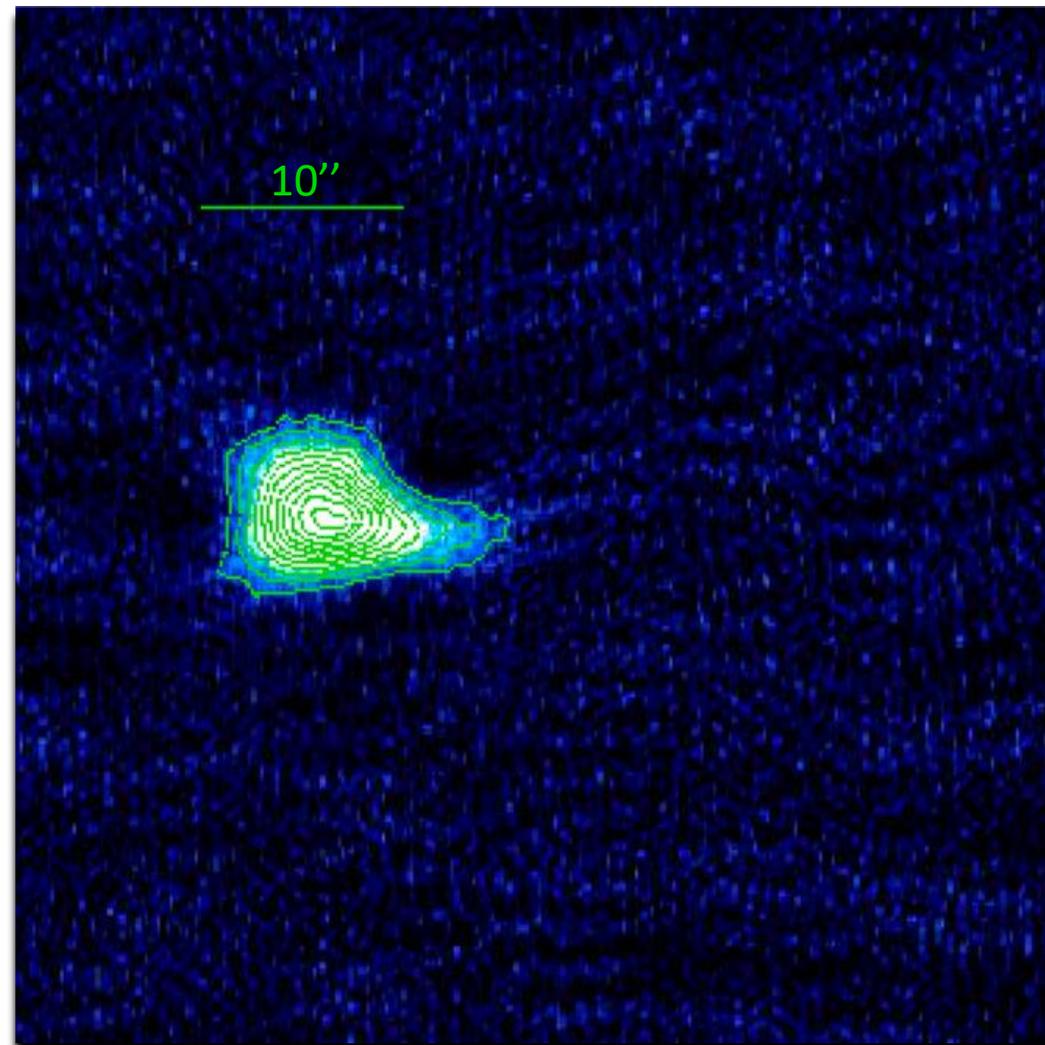


VLA



OJ 287 at 1.424 GHz
Mooney S. PhD Thesis (2020)

LOFAR

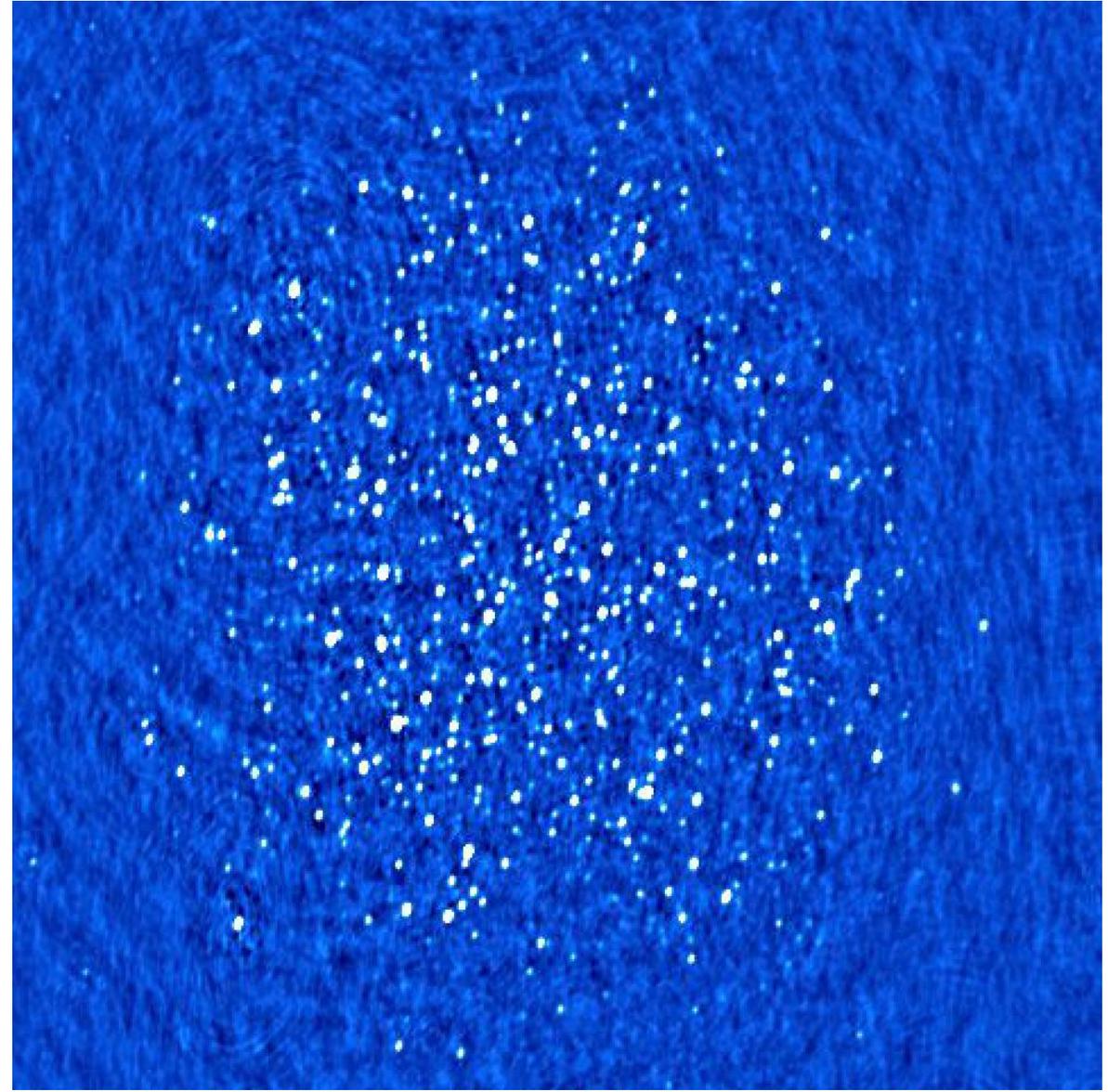


After one self-calibration pass,
With International Stations

10 Sub-bands

Low Sensitivity

Full Band



(9.7° x 9.7°)

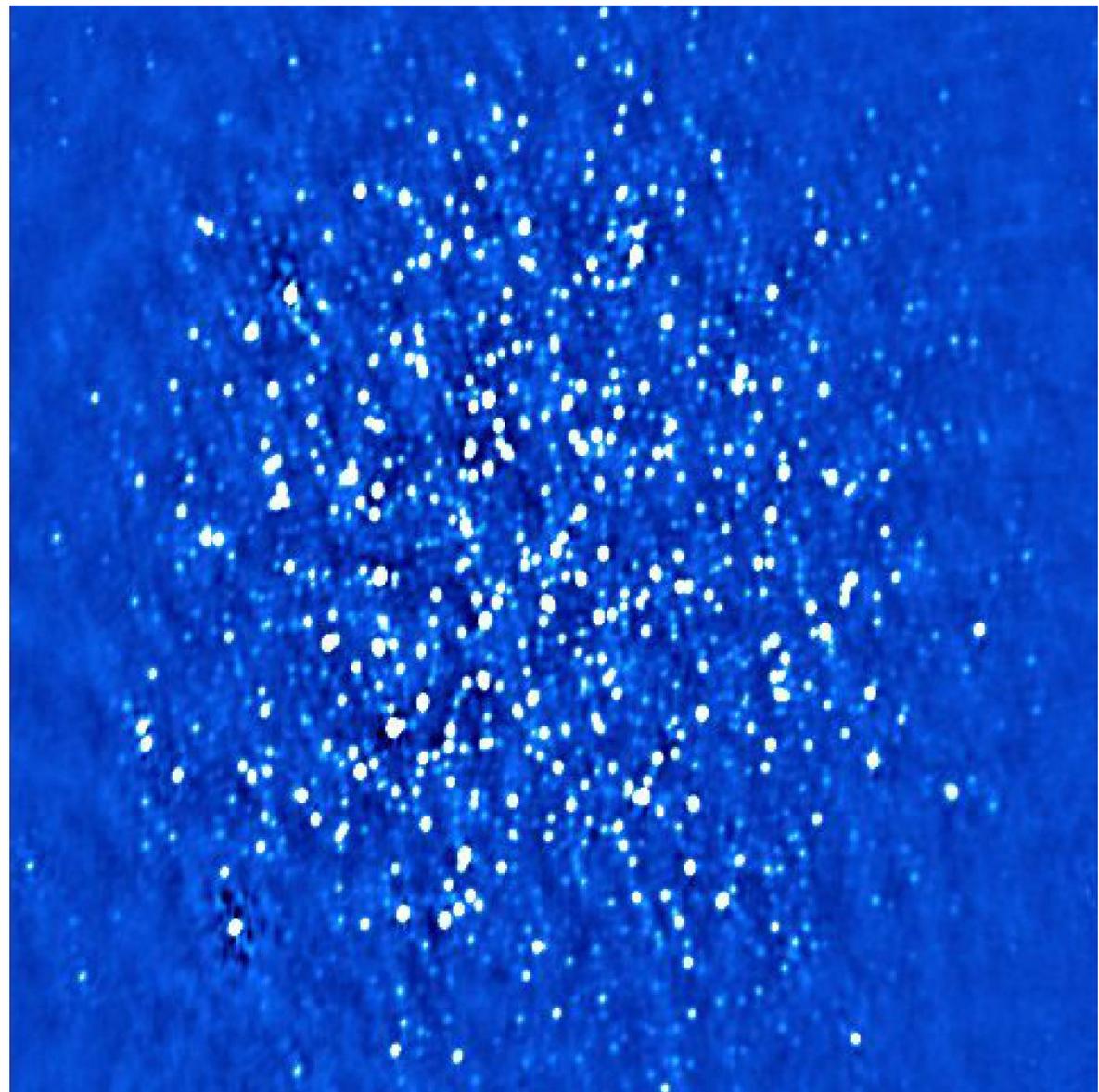
10 Sub-bands

Low Sensitivity

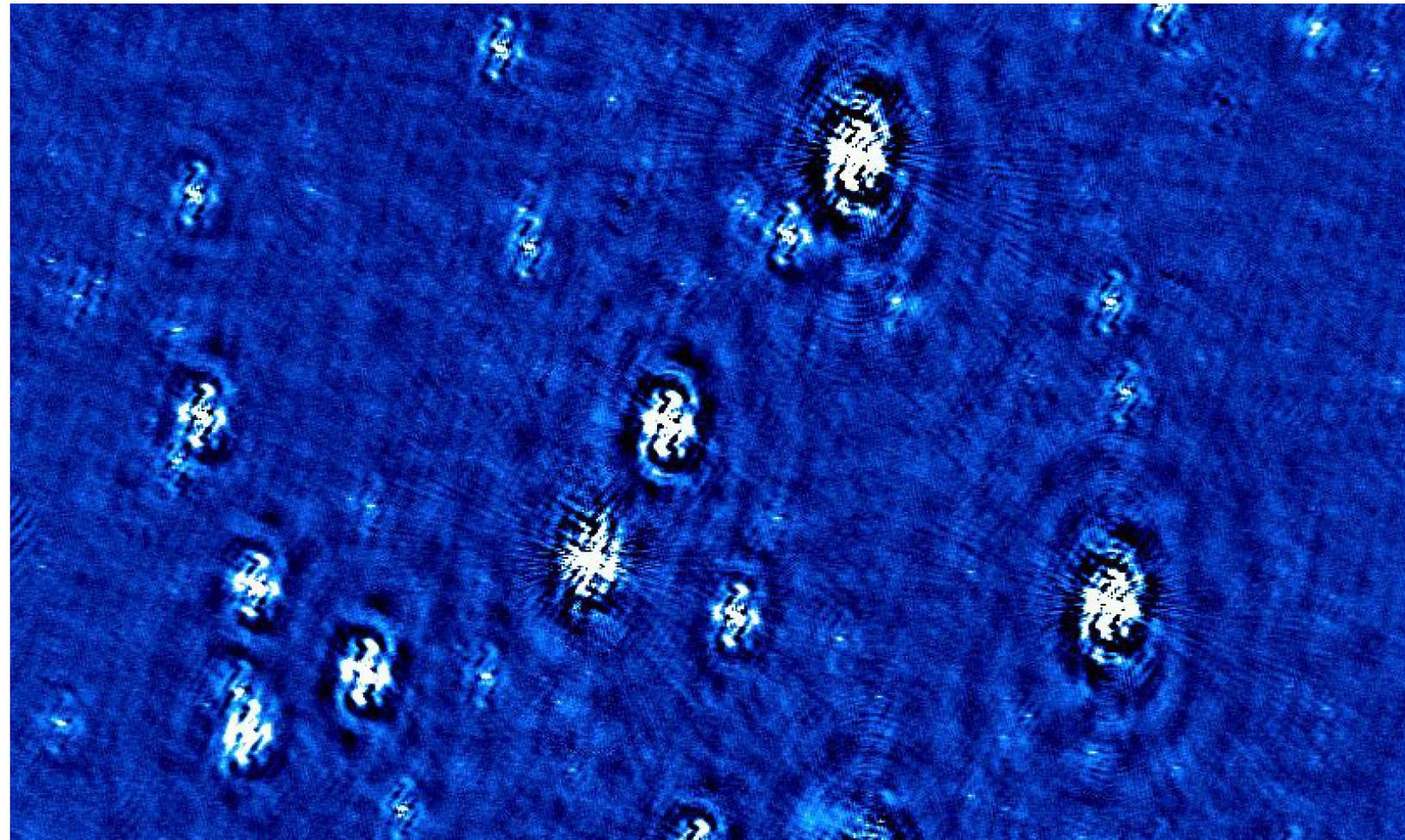
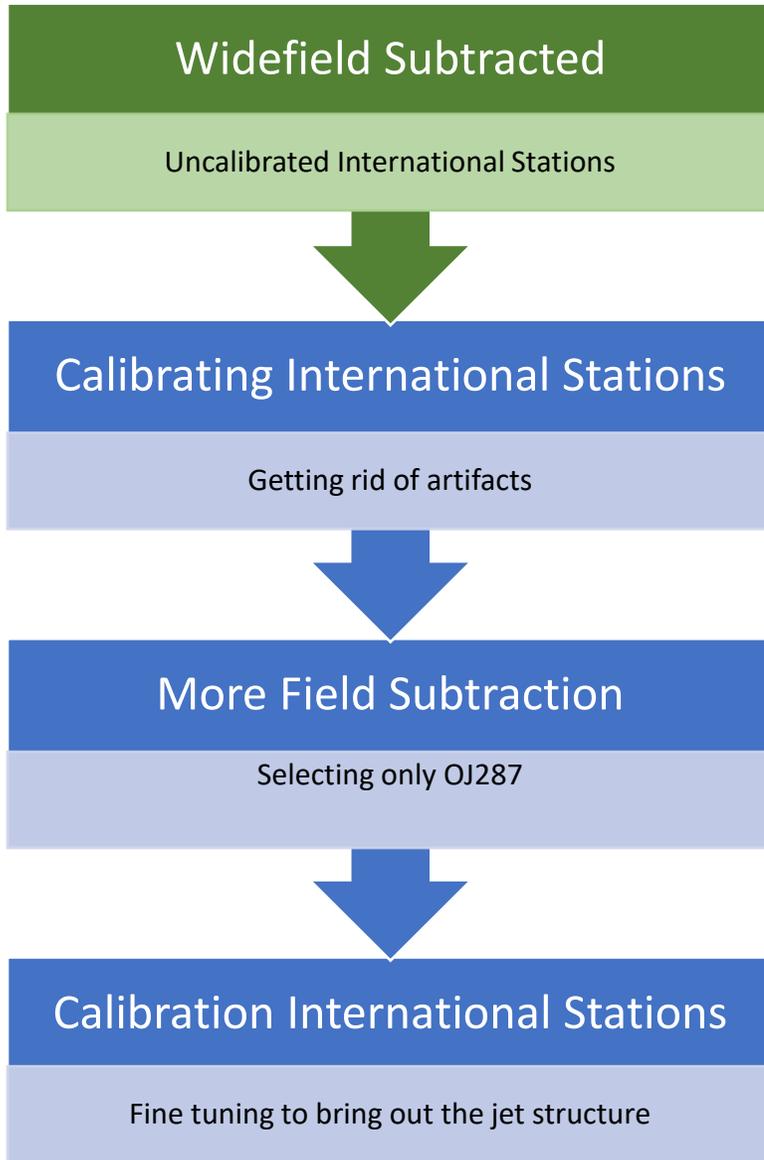
Full Band

Higher Sensitivity

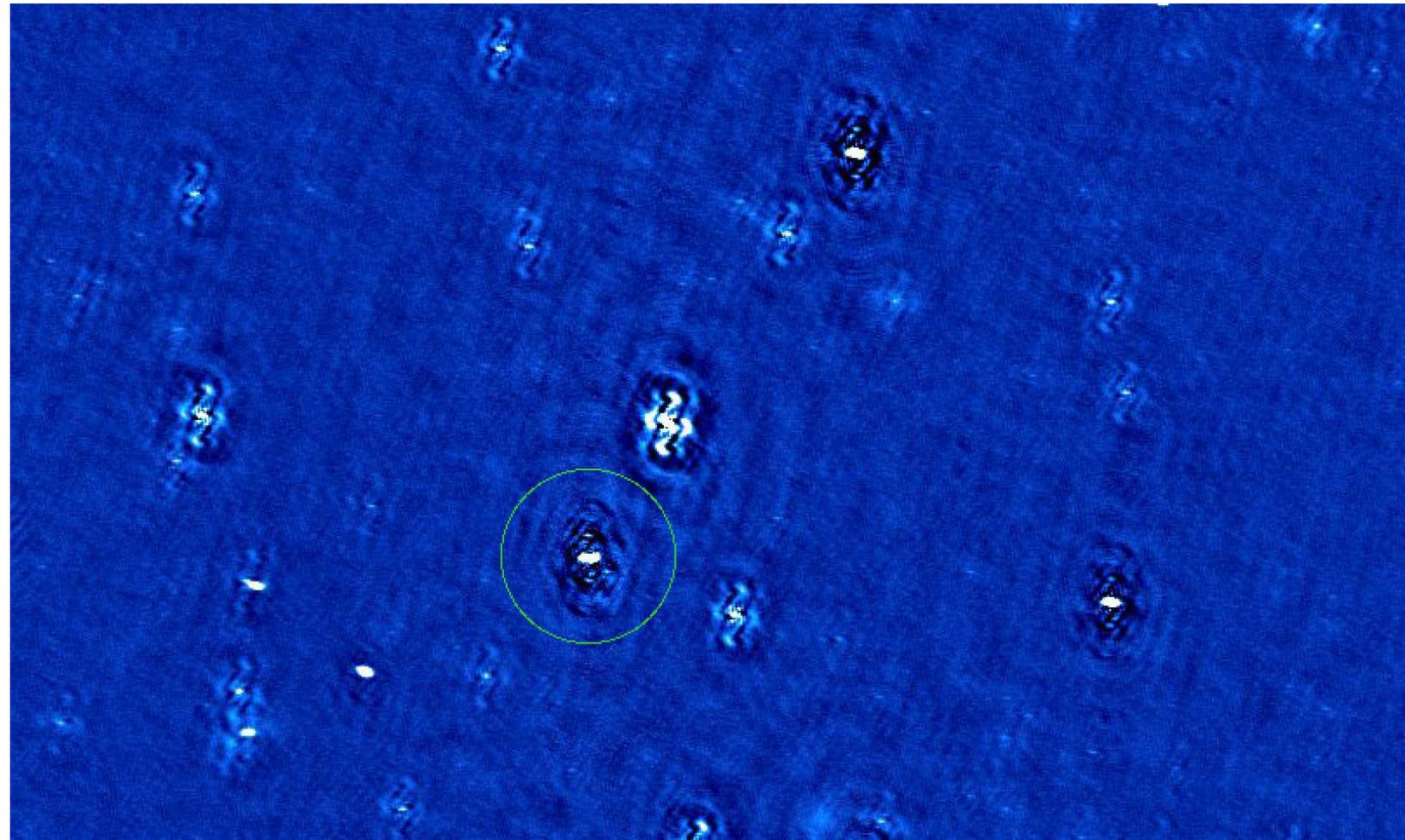
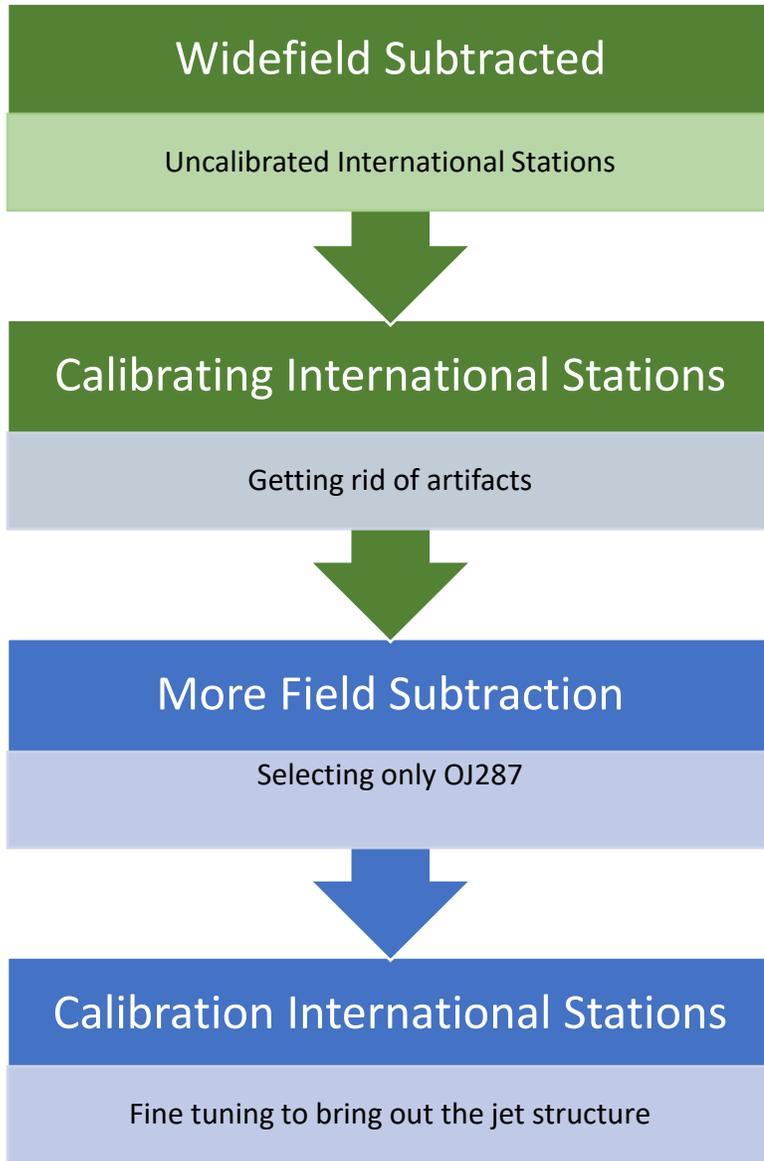
“More Data, More Problems”



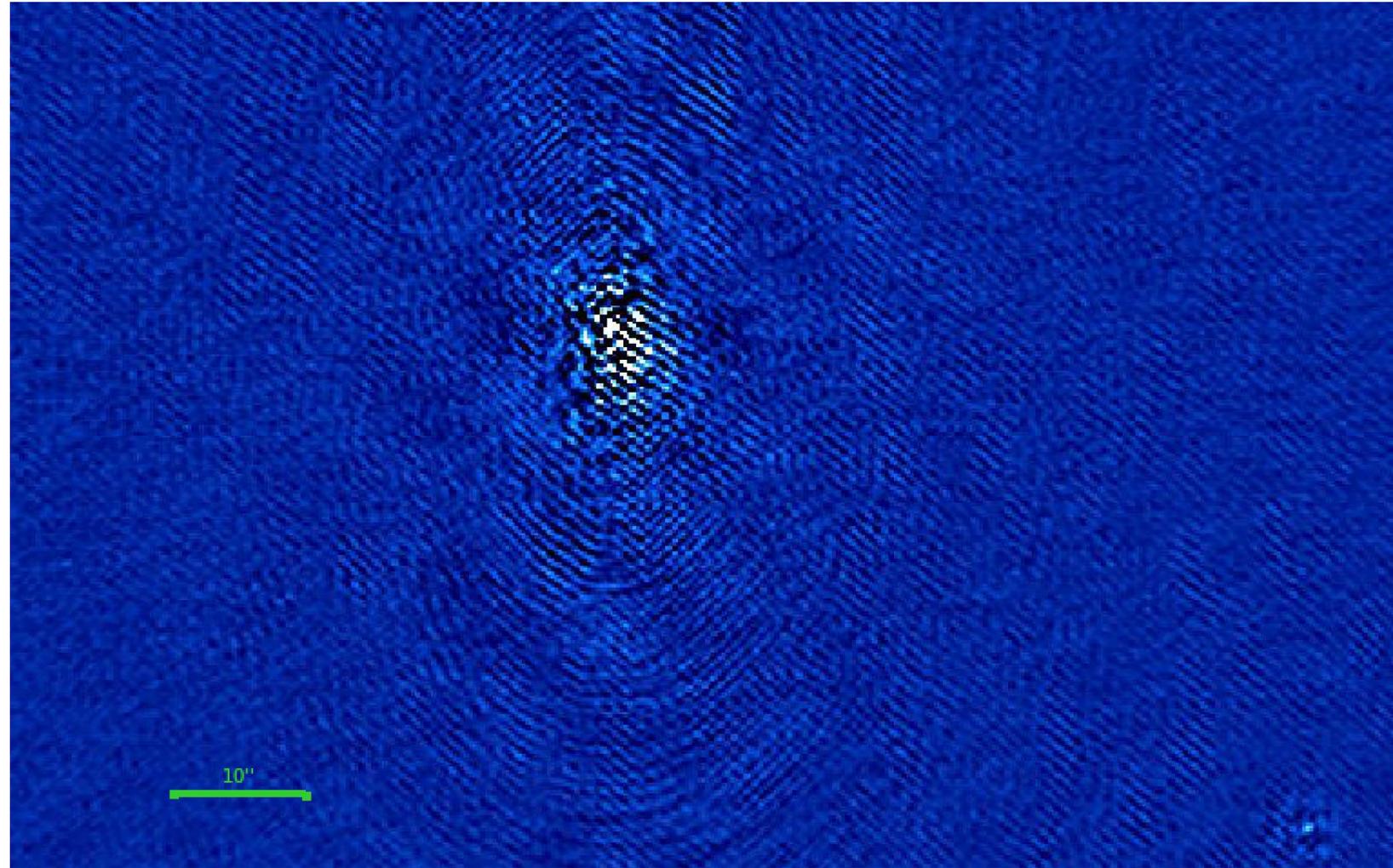
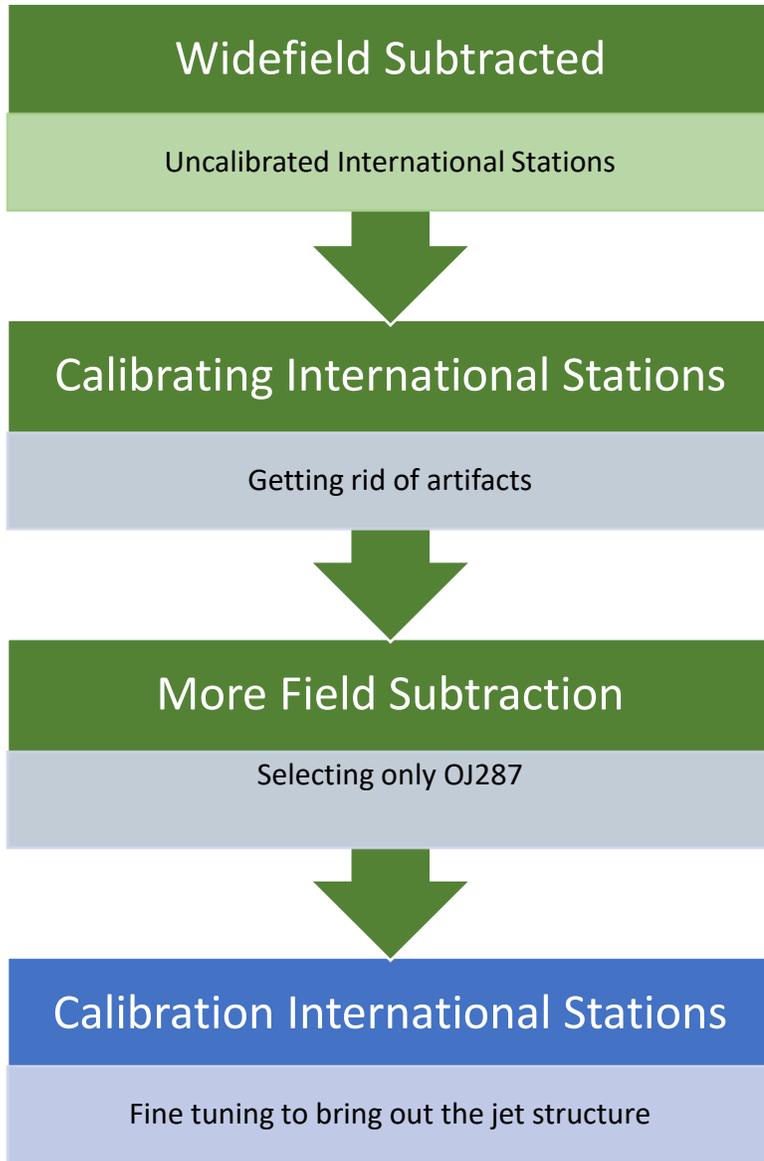
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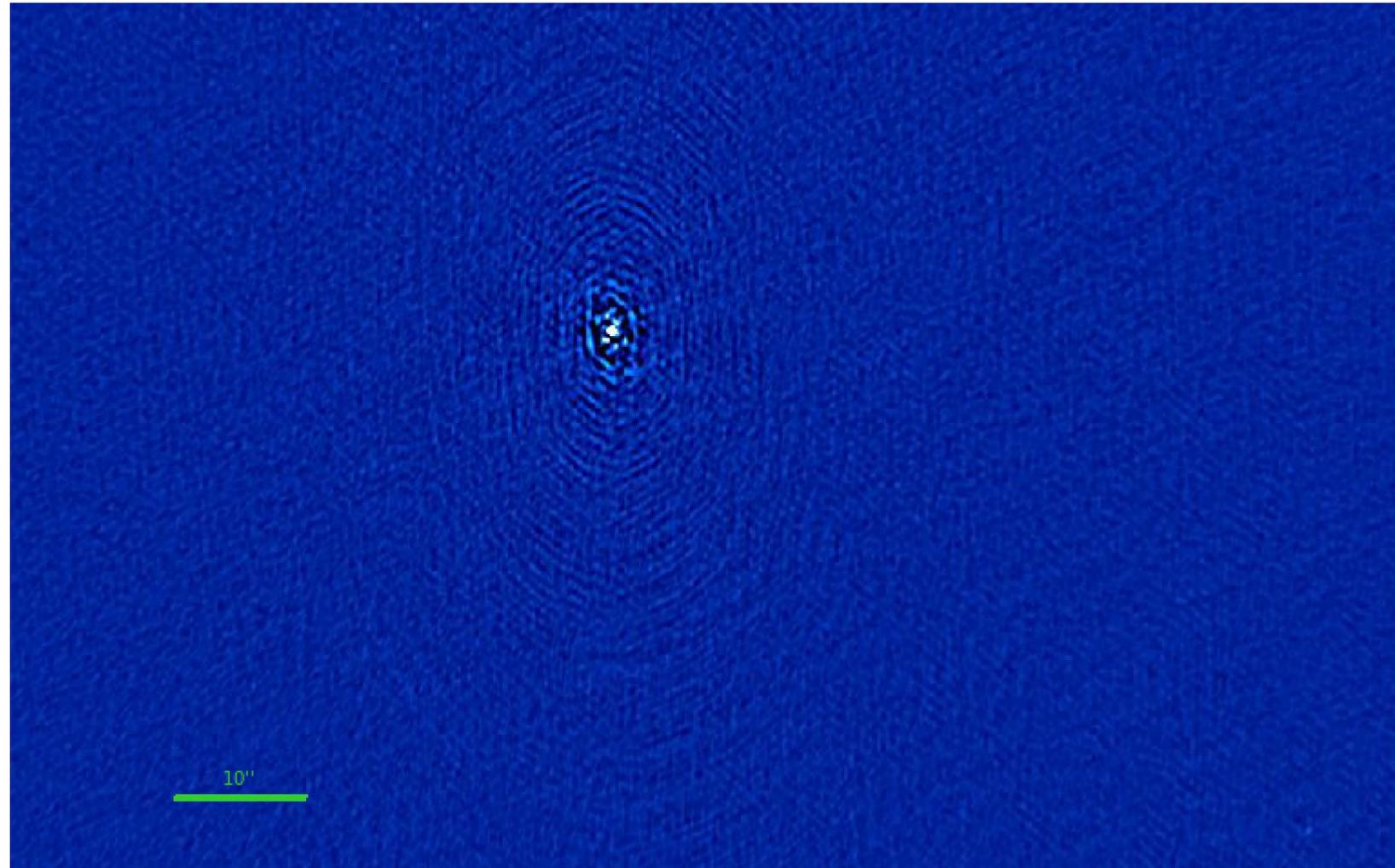
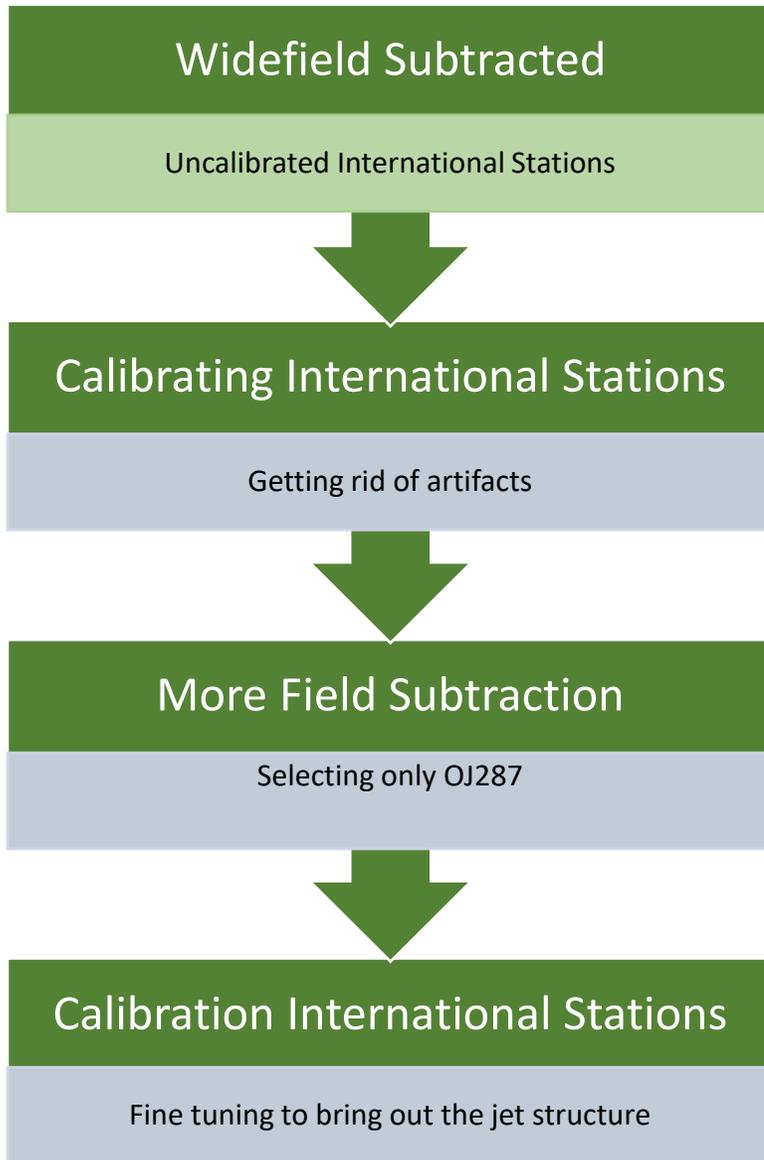


(4000'')



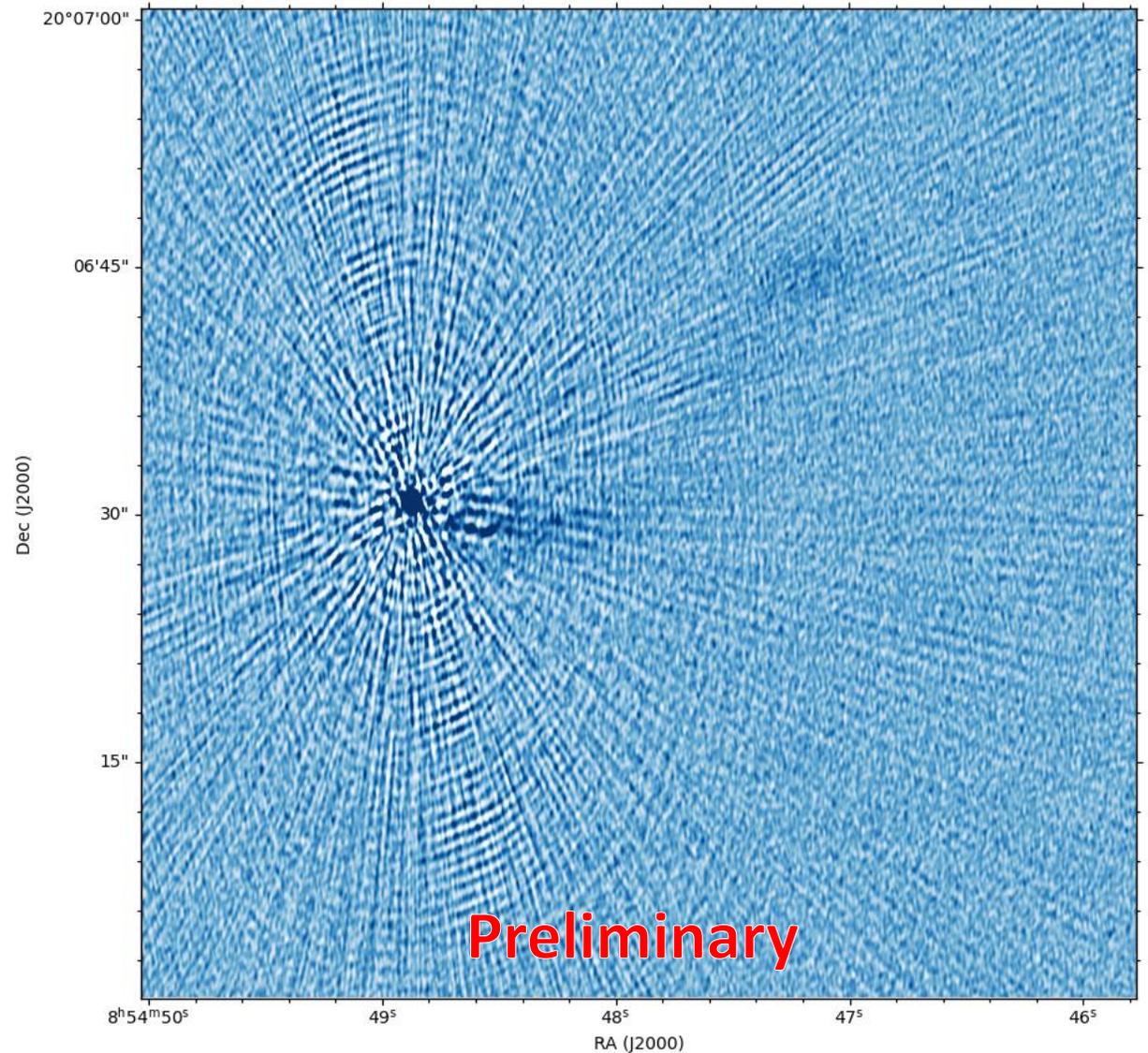
(4000'')





The journey so far...

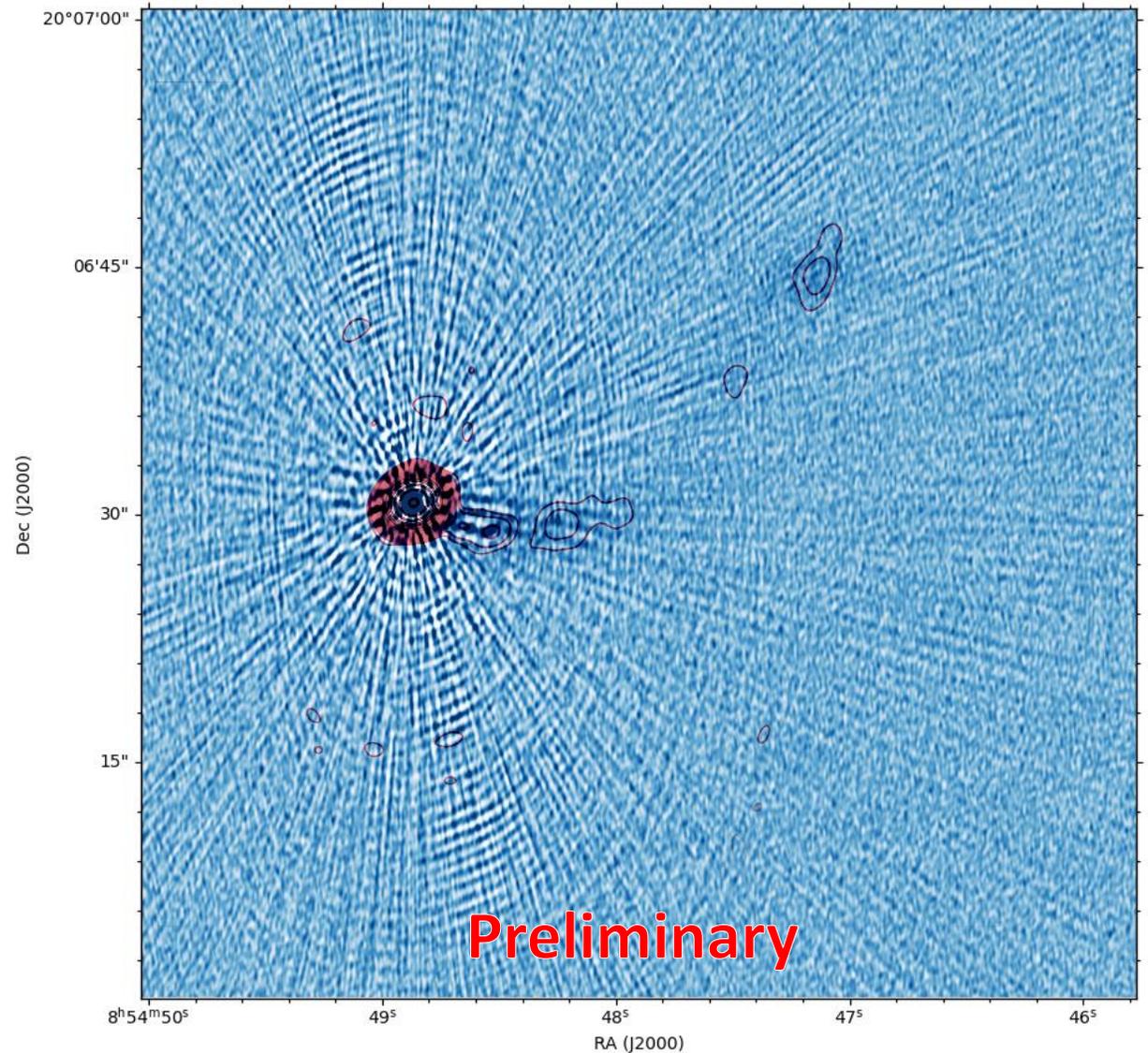
- 'Artisian' self calibration with **DDF** and **kMS**
 - Pro: Deep understanding of imaging
 - Con: Takes too long!
- Downloaded **LOFAR cycle 4 project 26** data for 6 blazars
- Setting up **LOFAR-VLBI pipeline** on local architecture in Würzburg
 - <https://github.com/lmorabit/lofar-vlbi>



Fullband LOFAR-VLBI image of OJ287

The journey so far...

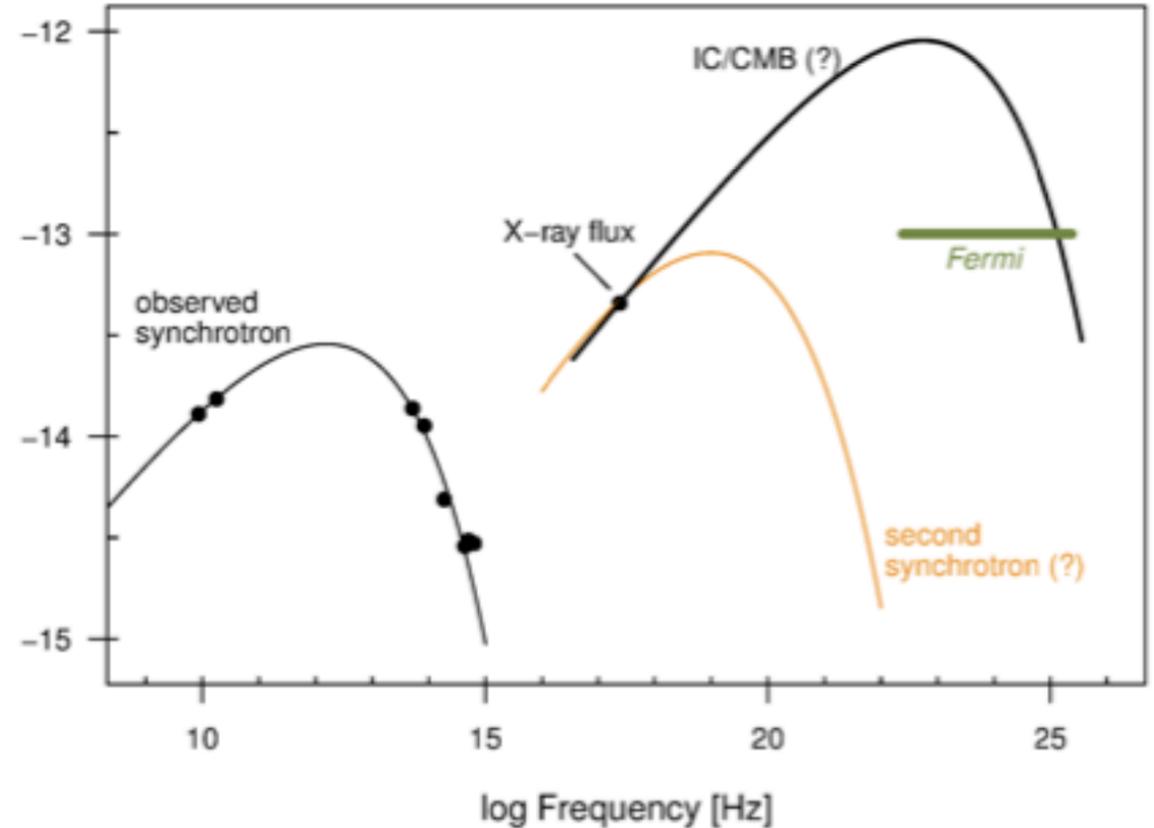
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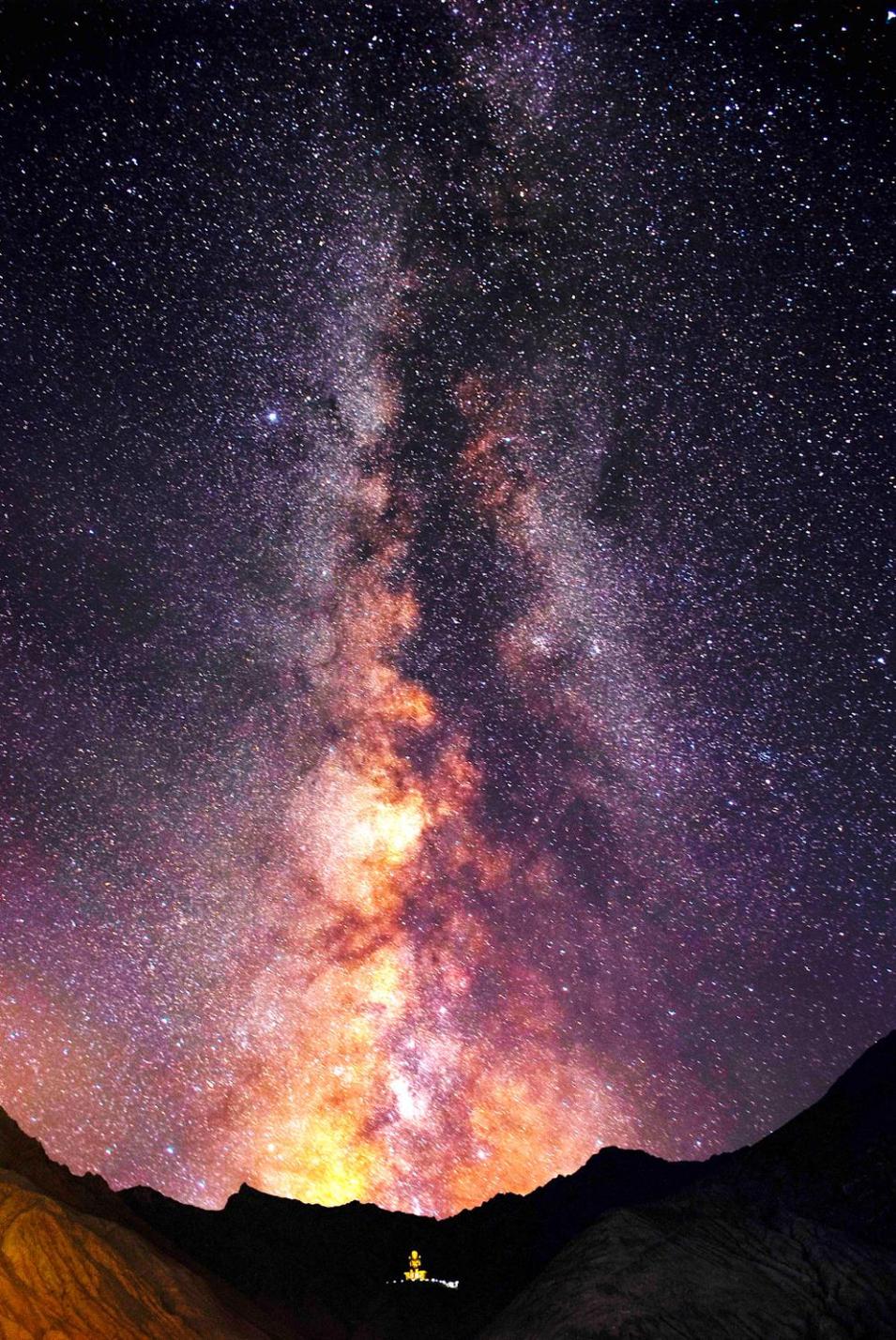
Fullband LOFAR-VLBI image of OJ287

The road ahead...

- Run LOFAR-VLBI pipeline and some more 'artisan' self calibration
- Complement with high-frequency radio, optical and X-ray data
- Spectral Energy Distribution (SED) modelling to test models
 - IC/CMB, SSC, second synchrotron
- Repeat the analysis for other X-ray selected blazars!



Radio-X-ray SED
Perlman et al. (2019)



Thank You!

Question?

Comments and suggestions?

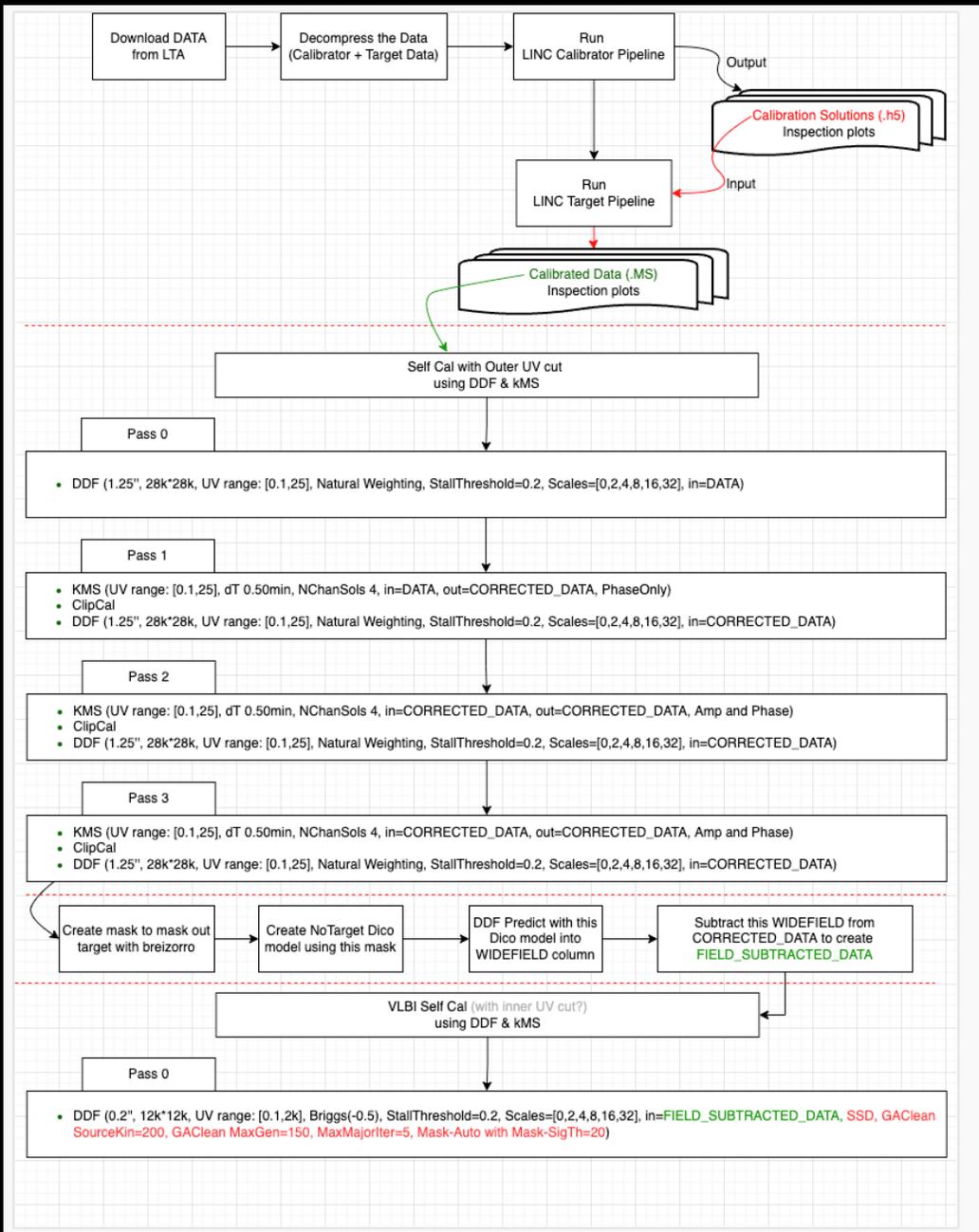
Email:

hrishikesh.shetgaonkar@uni-wuerzburg.de

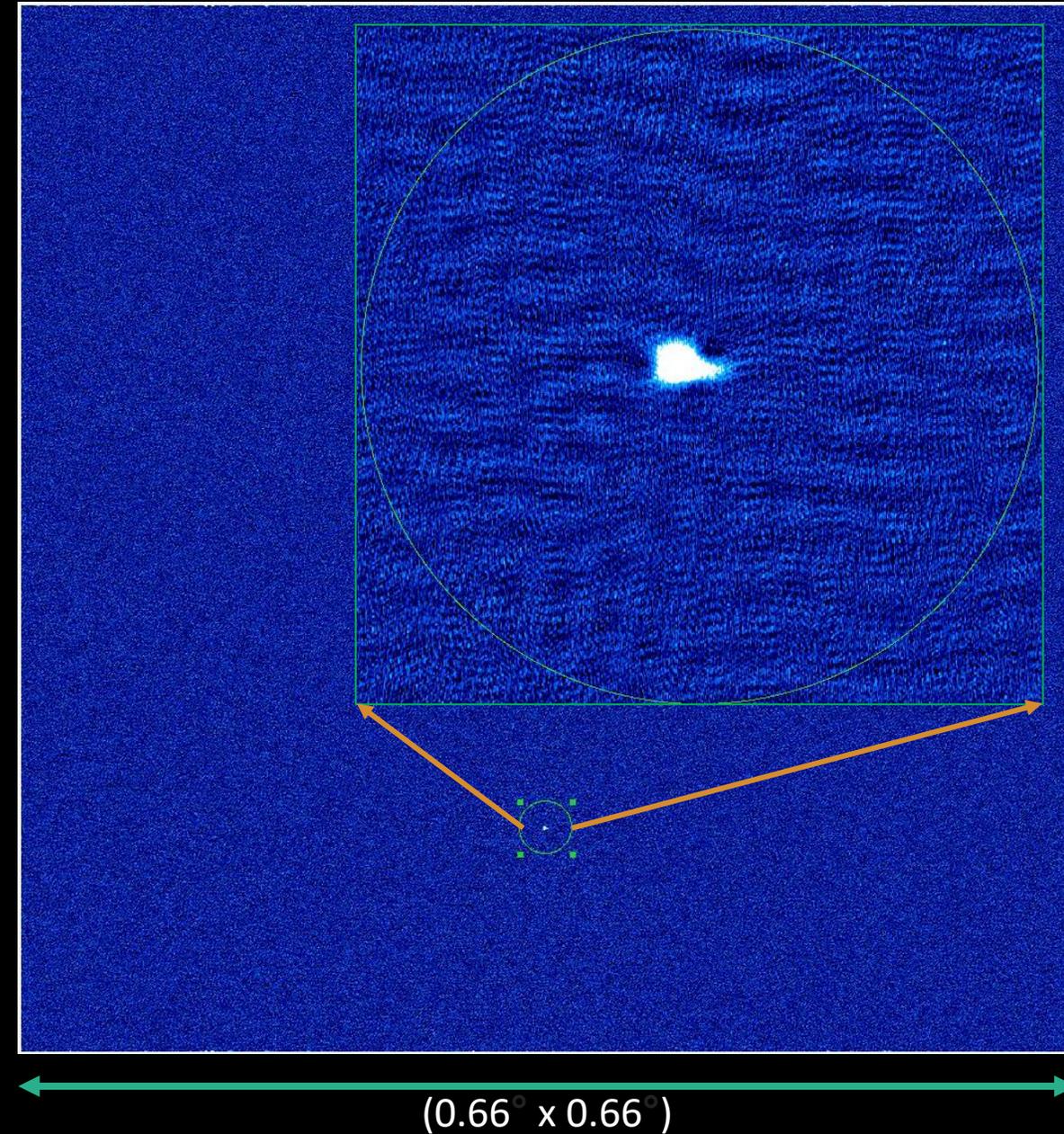
LinkedIn:

<https://www.linkedin.com/in/hrishikesh-shetgaonkar/>

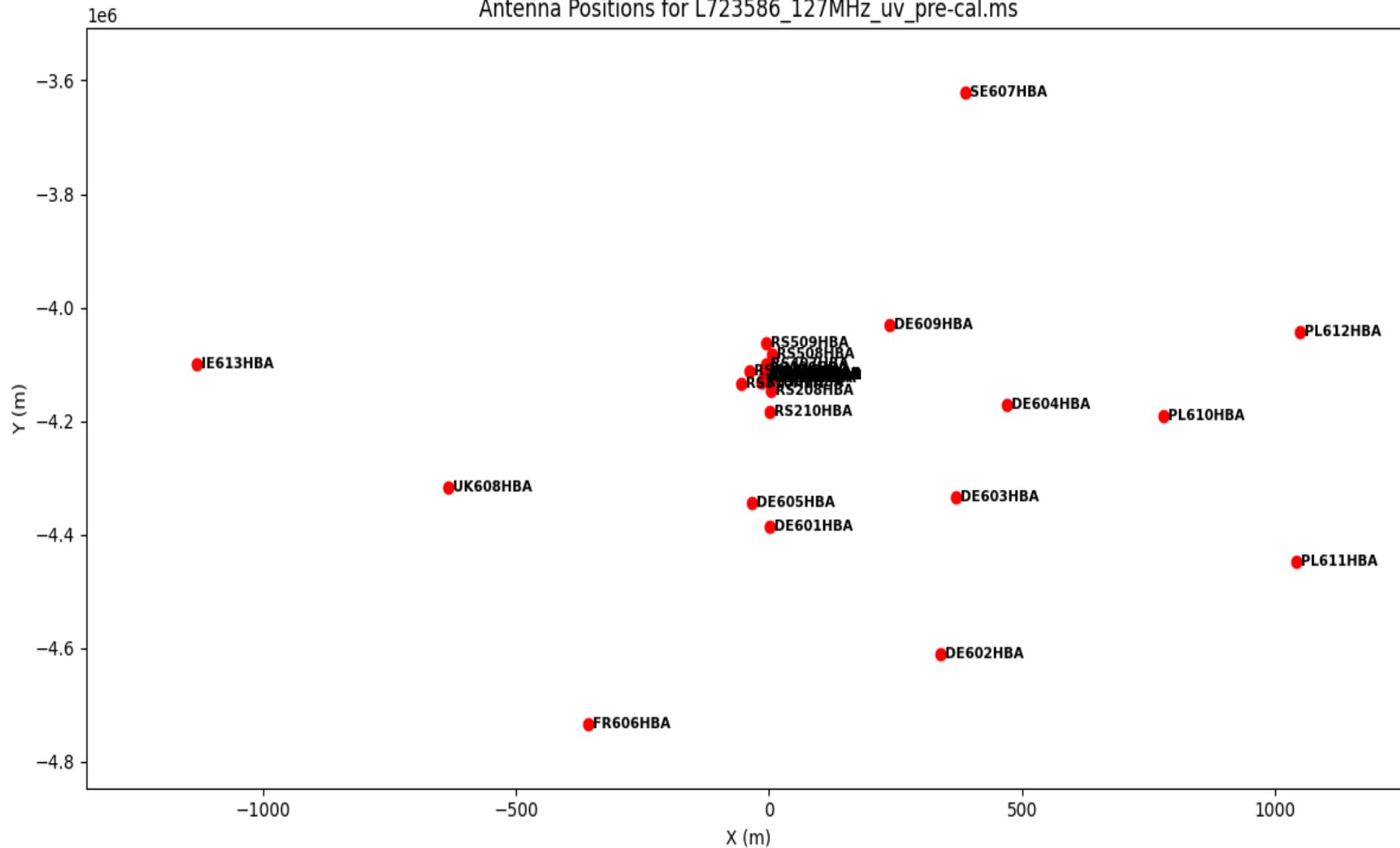




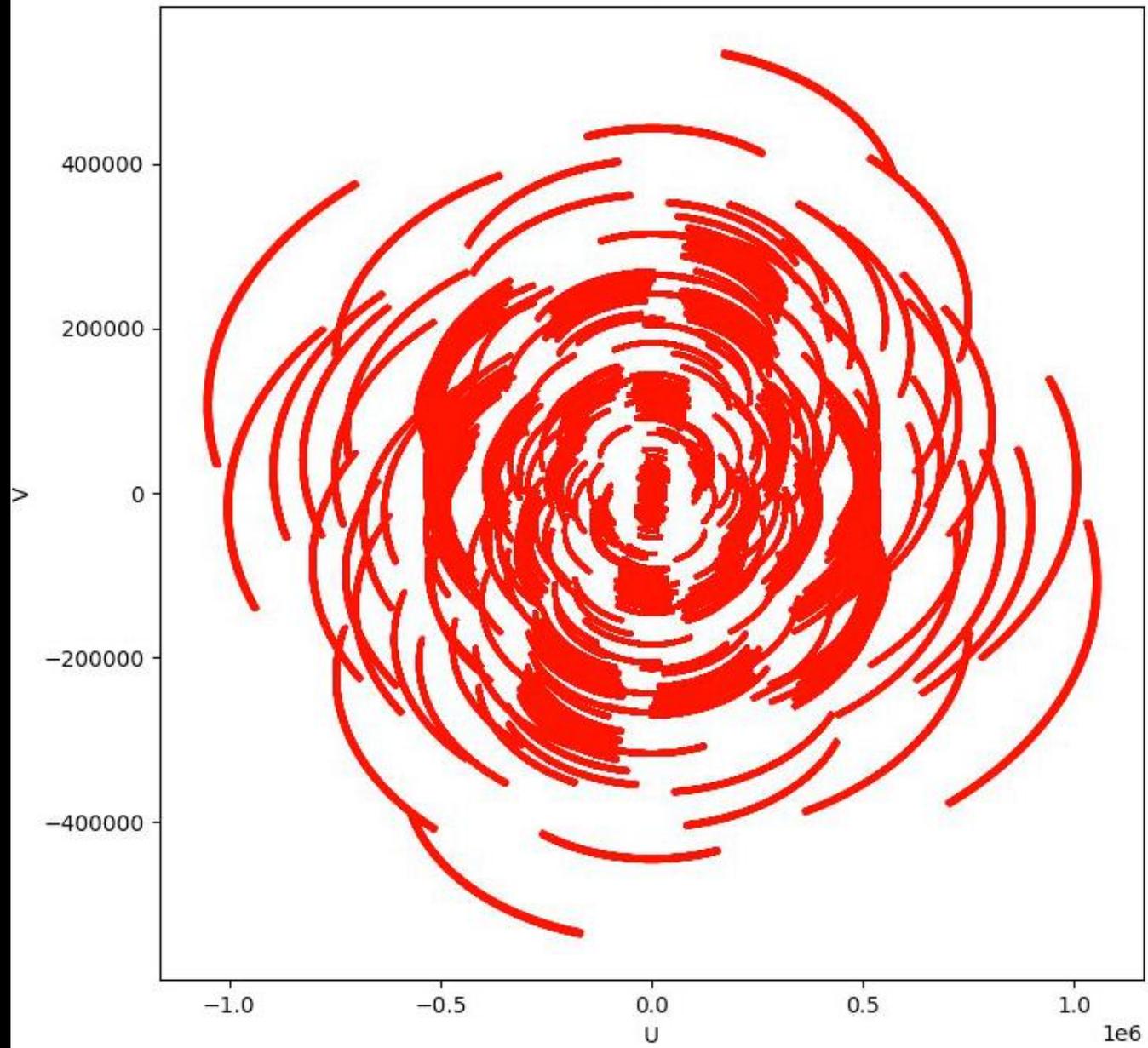
VLBI SelfCal, Pass 1, Target, Restored

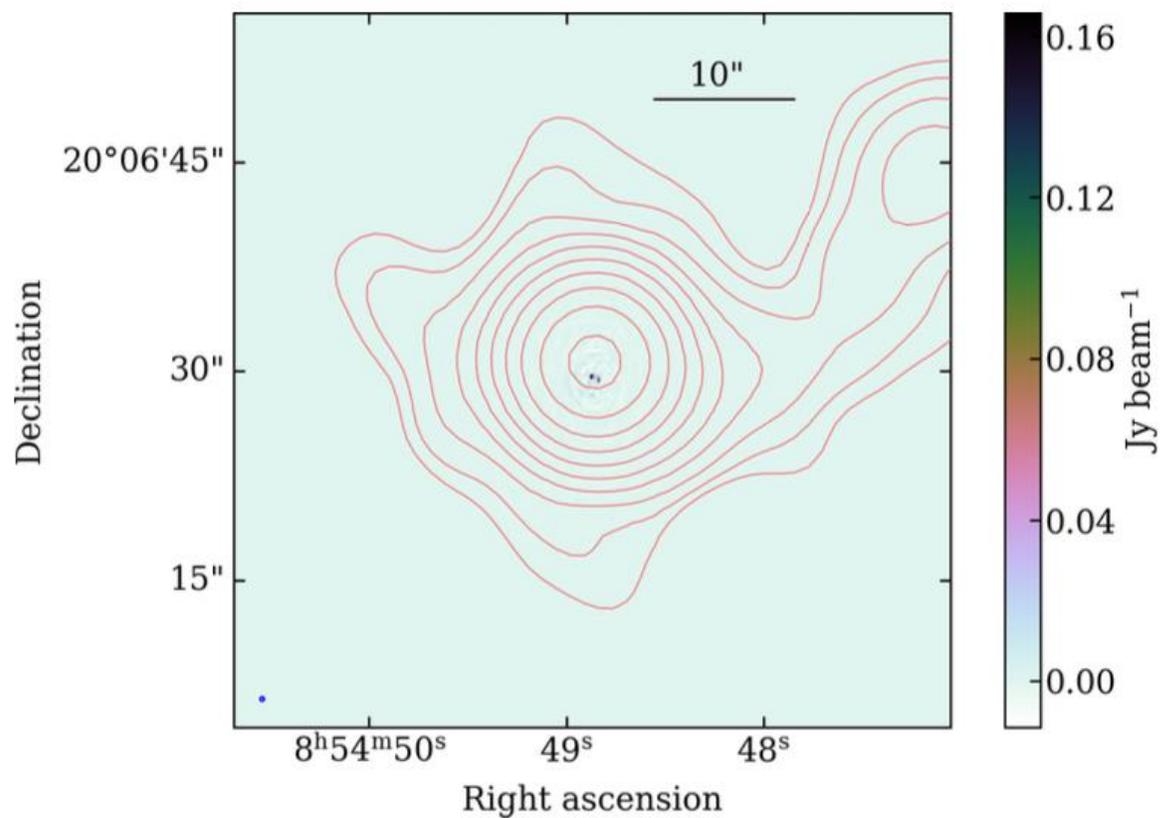


Antenna Positions for L723586_127MHz_uv_pre-cal.ms

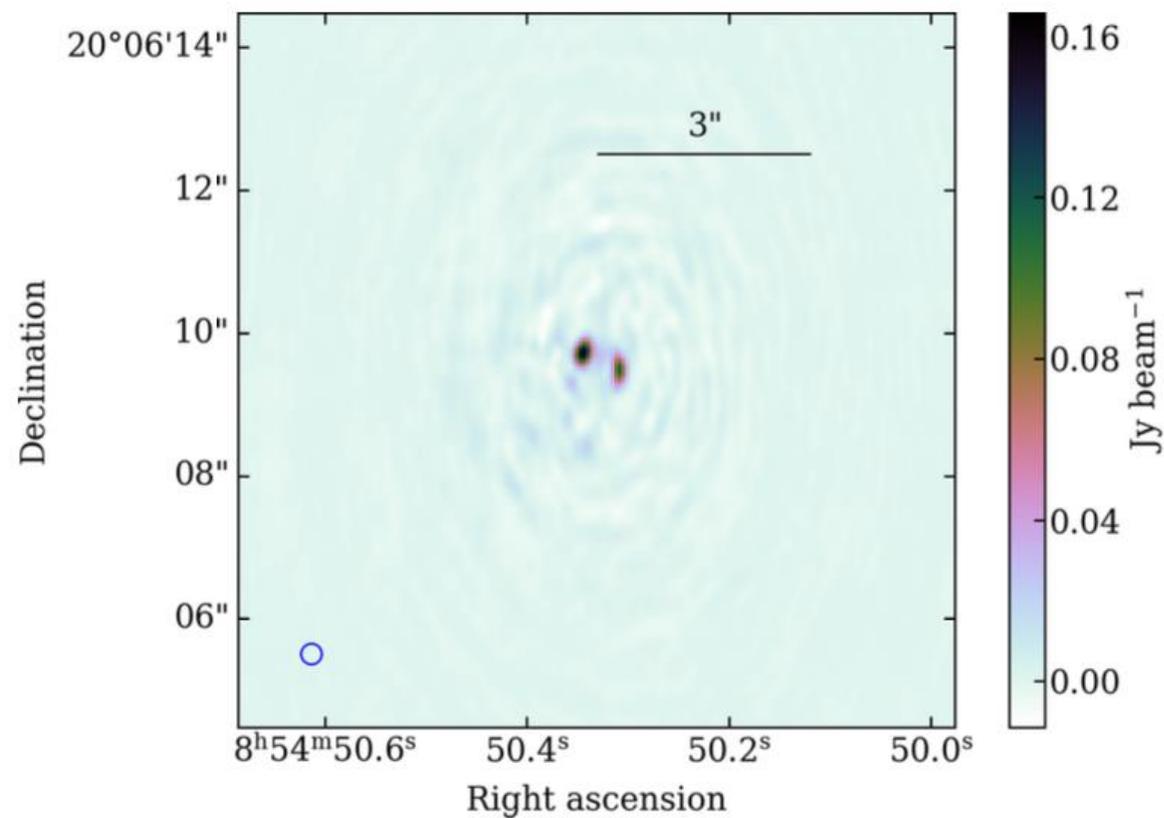


UV tracks

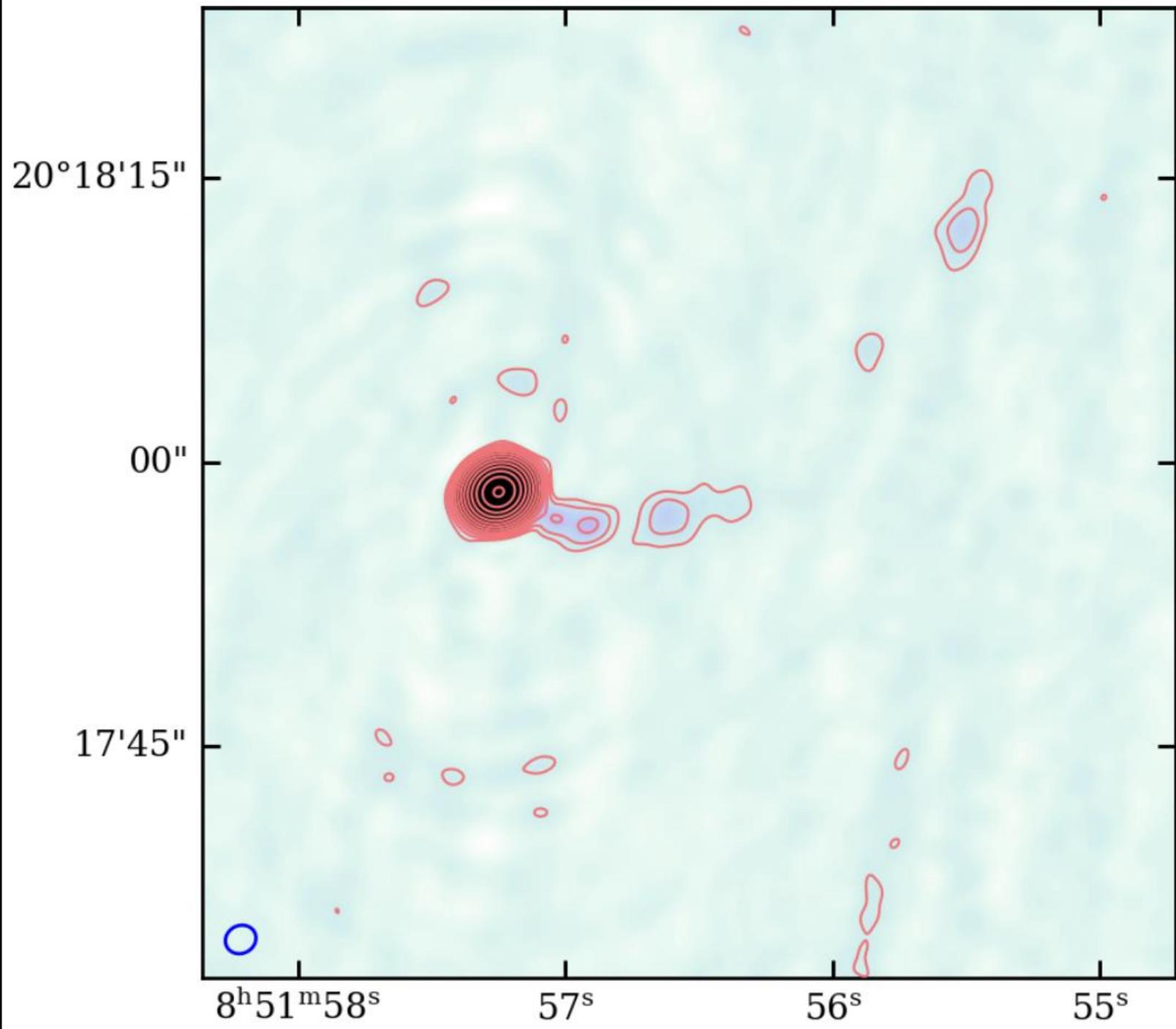


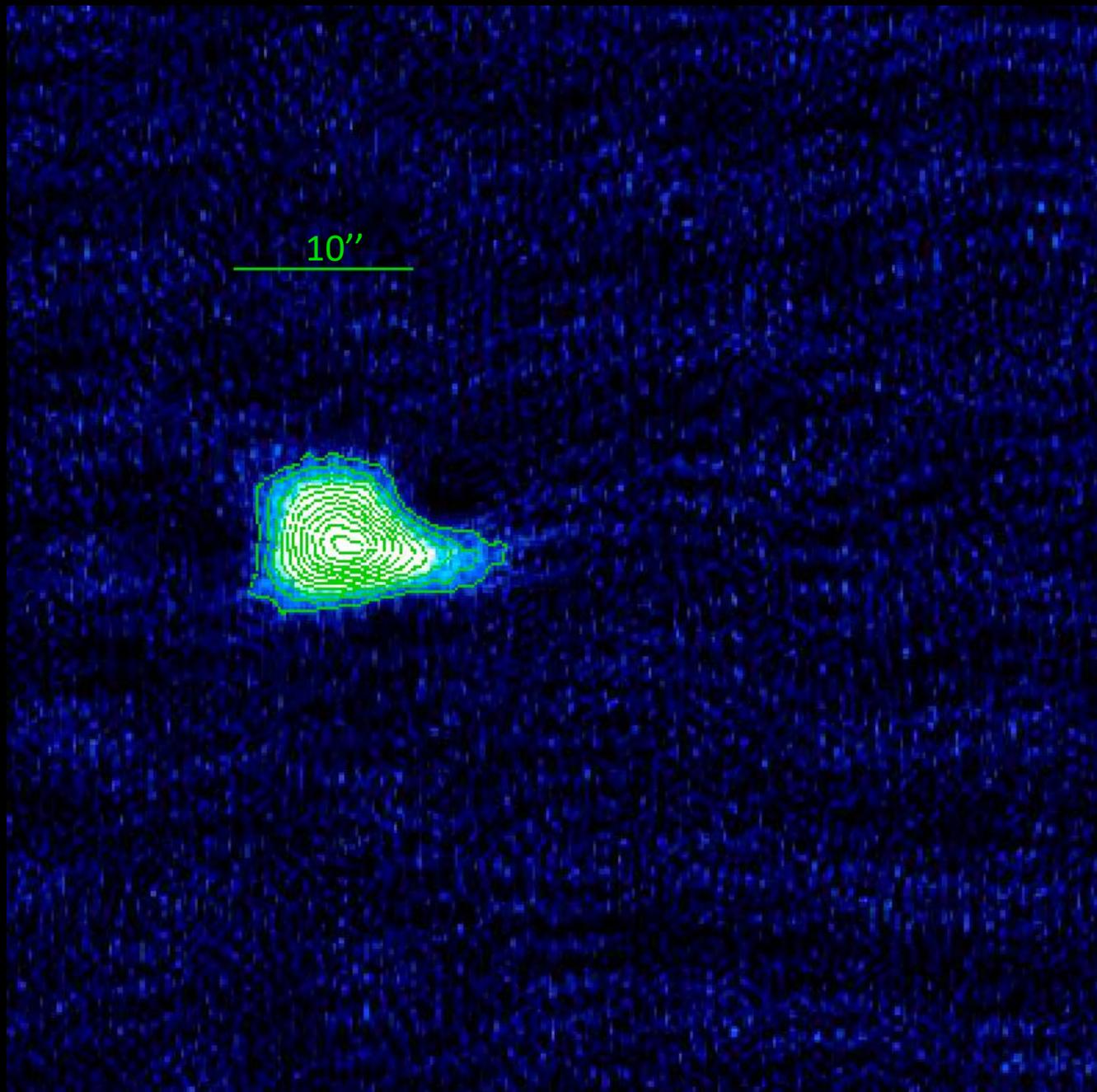


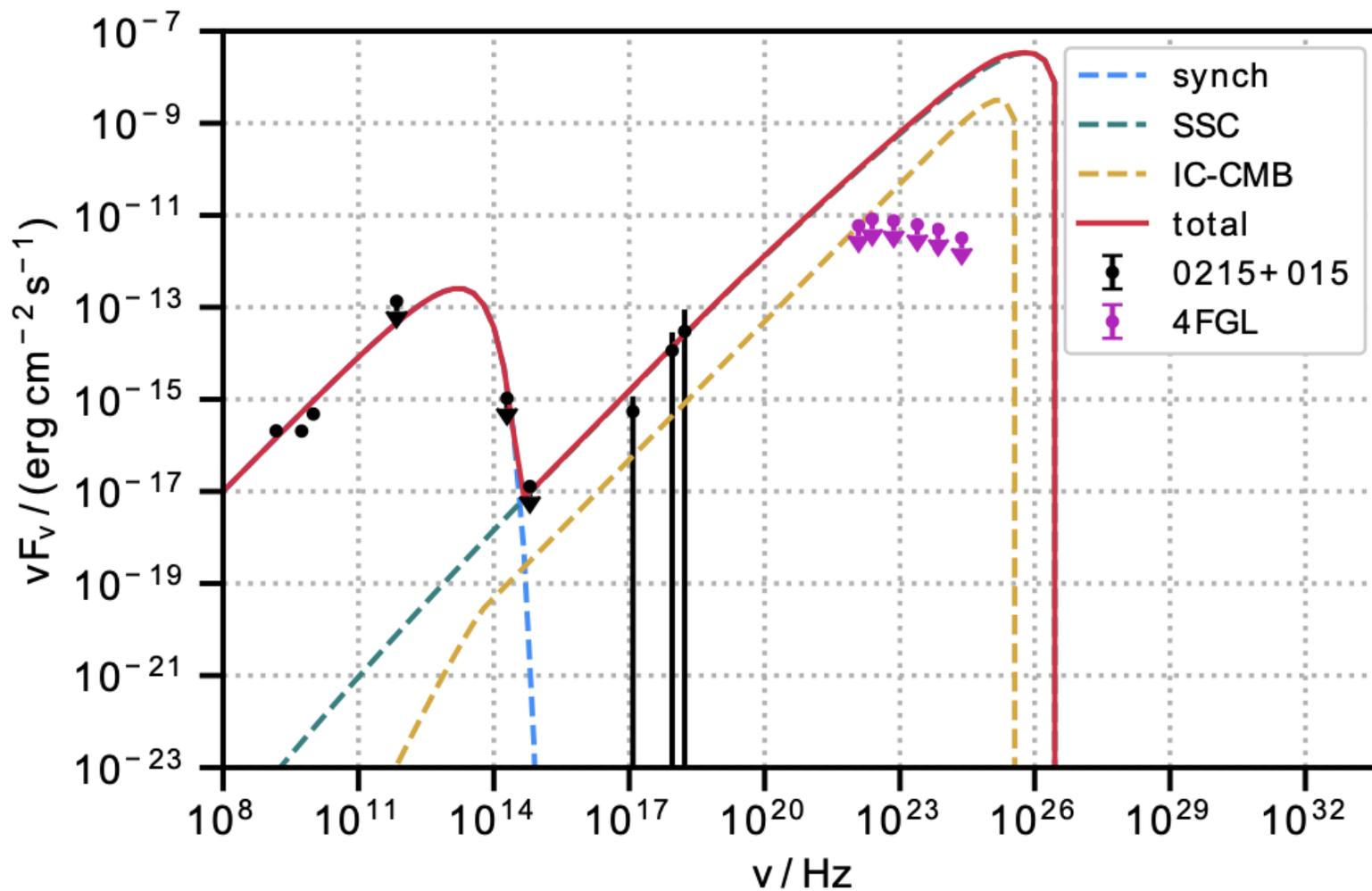
(a) Image that includes the international stations (colour map) overlaid with data from just the core and remote stations (contours).



(b) A 10" × 10" cutout of the image that includes the international stations.







Broadband SED models of extended jet regions of PKS 0215+015

Sebastian B., Kharb P., et al (2020)

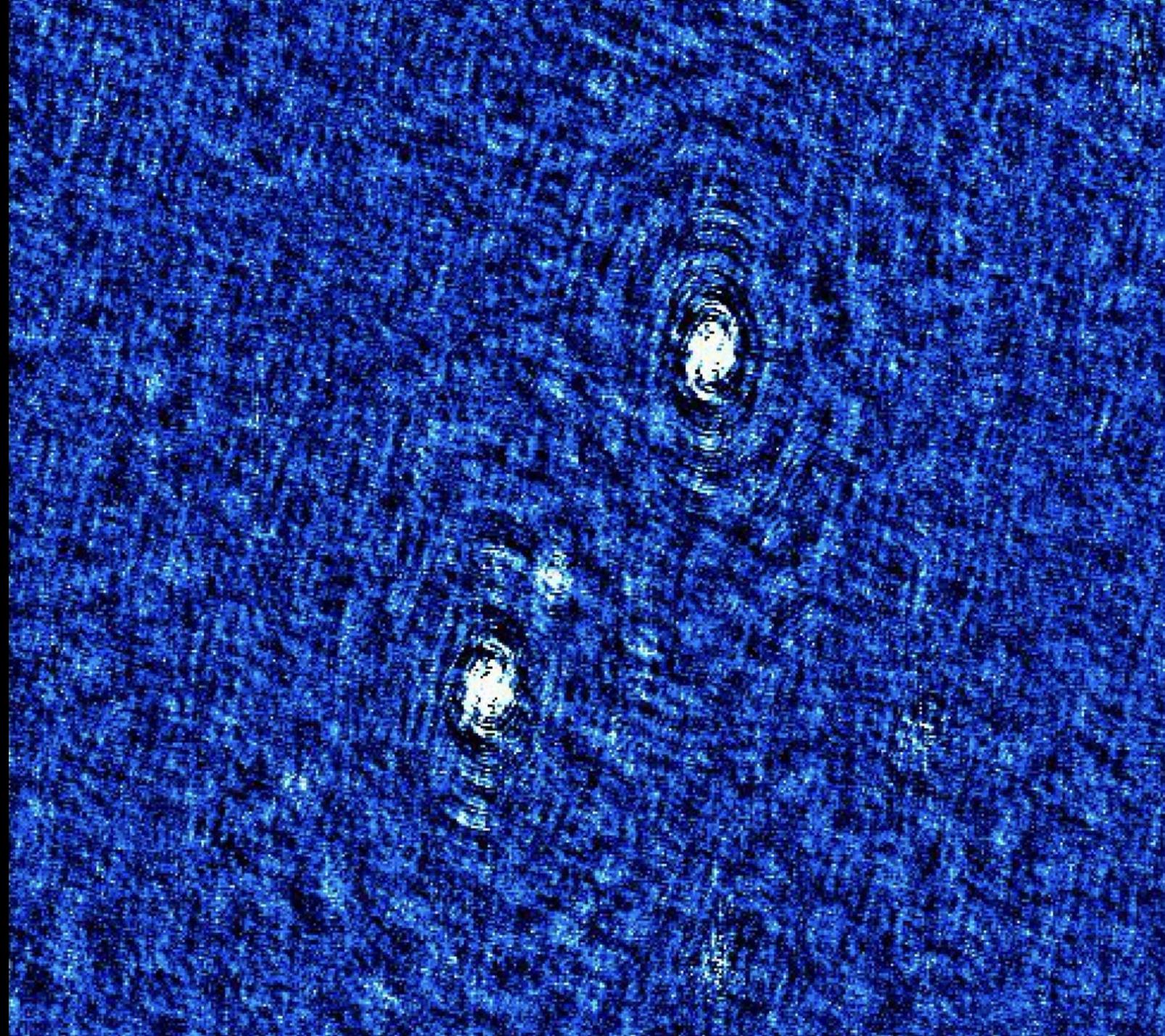
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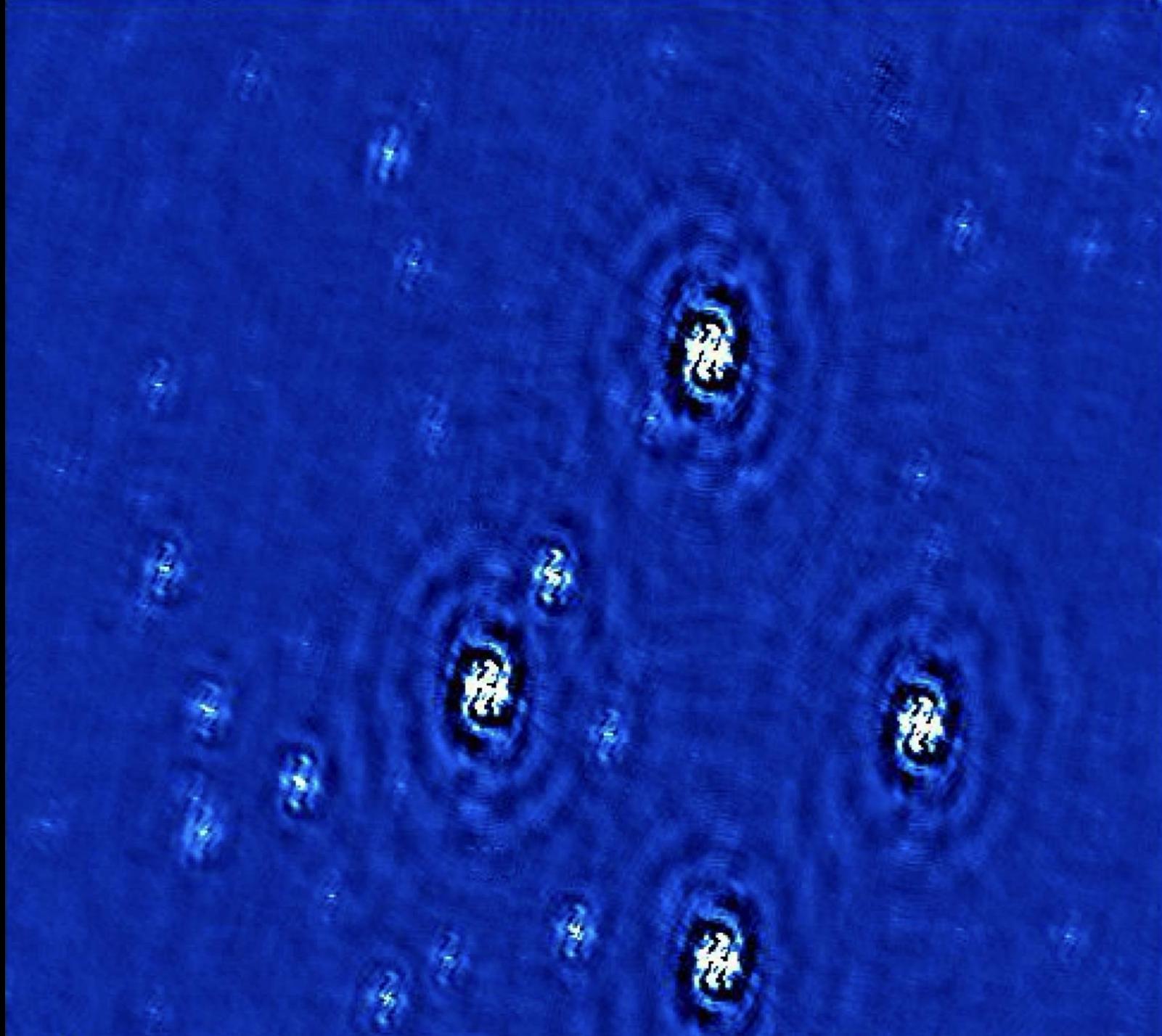
fk5

1 pixel = 0.2 arcsec

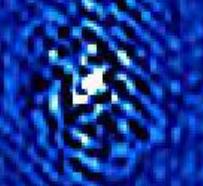
reg	sum	error	area	surf_bri	surf_err
			(arcsec**2)	(sum/arcsec**2)	(sum/arcsec**2)
---	---	-----	-----	-----	-----
1	-0.28891329	0.537507	1477.8	-0.000195502	0.000363721

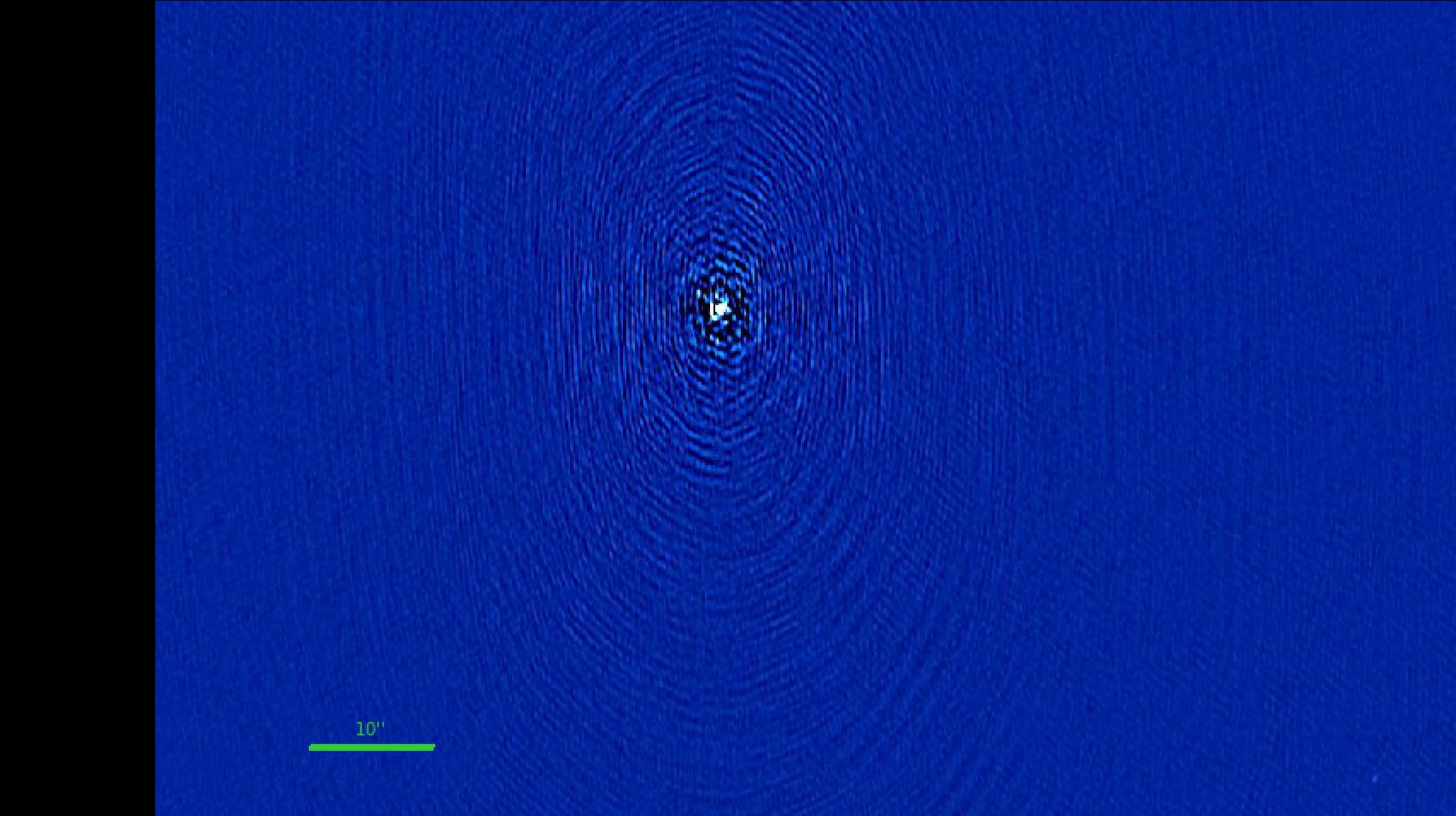
reg	sum	npix	mean	median	min	max	var	stddev	rms
---	---	-----	-----	-----	---	---	---	-----	---
1	-0.28891329	36945	-7.82009e-06	-2.4712e-05	-0.016241	0.0152849	1.55683e-05	0.00394567	0.00394568





10''





10''

Observation Description	Assessment
1641+399/1/TO (Target Observation)	
3C196/1/CO (Calibration Observation)	
1803+784/1/TO (Target Observation)	
3C48/1/CO (Calibration Observation)	
2230+114/1/TO (Target Observation)	
3C48/1/CO (Calibration Observation)	
1226+023/1/TO (Target Observation)	
3C196/1/CO (Calibration Observation)	
1823+568/1/TO (Target Observation)	
3C48/1/CO (Calibration Observation)	
0838+133/1/TO (Target Observation)	
3C196/1/CO (Calibration Observation)	

“Sto

Stay on target

