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Leiden



European Research Council

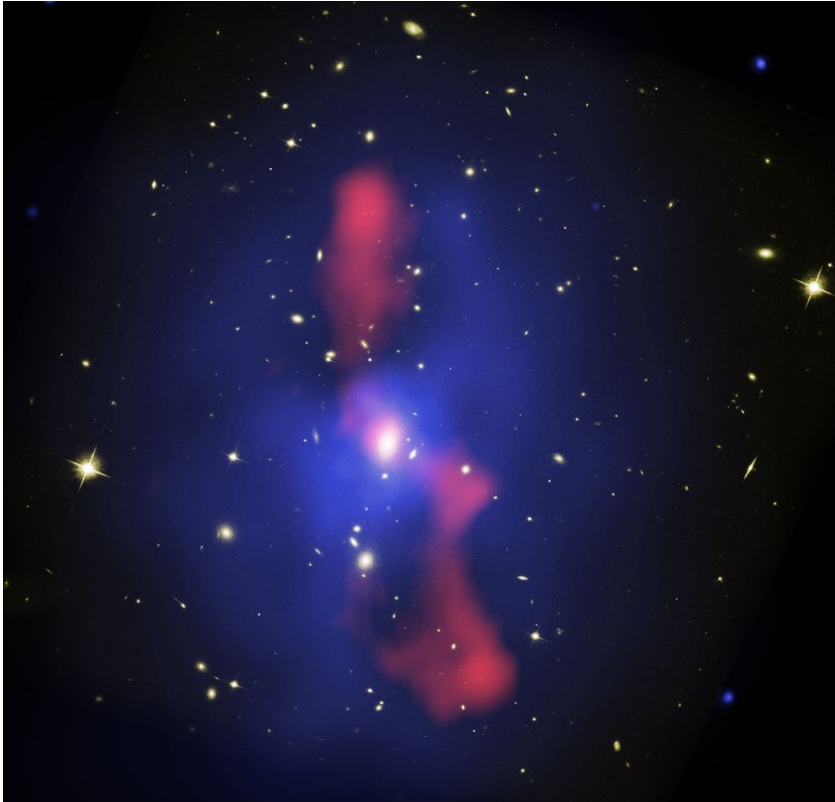
Established by the European Commission

Radio-mode feedback in high-redshift galaxy clusters with the International LOFAR Telescope

A new window on feedback opened by subarcsecond LOFAR-VLBI observations

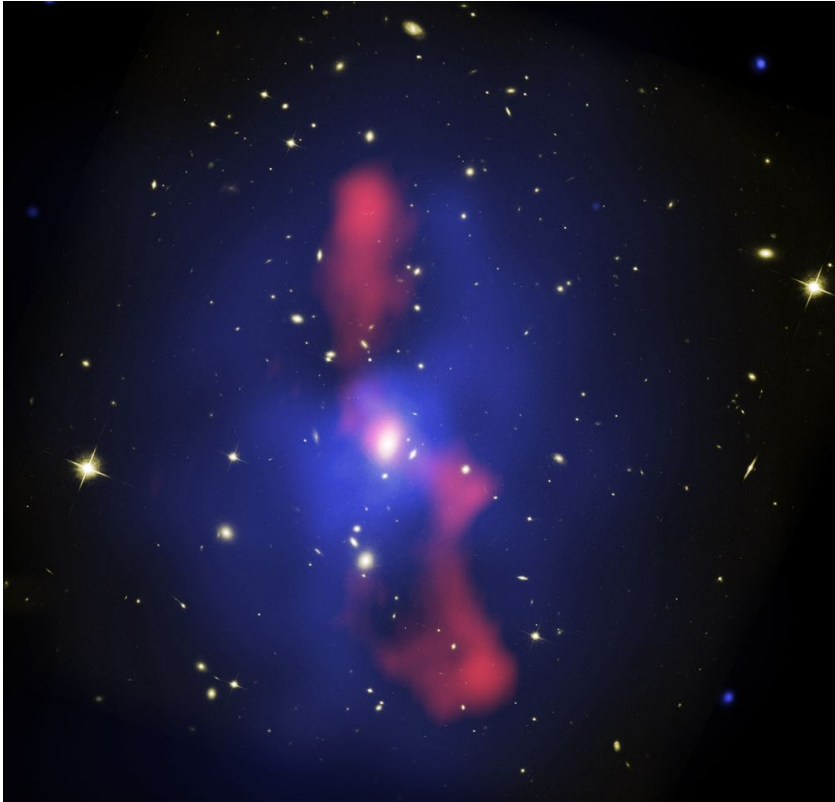
R. Timmerman, R. J. van Weeren, A. Botteon, L. K. Morabito,
F. Sweijen, H. J. A. Röttgering, L. Bîrzan, B. R. McNamara

Measuring the amount of radio-mode feedback



(NASA, ESA, CXC, STScI, B. McNamara, NRAO/AUI/NSF, and L. Birzan & team)

Measuring the amount of radio-mode feedback

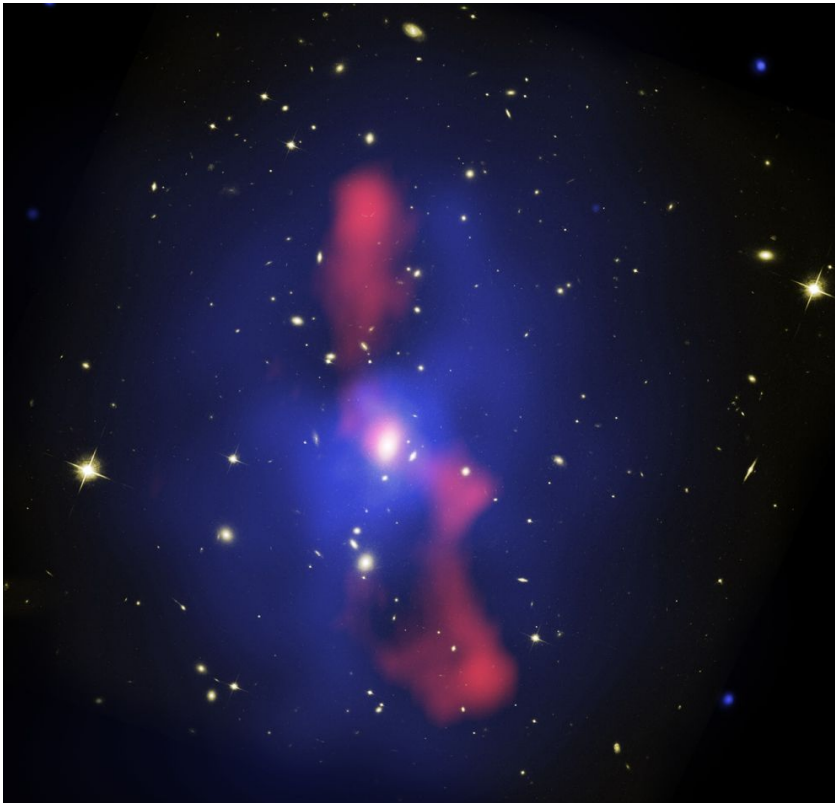


Radio

- Radio-mode feedback power derived from the **luminosity** and **spectral properties** of the radio lobes

(NASA, ESA, CXC, STScI, B. McNamara, NRAO/AUI/NSF, and L. Birzan & team)

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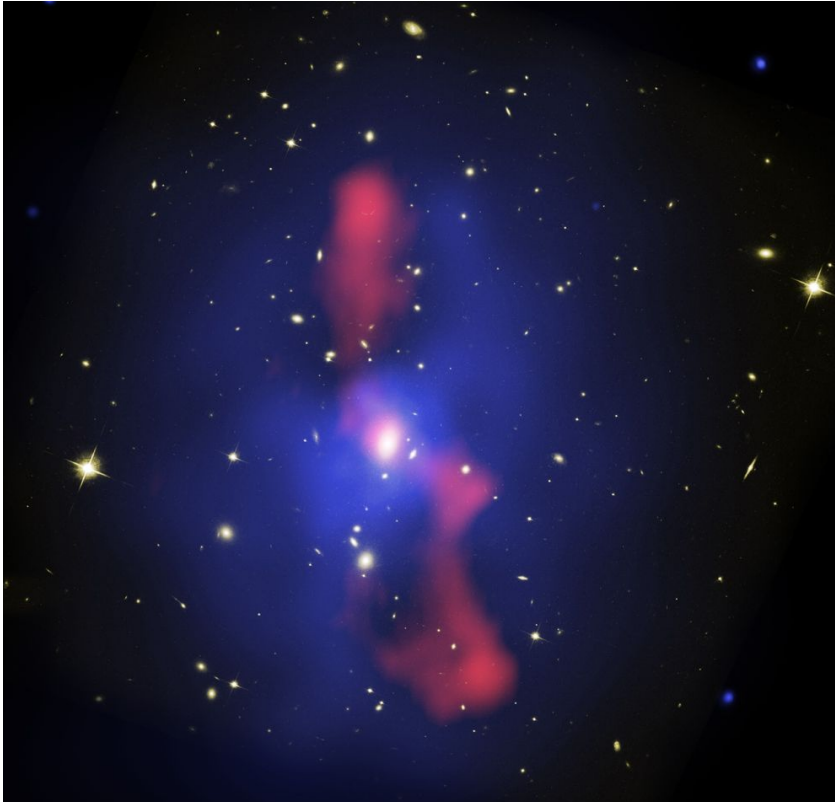


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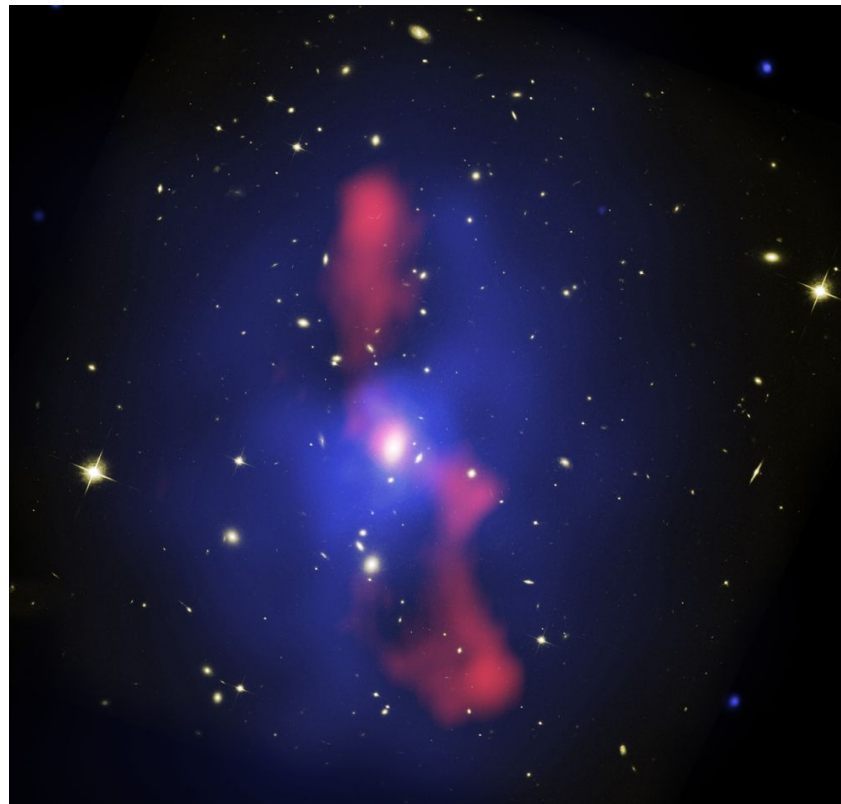


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- Radio-mode feedback power derived from the **luminosity** and **spectral properties** of the radio lobes
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- Unknown parameters result in large **scatter**

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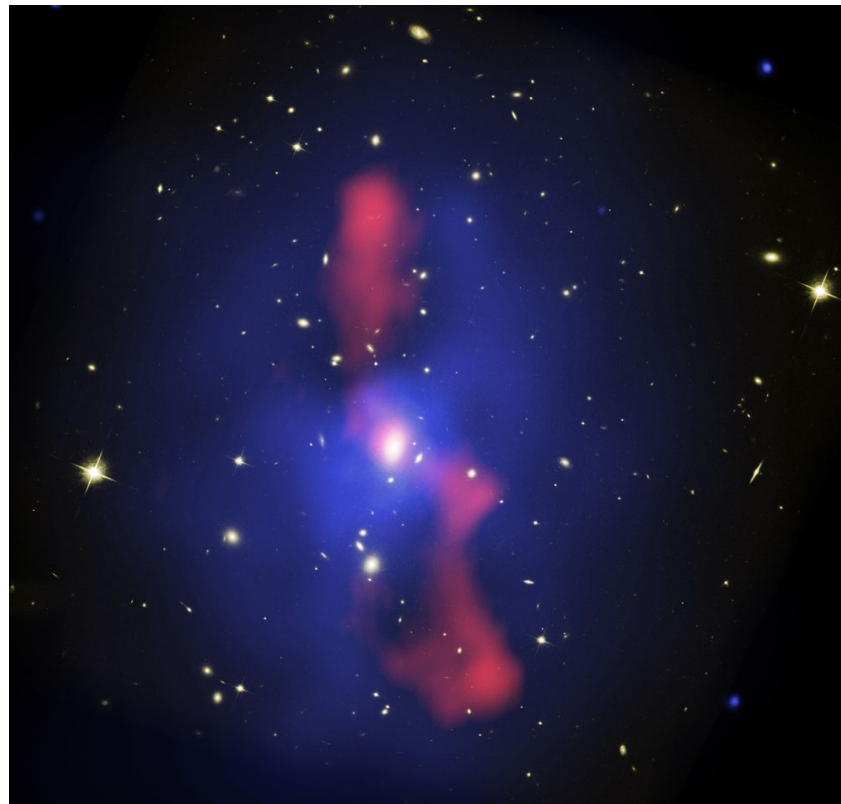
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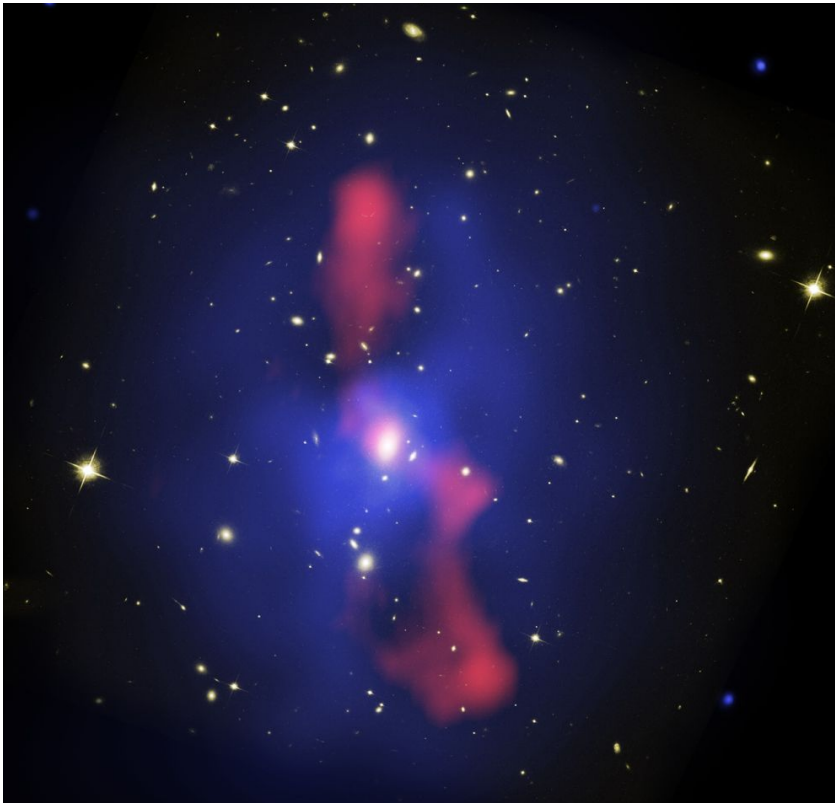
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Measuring the amount of radio-mode feedback



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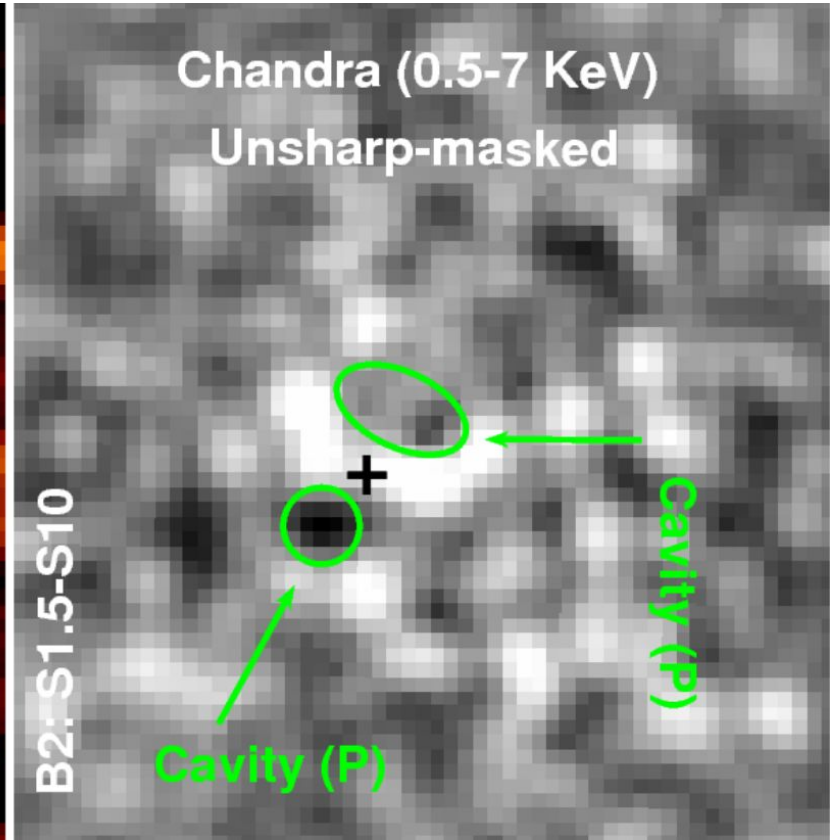
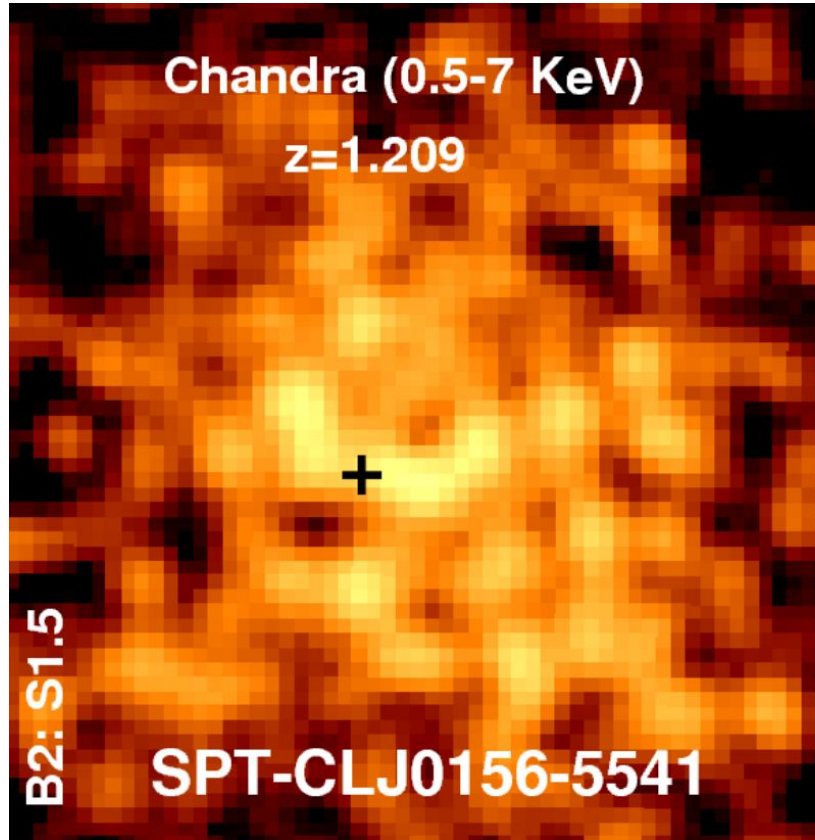
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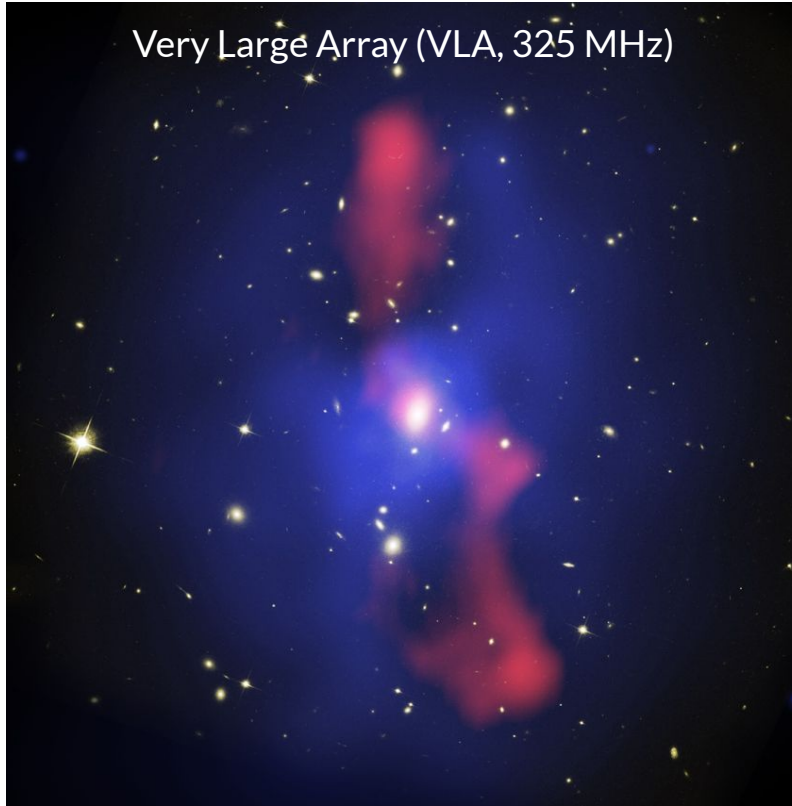
- Radio-mode feedback power derived from the **volume** of the cavities and their **distance** to the core, and the **ICM pressure** surrounding the cavities
- Observables **directly** provide power through fundamental physics
- Generally **preferred** for measuring amount of feedback

(NASA, ESA, CXC, STScI, B. McNamara, NRAO/AUI/NSF, and L. Birzan & team)

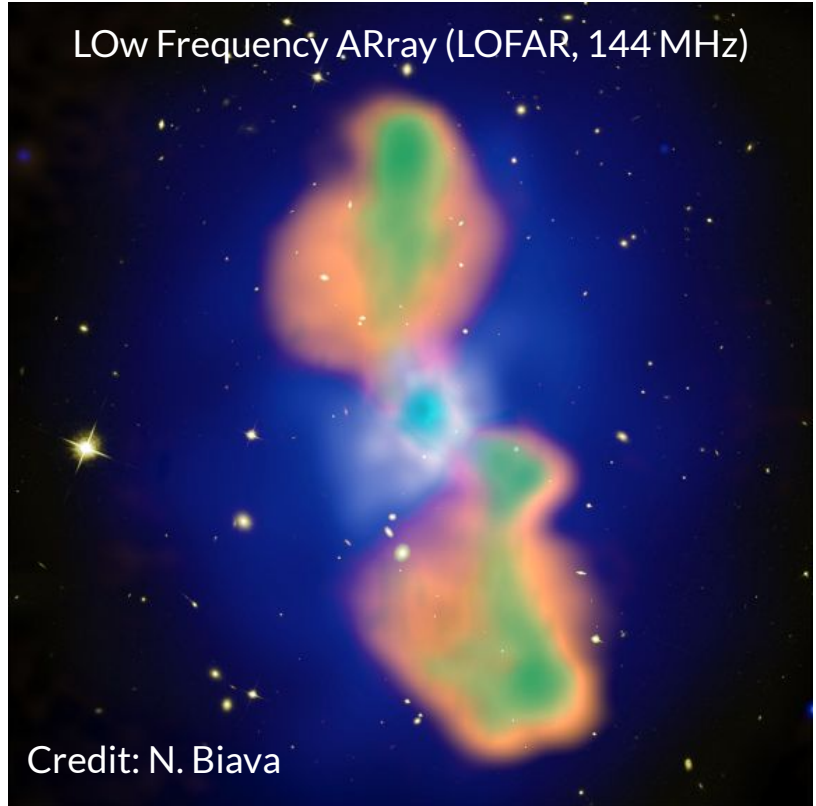
The challenge of high redshifts



The benefit of low-frequency radio

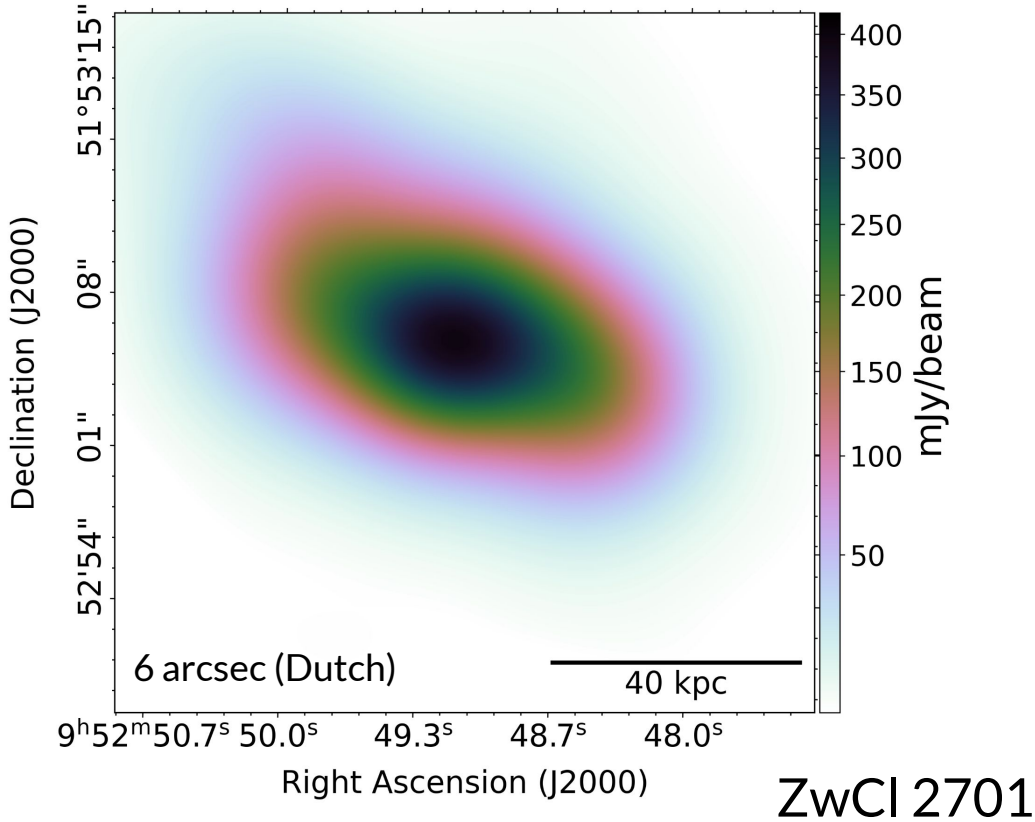


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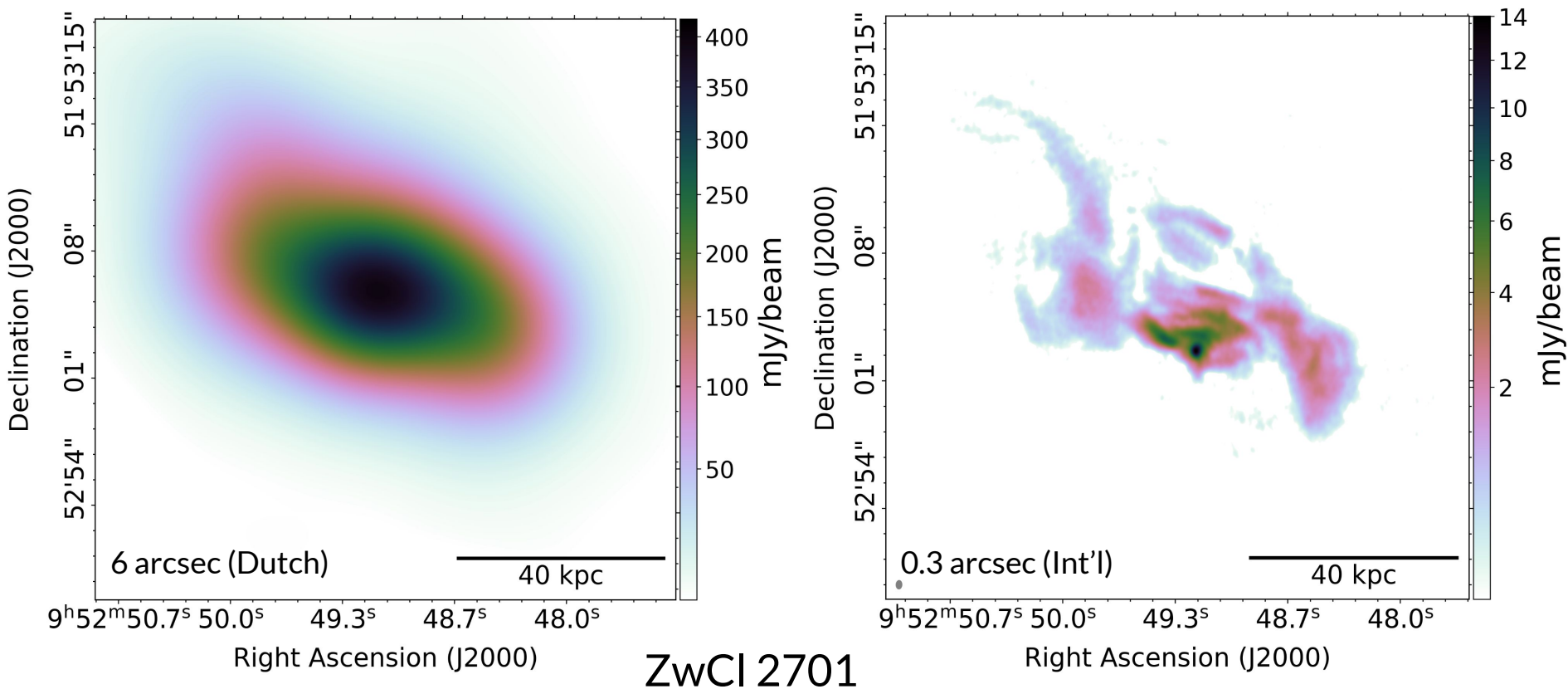


(N. Biava et al. 2021 A&A 650 A170)

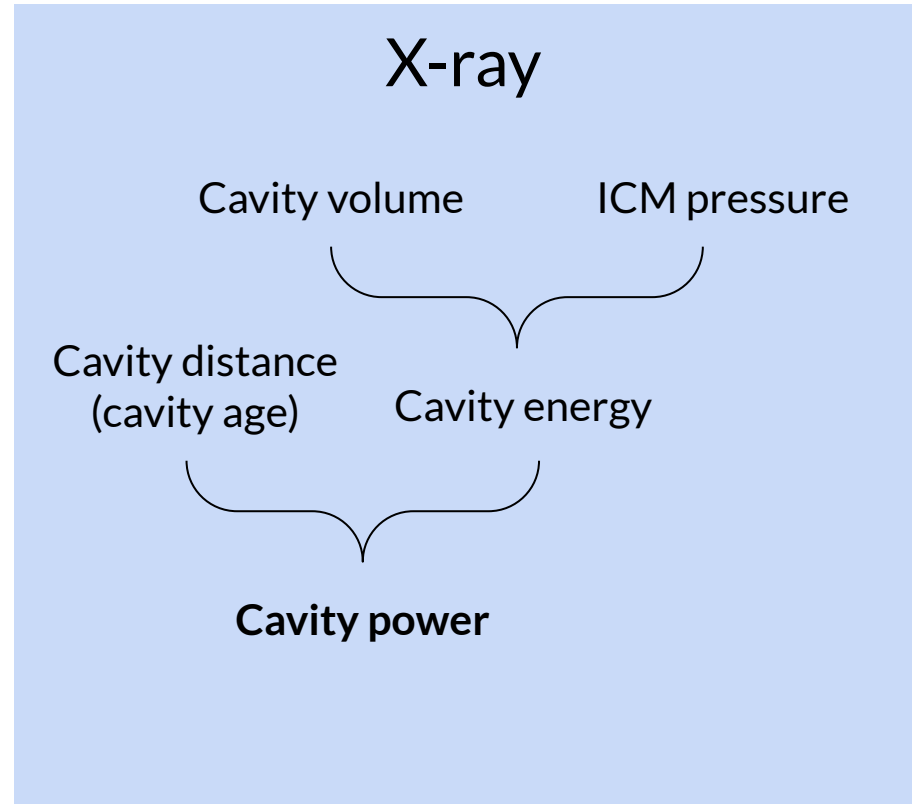
Low frequencies at high angular resolutions



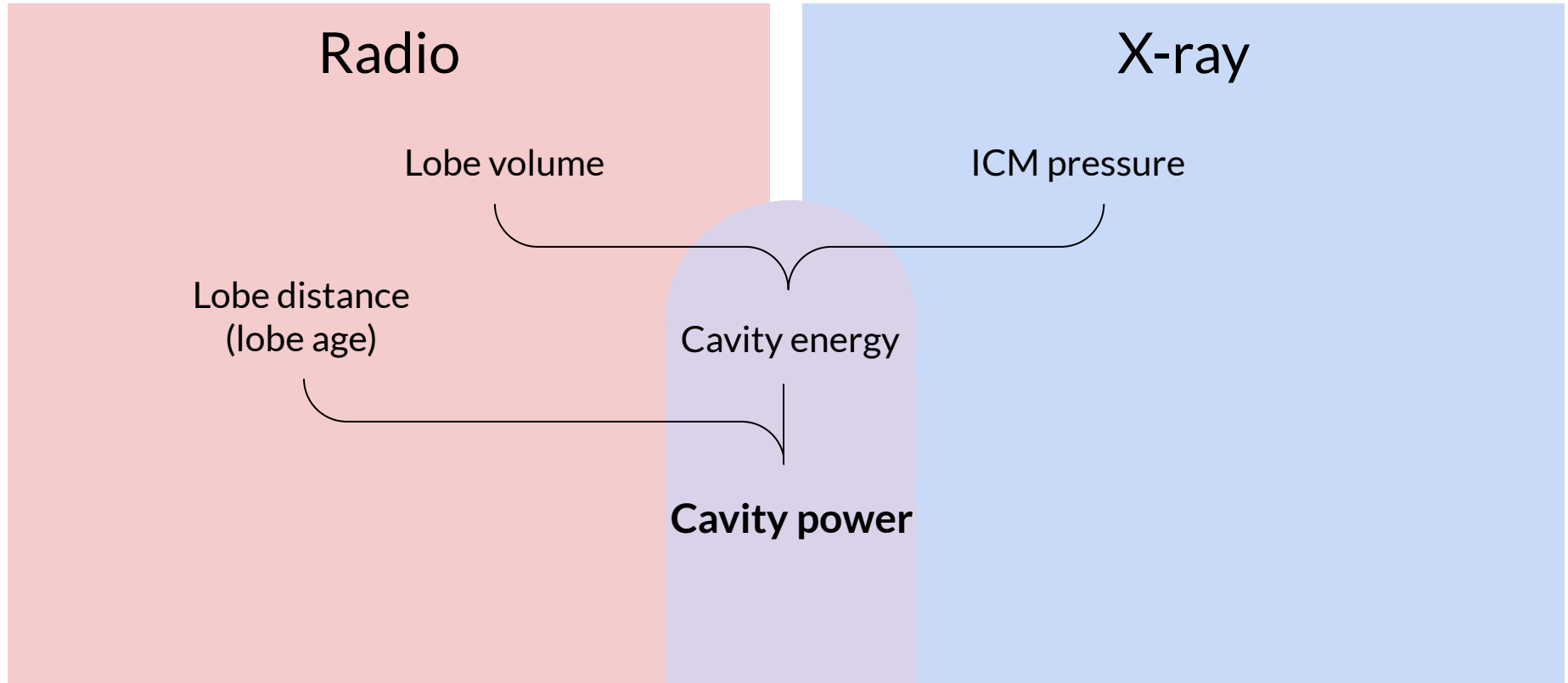
Low frequencies at high angular resolutions



The normal X-ray-based approach



The hybrid radio—X-ray approach



Low-redshift observations

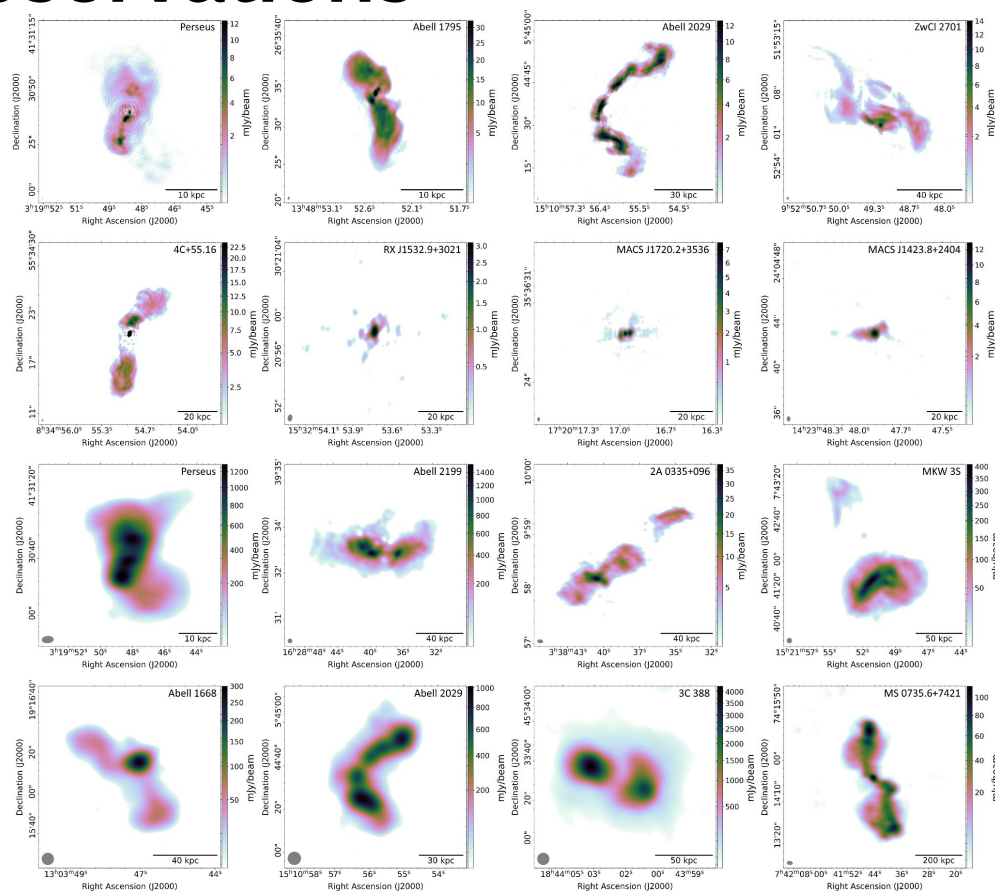
Sample:

- 14 cool-core clusters
- Known X-ray cavities
- Presence of radio source
- $0 < z < 0.6$

Rafferty et al. (2006)

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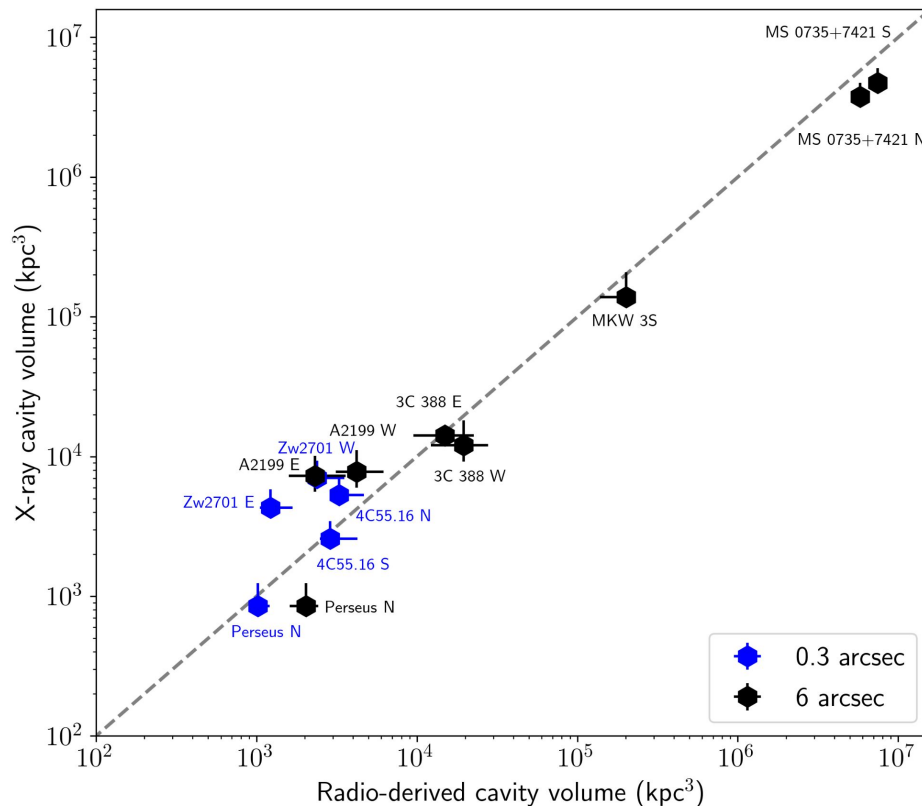
Bîrzan et al. (2008, 2020)



$\Theta=0.3''$

$\Theta=6''$

Cavity/lobe volume comparison



Selected cavities!

Overall **good** agreement
($\sigma = 0.30$ dex)

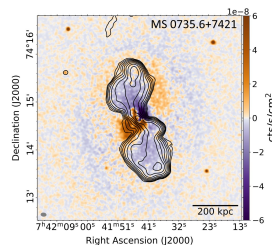
Previous monochromatic
radio-based measurements
resulted in $\sigma = 0.8$ dex

→ Low-frequency radio
observations detect the
complete radio lobes!

Further discussion points

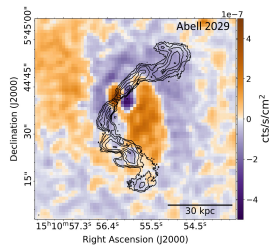
Correlation **improves** with
how well-defined cavities are

FOM = 2



$\sigma = 0.34$ dex

FOM = 3



$\sigma = 0.49$ dex

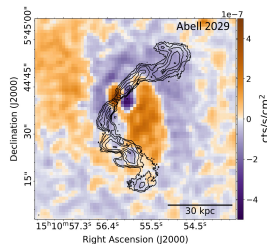
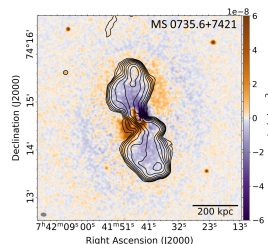
Low-frequency radio lobes likely
to be **valuable proxy**

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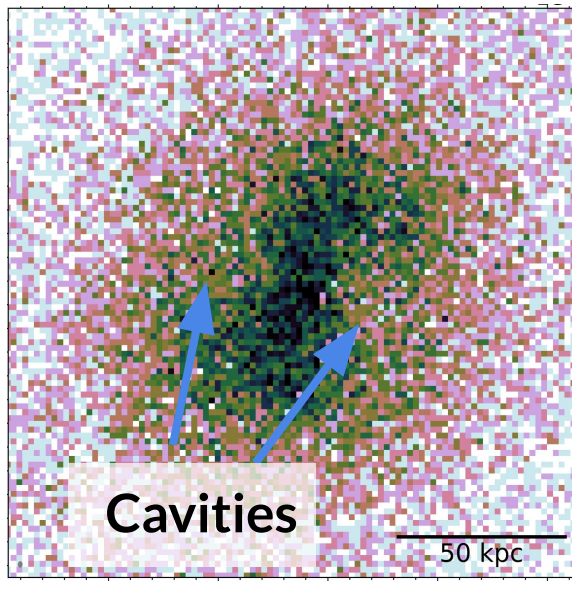


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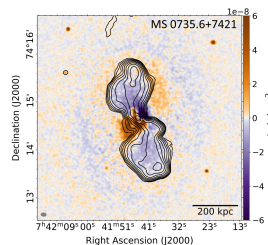
An example of an outlier:
ZwCl 2701



Further discussion points

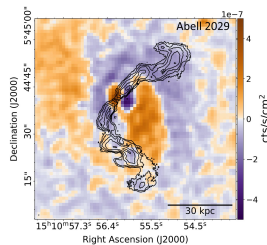
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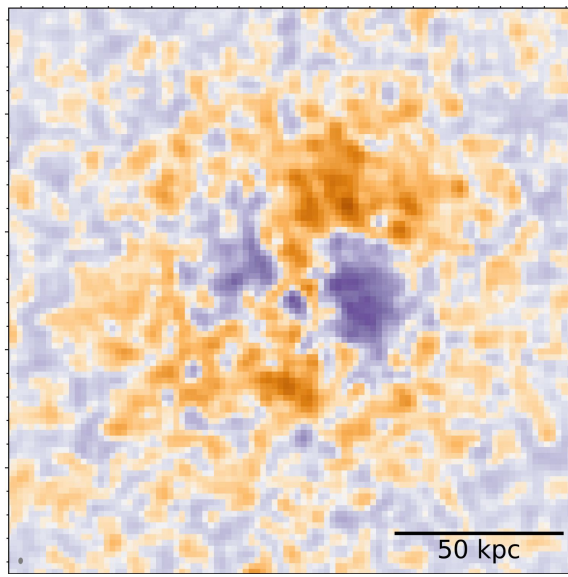
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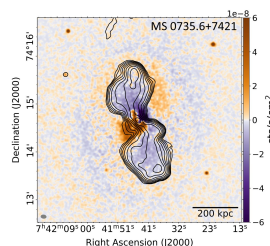
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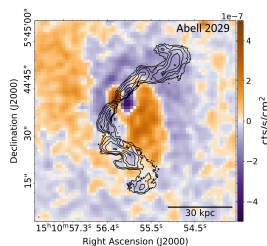
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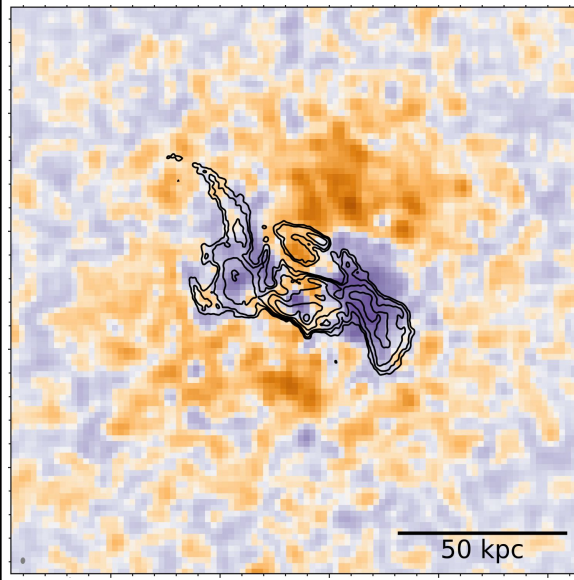
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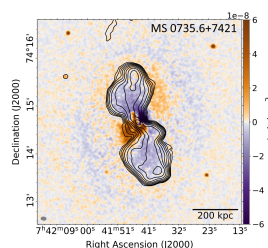
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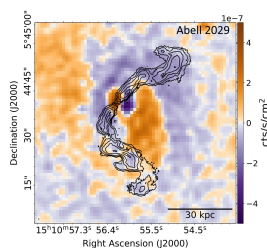
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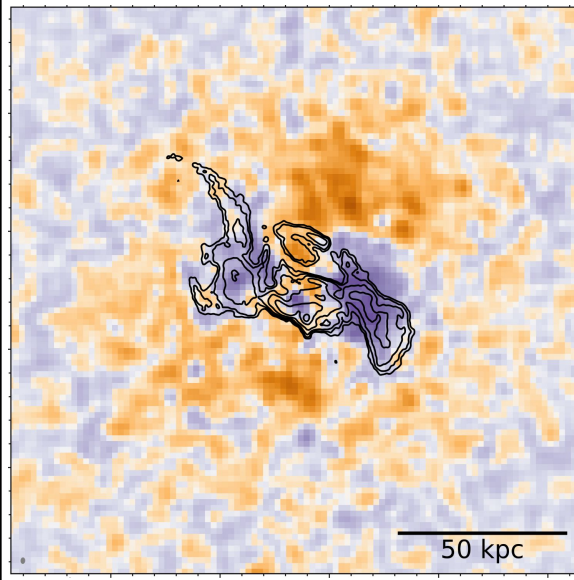
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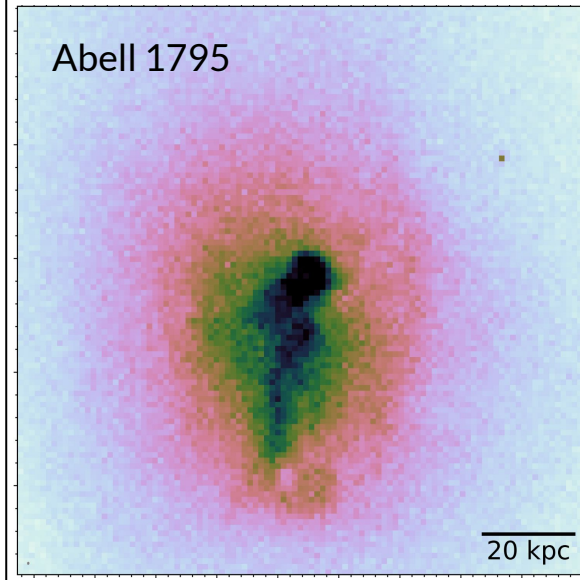
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Radio lobes especially valuable with complex ICM structures

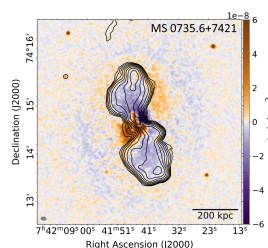
Abell 1795



Further discussion points

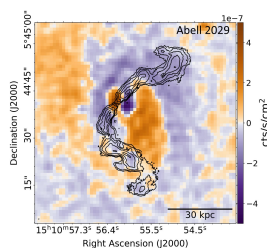
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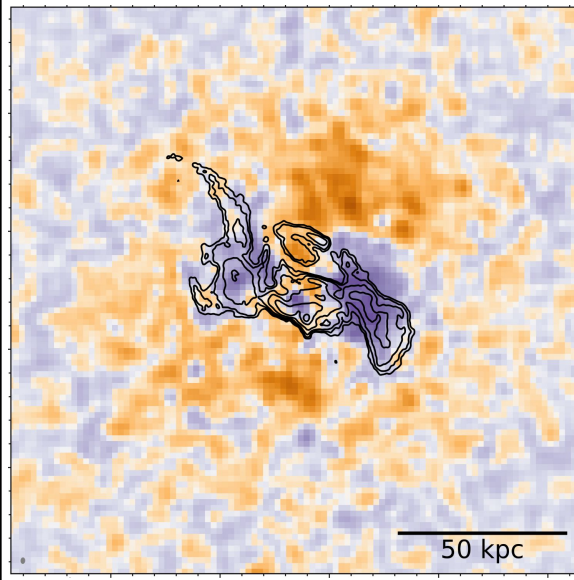
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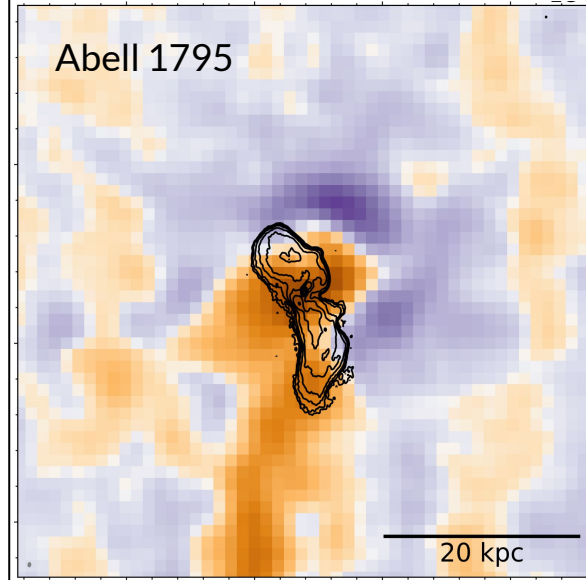
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High-redshift observations

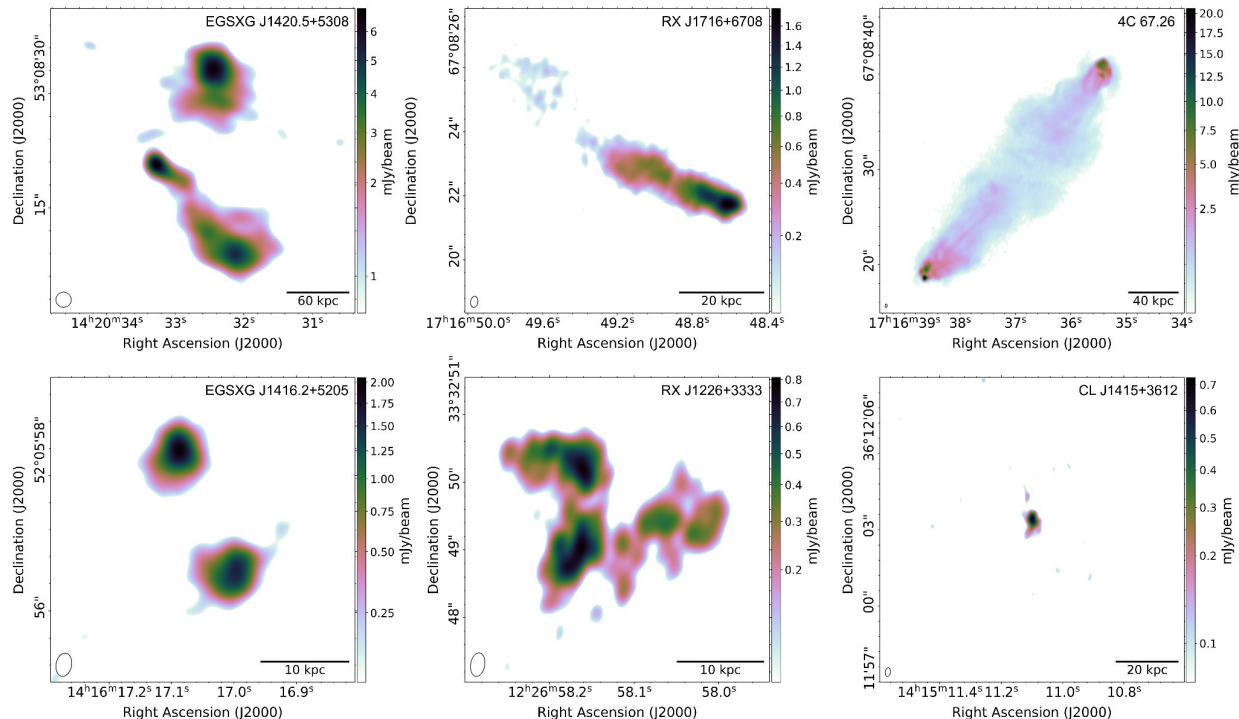
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- $z > 0.6$

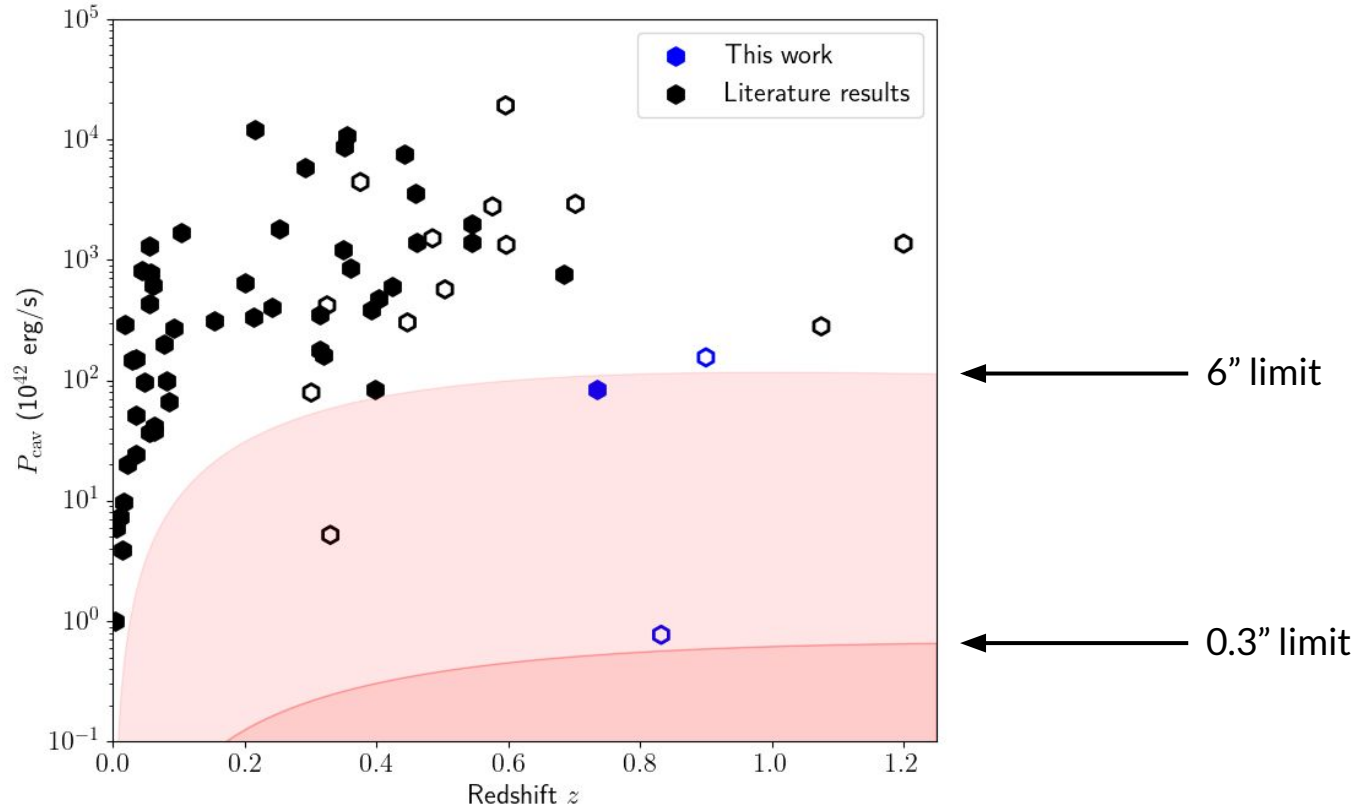
5 clusters with significant
radio emission →
(+1 bright member galaxy)

3 central galaxies with
clear radio lobes

→ 23% success rate
(compared to 4.7% based on
previous X-ray observations)



Progress towards the high-redshift regime

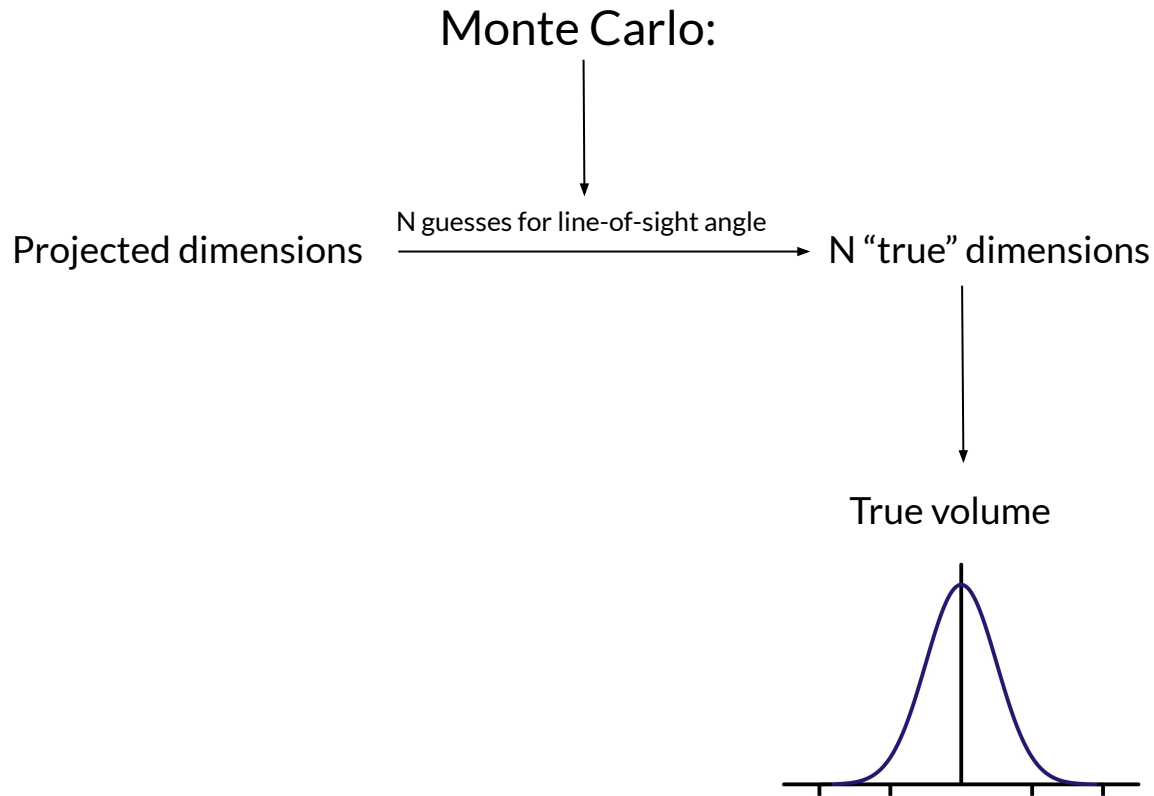
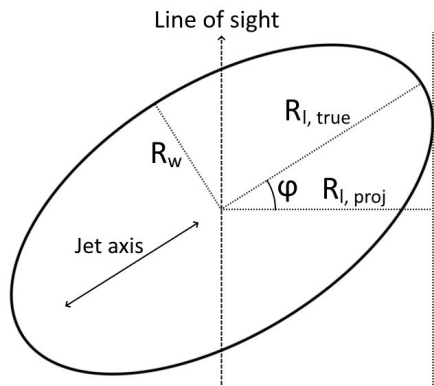


Take-home messages

- Observing a large sample of targets with the ILT has been proven to be a realistic endeavour
- The radio lobe 'cavity' volume measurements at 144 MHz are in good agreement with X-ray estimates, offering an additional opportunity to measure this quantity.
- **Recommendation** is to decide on a per-case basis whether to use X-ray cavities or radio lobes. **Follow the S/N!**
- We've started exploring the epoch of early formation and evolution of galaxy clusters!

Backup slides

Projection-based uncertainties



Observations

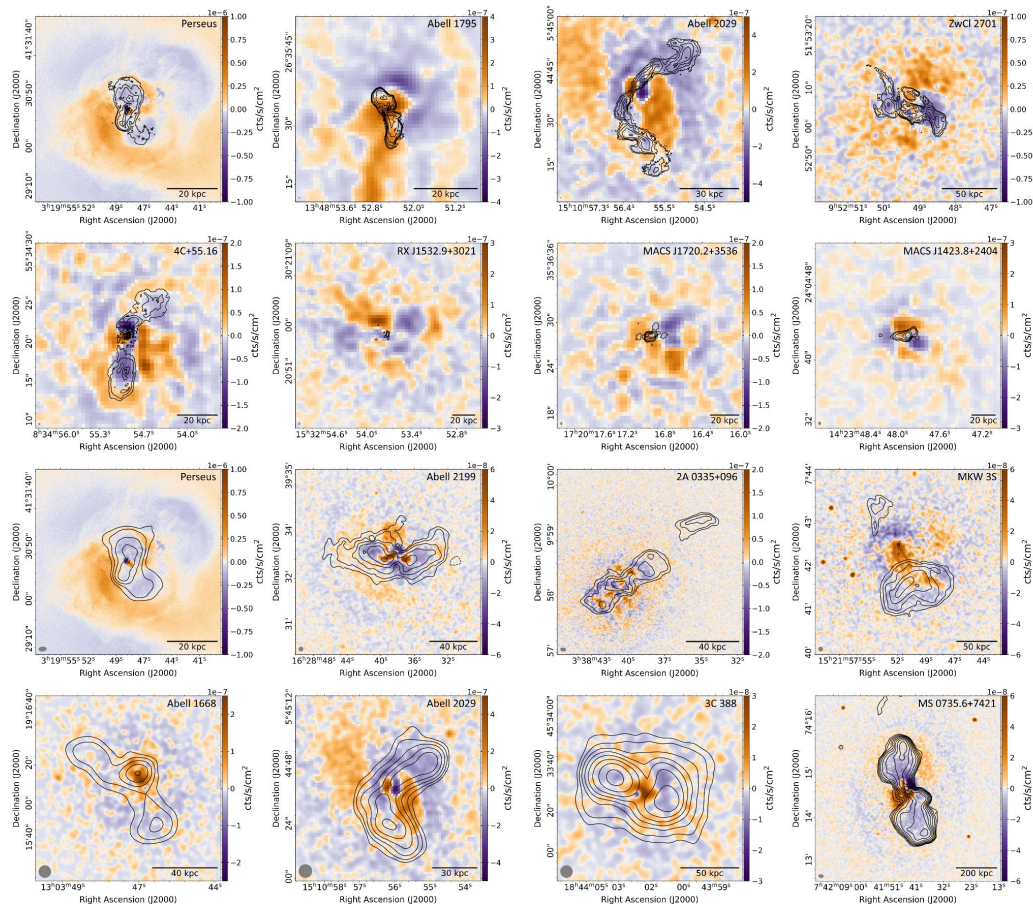
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